



Progress on the SOXS transients chaser for the ESO-NTT

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SPIE. ASTRONOMICAL
TELESCOPES +
INSTRUMENTATION

Montreal

17-22 July 2022



SOXS

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Son Of X-Shooter

- Single-object wide band spectrograph from U to H band @ESO-NTT 350-2000 nm
- 'Similar' to X-Shooter @VLT
- Two arms (VIS + NIR) with partial overlap around 800 nm to cross-calibrate spectra
- R~4,500 (3,500-6,000)
- Acquisition camera to perform photometry ugrizY-V (3.5'x3.5')



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Consortium

Institutes from 6 Countries

- Istituto Nazionale di AstroFisica (INAF), Italy
- Department of Particle Physics and Astrophysics, Weizmann Institute of Science, Rehovot, Israel
- Universidad Andres Bello & Instituto Milenio de Astrofisica (MAS), Santiago, Chile
- FINCA - Finnish Centre for Astronomy with ESO & Turku University, Turku, Finland
- Queen's University Belfast, UK
- Tel Aviv University, Israel
- Niels Bohr and Aarhus University, Copenhagen, Denmark



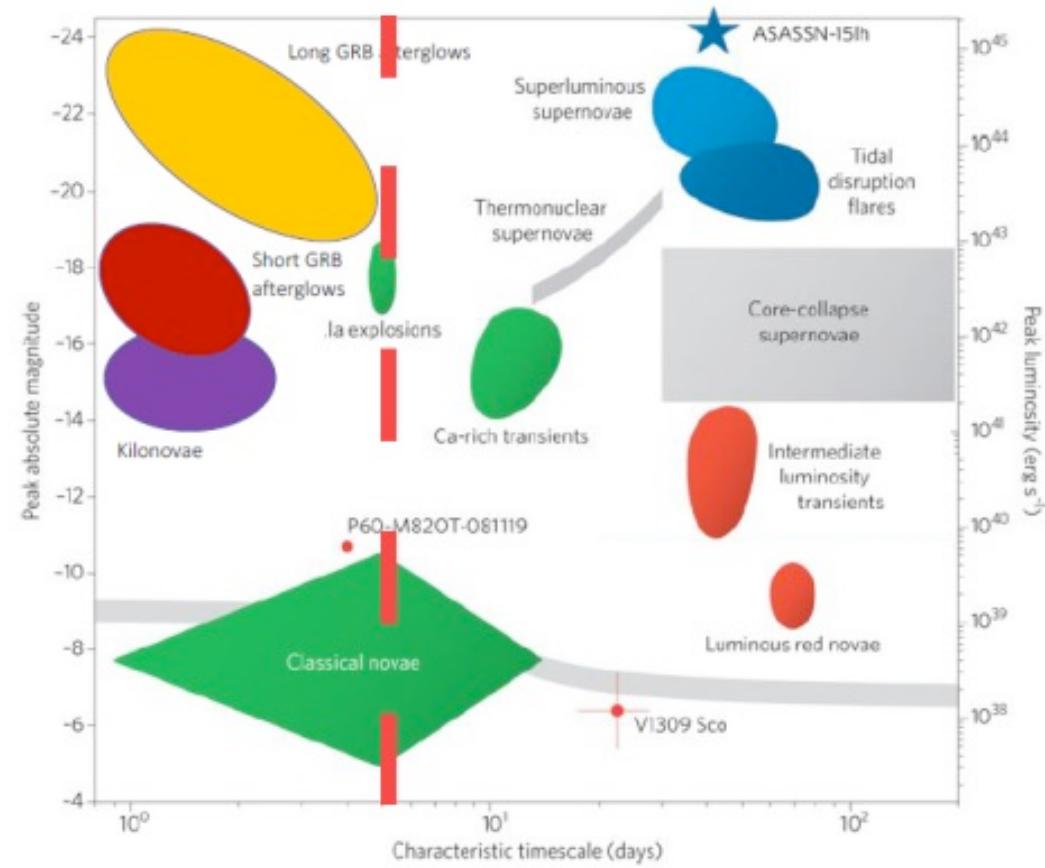
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Spectroscopic follow up of transients

- ❑ Classification of transients
- ❑ Supernovae (all flavours)
- ❑ Gravitational Wave events
- ❑ Neutrino events
- ❑ Blazars and AGN
- ❑ Nuclear transients and Tidal Disruption Events
- ❑ GRB and FRB
- ❑ Transient X-ray binaries, magnetars, ultra-luminous X-ray sources (NS & BH)
- ❑ Asteroids and Comets
- ❑ Young Stellar Objects, stellar variability, exoplanets
- ❑ The Unknown



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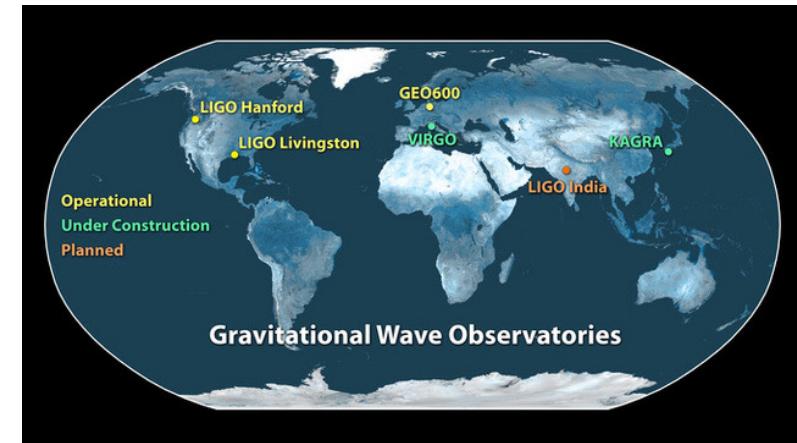


Sinergies

SOXS will have 180 n/yr (for ≥ 5 yr)
 $\sim 3,000\text{-}4,000$ spectra/yr

A spectroscopic machine for the transient sky

- New deeper survey: Vera Rubin, PanSTARSS, DES, ZTF, ...
- Space optical missions: Gaia, EUCLID, ...
- Space high-energy missions: Swift, Fermi, SVOM, ...
- Radio new facilities: MeerKAT, SKA, ...
- VHE: CTA
- Messengers: aLIGO-Virgo, KM3Net, ANTARES, ...

**SKAO****cta**
cherenkov telescope array**SVOM**

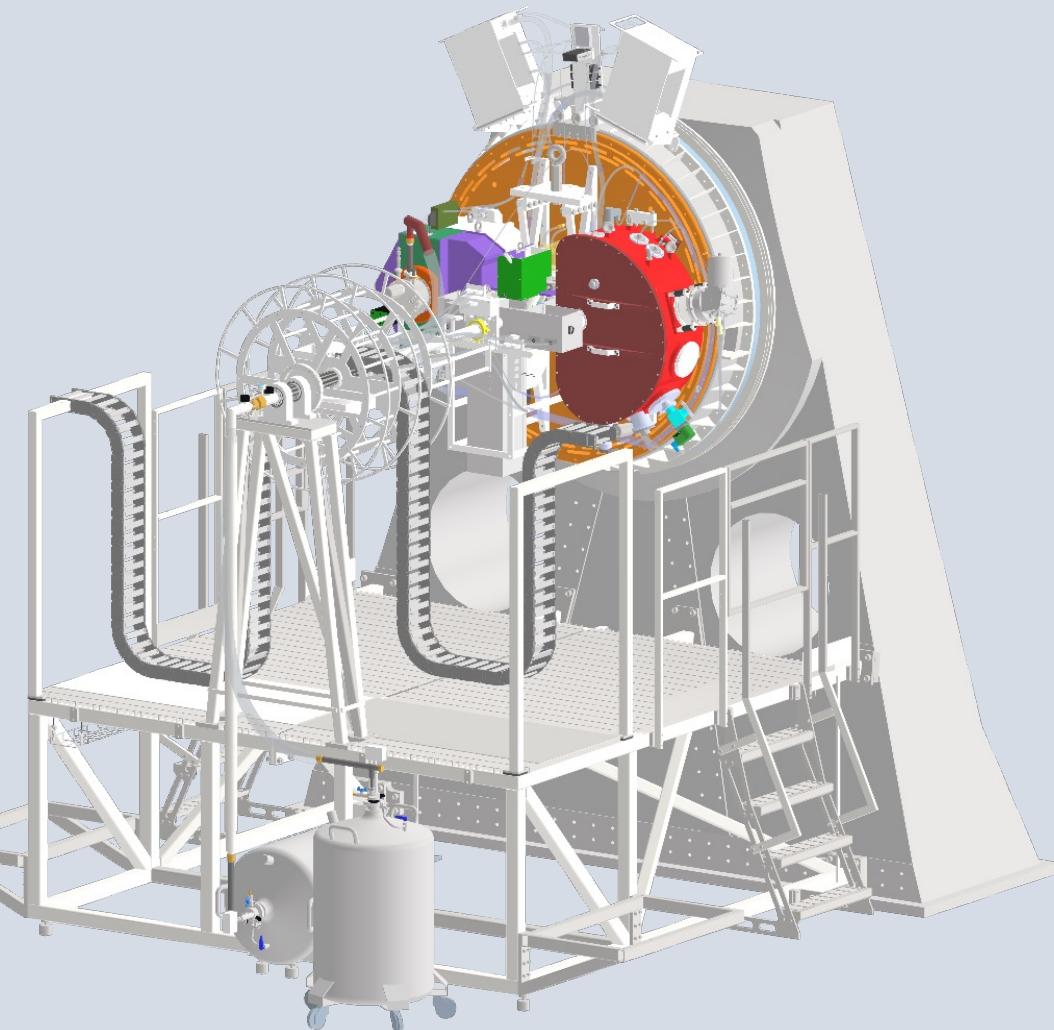
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	UV-VIS	NIR
Spectral range	350-850 nm	800-2000 nm
Resolution (1" slit)	>3600 (\approx 4500 avg)	5000
Slit widths	0.5 - 1 - 1.5 - 5 arcsec	0.5 - 1 - 1.5 - 5 arcsec
Slit height	12 arcsec	12 arcsec
Detector	e2V CCD44-82 2Kx4K	Teledyne H2RG 2Kx2K
Pixel Size	15 μm	18 μm
Detector Scale	0.28"/pixel	0.25"/pixel

	Camera
Spectral range	360-970 nm
Detector	Andor iKon M-934 1Kx1K
Field of View	3.5'x3.5'
Pixel Size	13 μm
Detector Scale	0.205"/pixel





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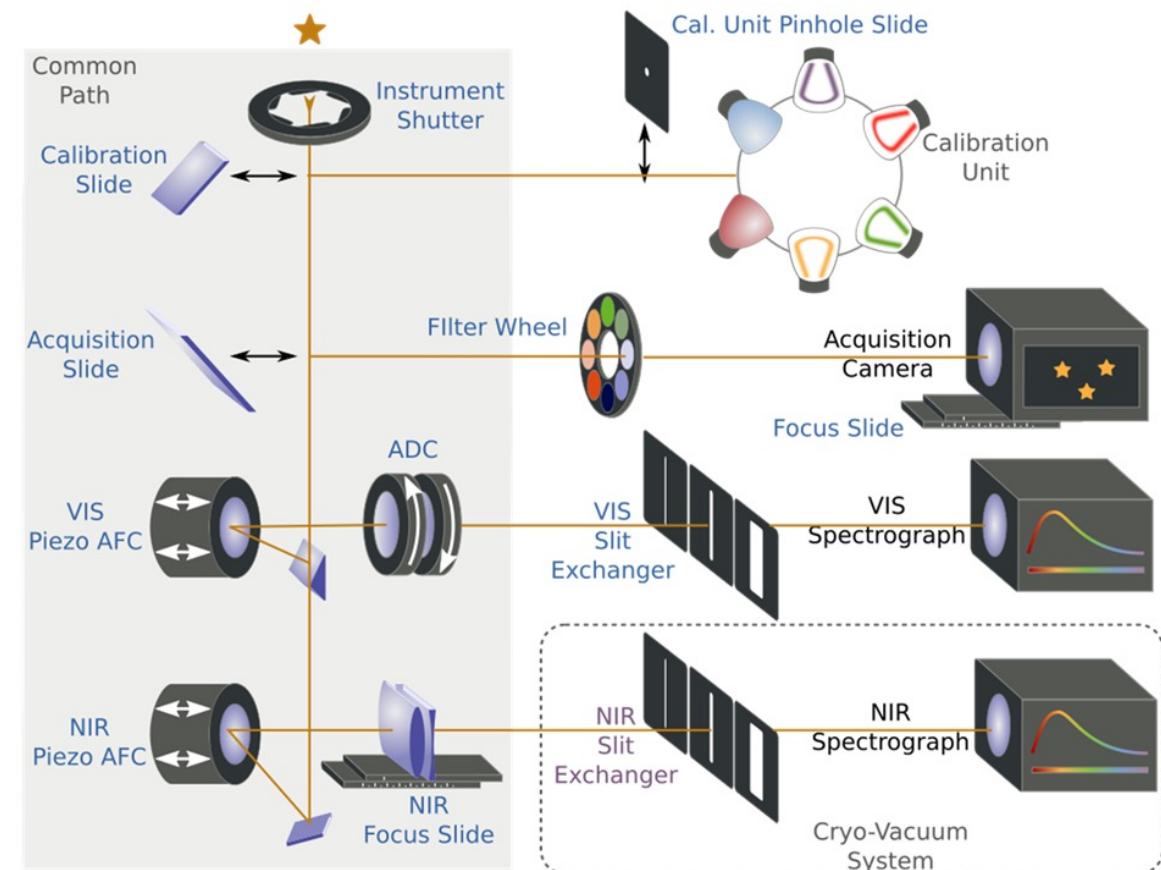
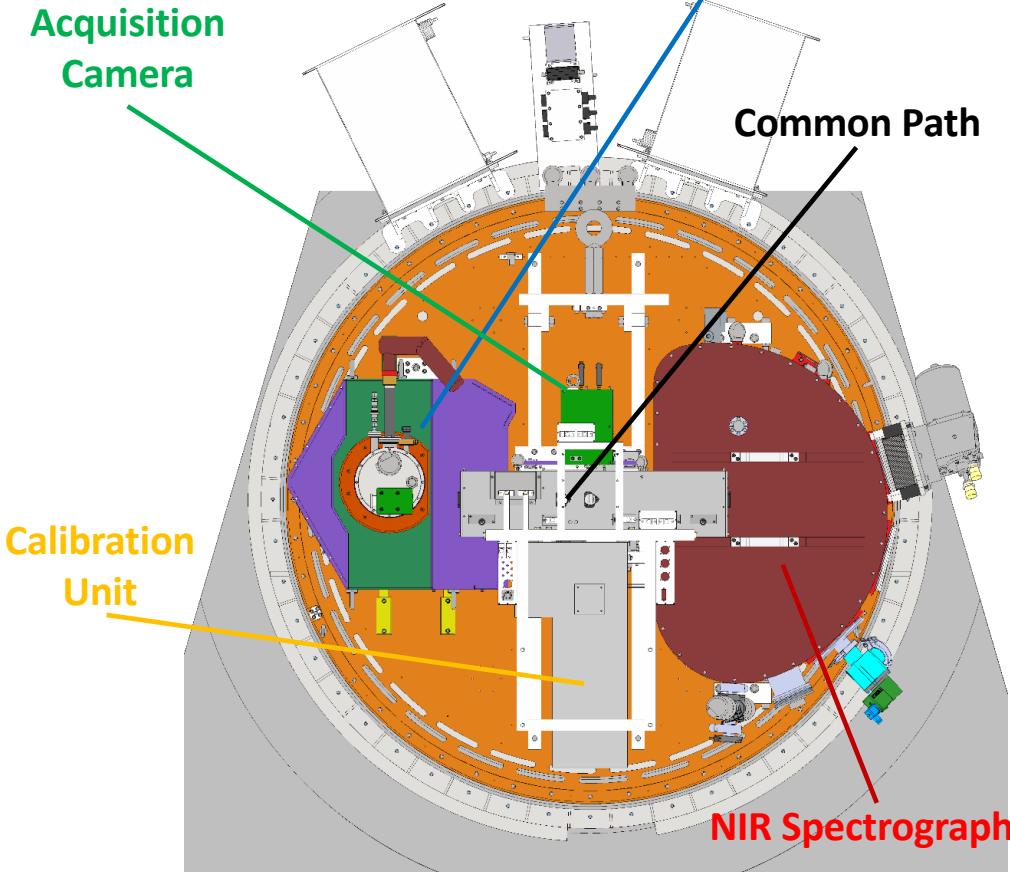
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UV-VIS Spectrograph





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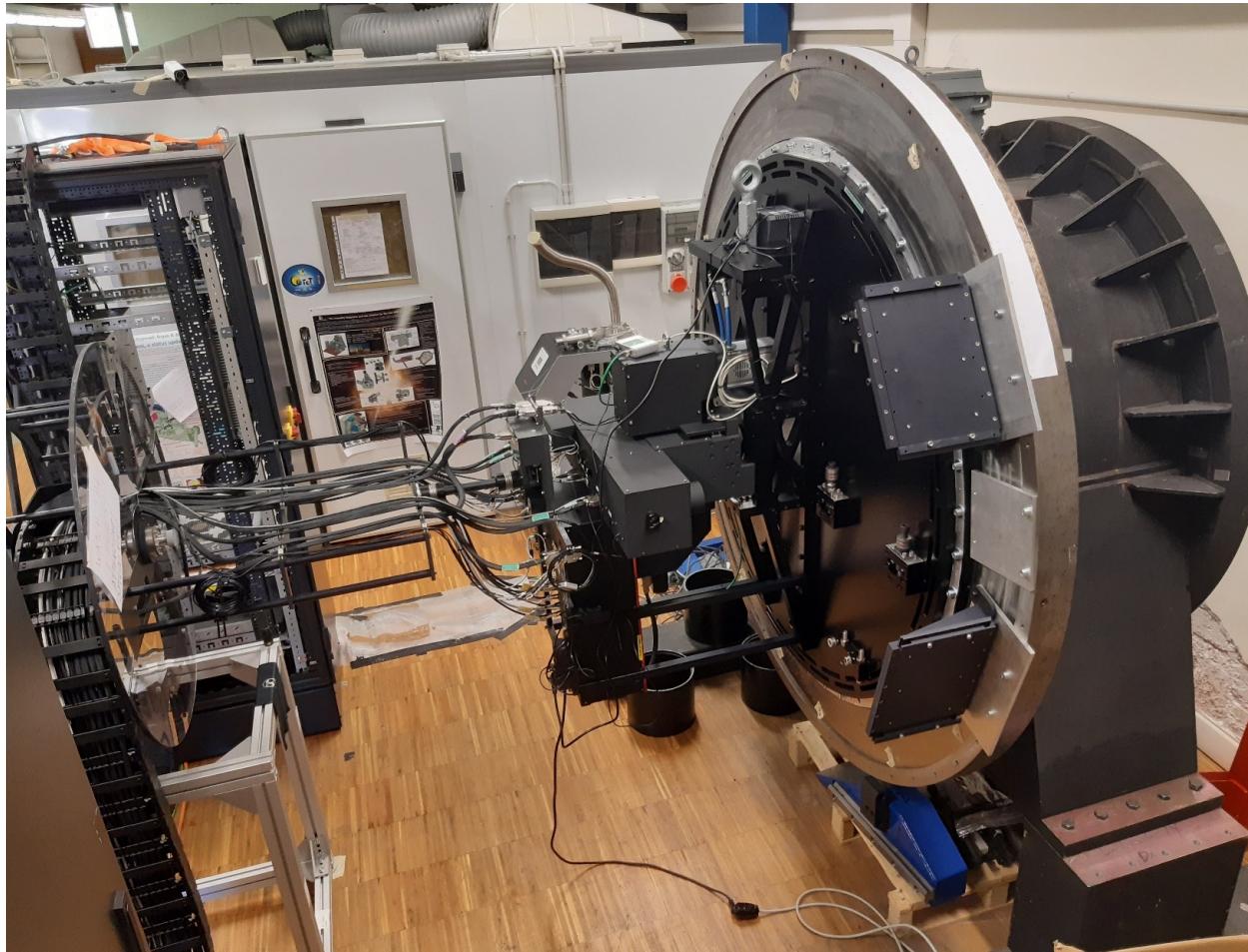
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AIT @INAF Padova





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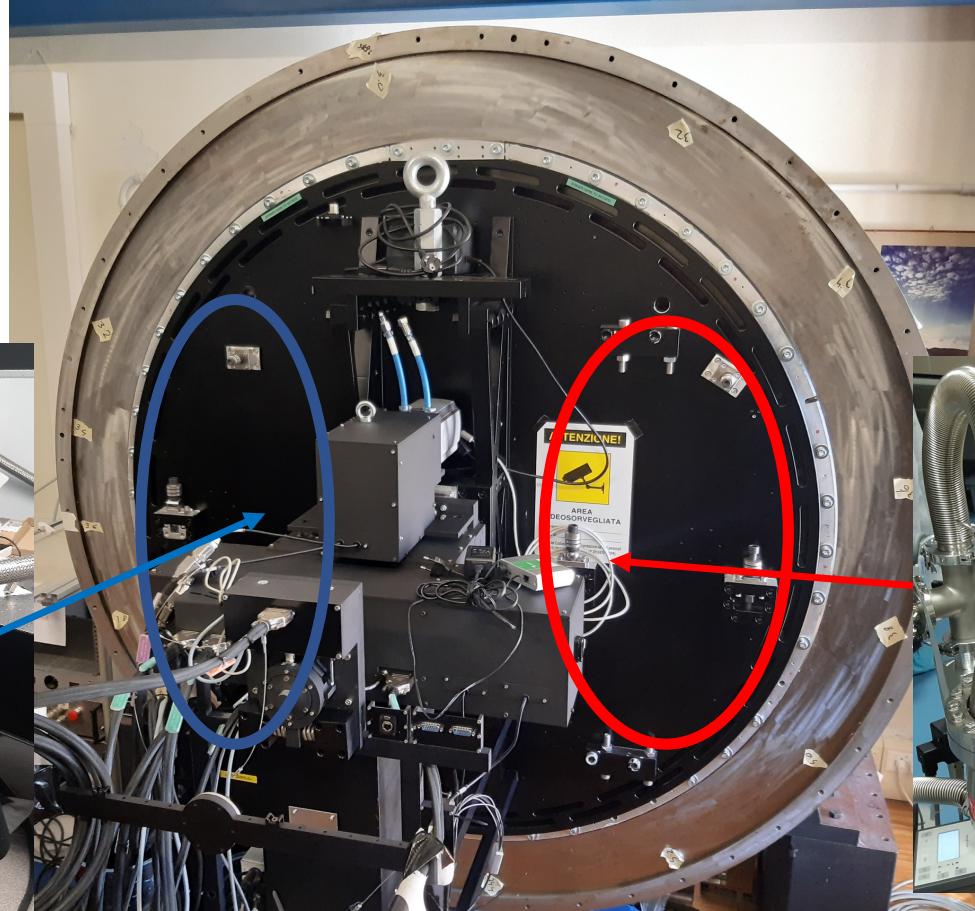
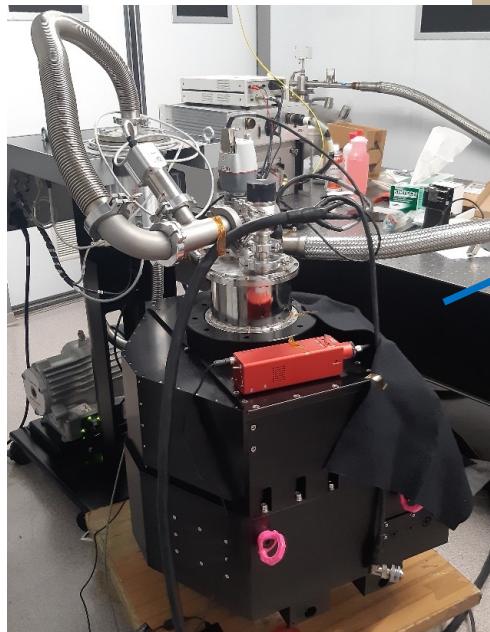
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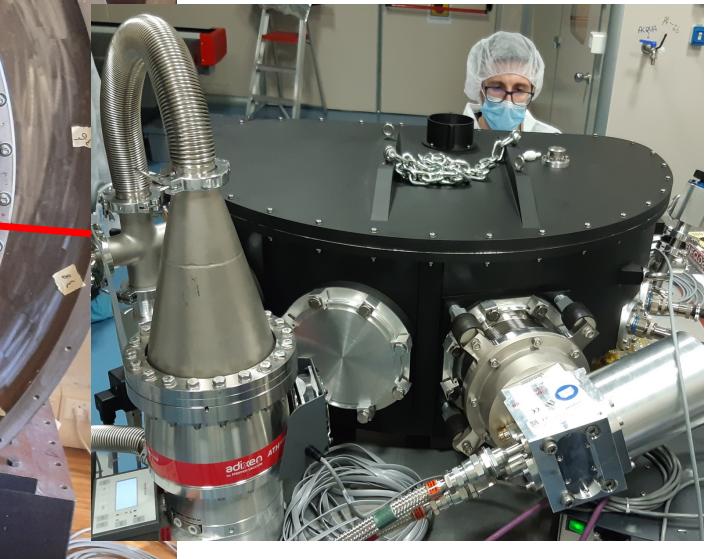
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UV-VIS Spectrograph @INAF Padova



NIR Spectrograph AIT Phase @INAF Merate

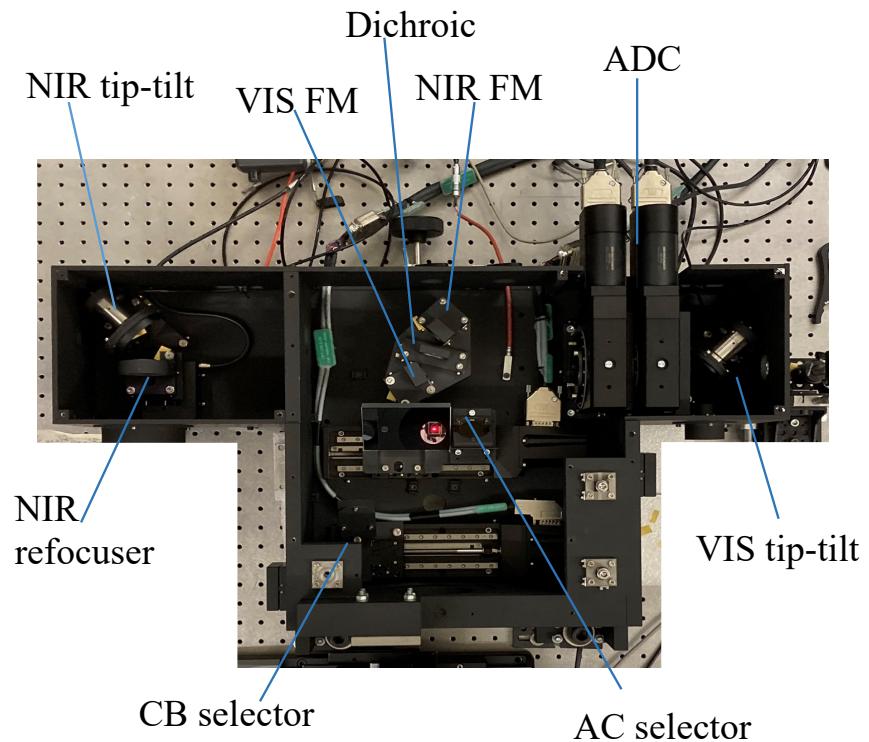
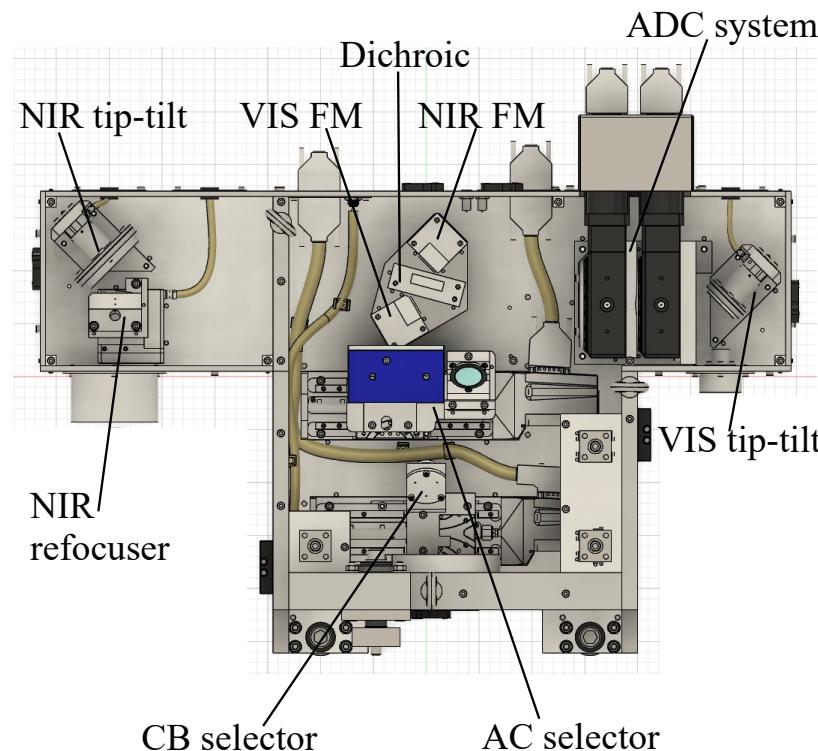


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Common Path

**Archer OptX**
Optical Precision. Optimal Outcome.**OPTIMAX****ASAHI SPECTRA** USA**PECCHIOLI RESEARCH**
OPTICS, IDEAS AND MORE



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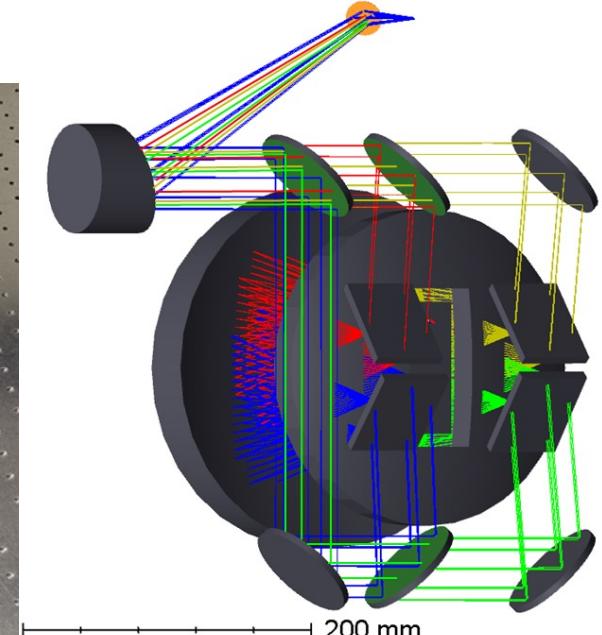
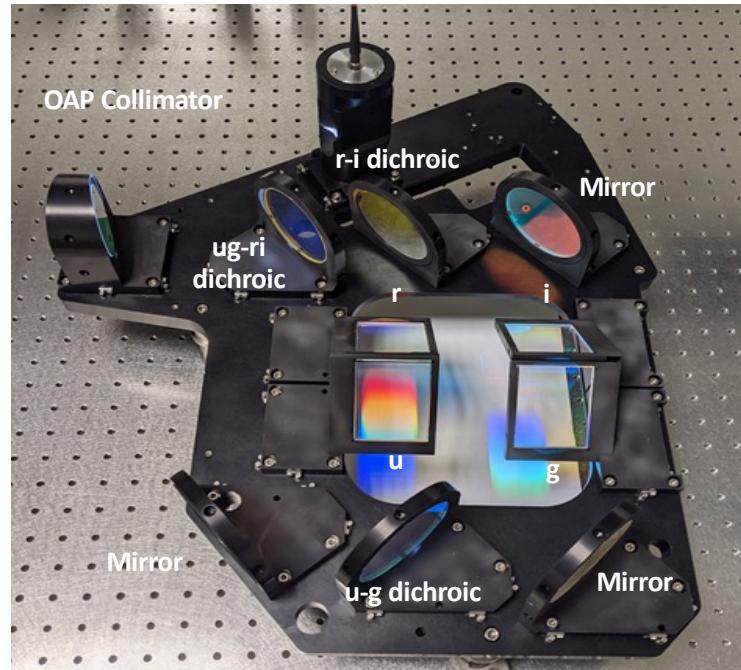
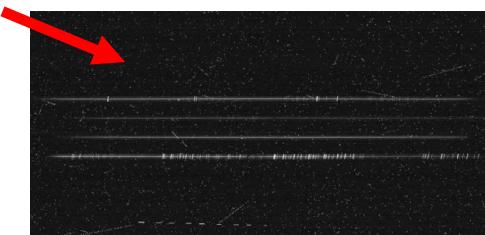
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UV-VIS: Multi-Imaging Transient Spectrograph

- ❑ Collimated beam is divided to 4 bands using 3 dichroics.
- ❑ Each band has its own optimized disperser
- ❑ Single camera
- ❑ 1st order dispersion, $\mathcal{R} \sim 4500$ at $\alpha_{Lit.}$
- ❑ 4 bands quasi-orders are imaged onto a single 4k×2k CCD.



Winlight System ASAHI SPECTRA USA

Fraunhofer
IOF

Quasi-Order	Wavelength Range [nm]
u	350 – 439.5
g	427 - 547
r	527 - 680
i	664 – 850



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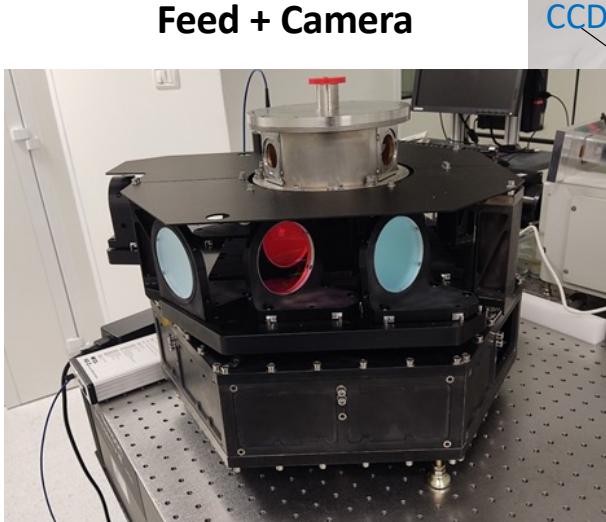
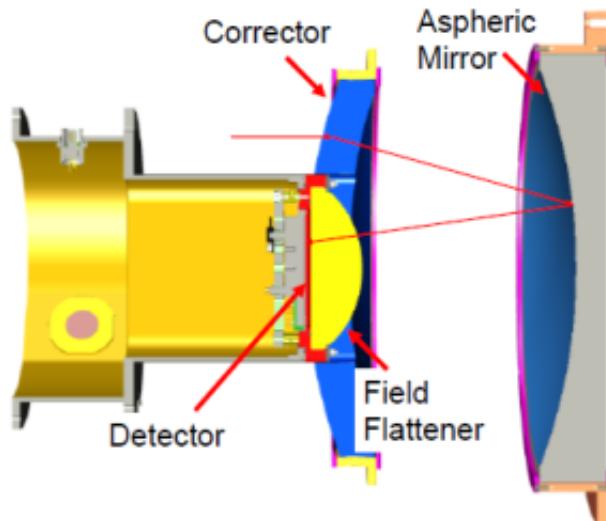
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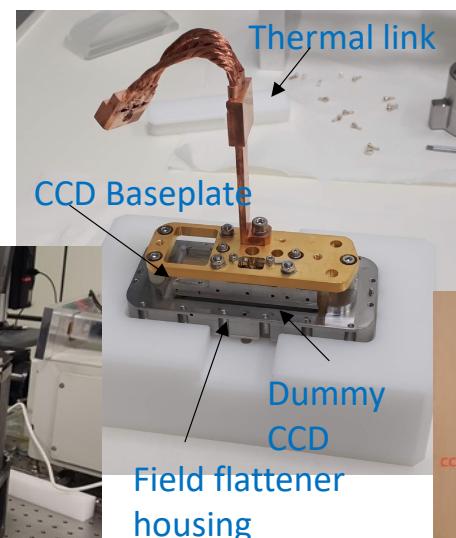
UV-VIS Camera

- Three element catadioptric camera: all aspheric
- Used as 4 off axis F/3.1 cameras.
- CaF₂ corrector + Fused Silica Field Flattener
- Low CTE=>Athermal camera

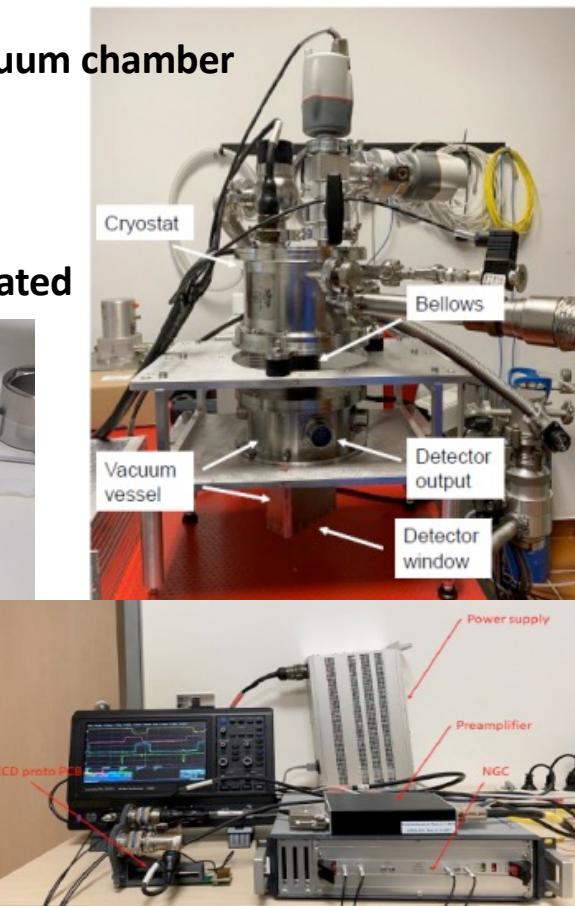


Feed + Camera

CCD Invar Baseplate gold plated



CCD vacuum chamber



ESO NGC Controller



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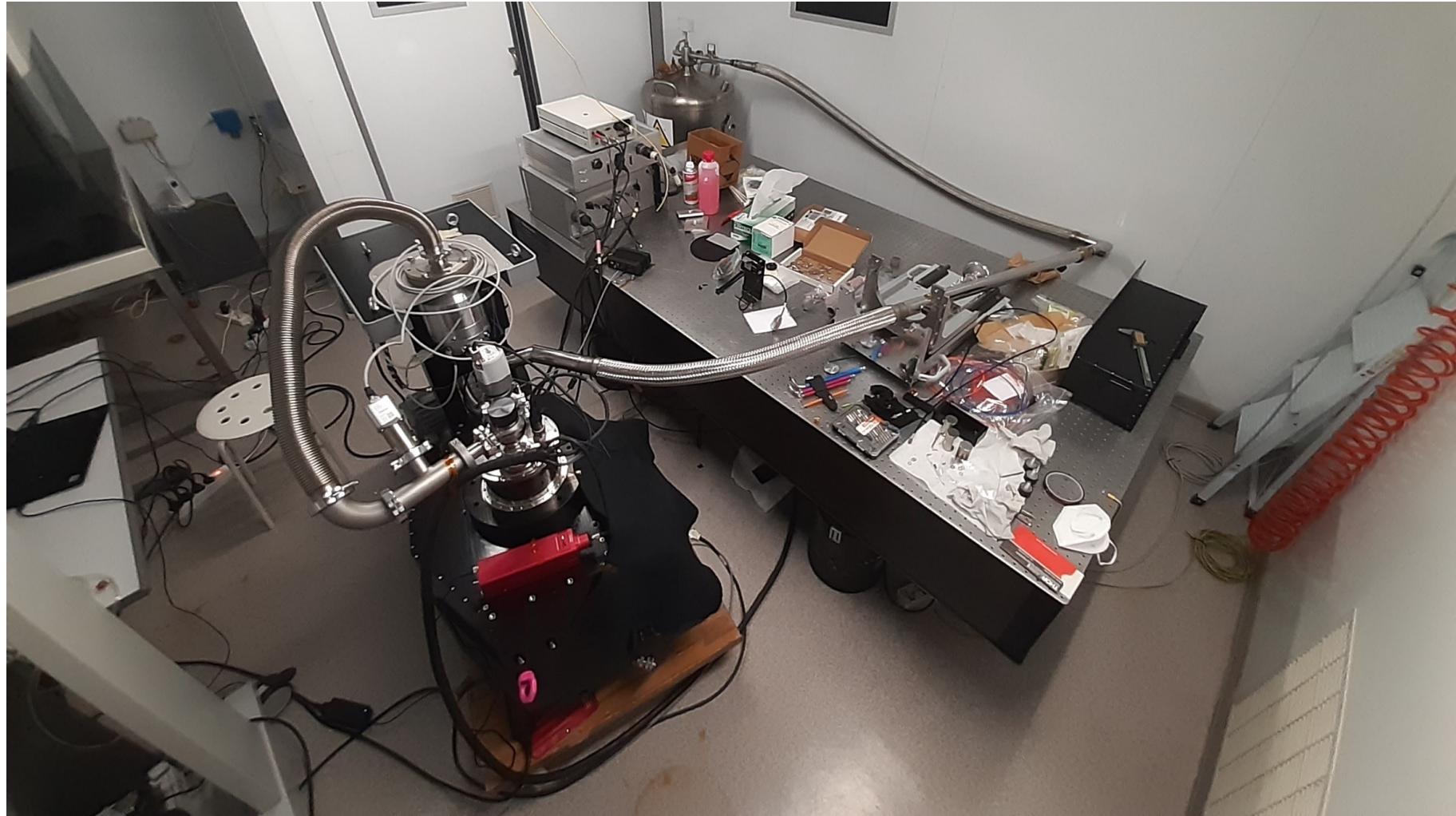
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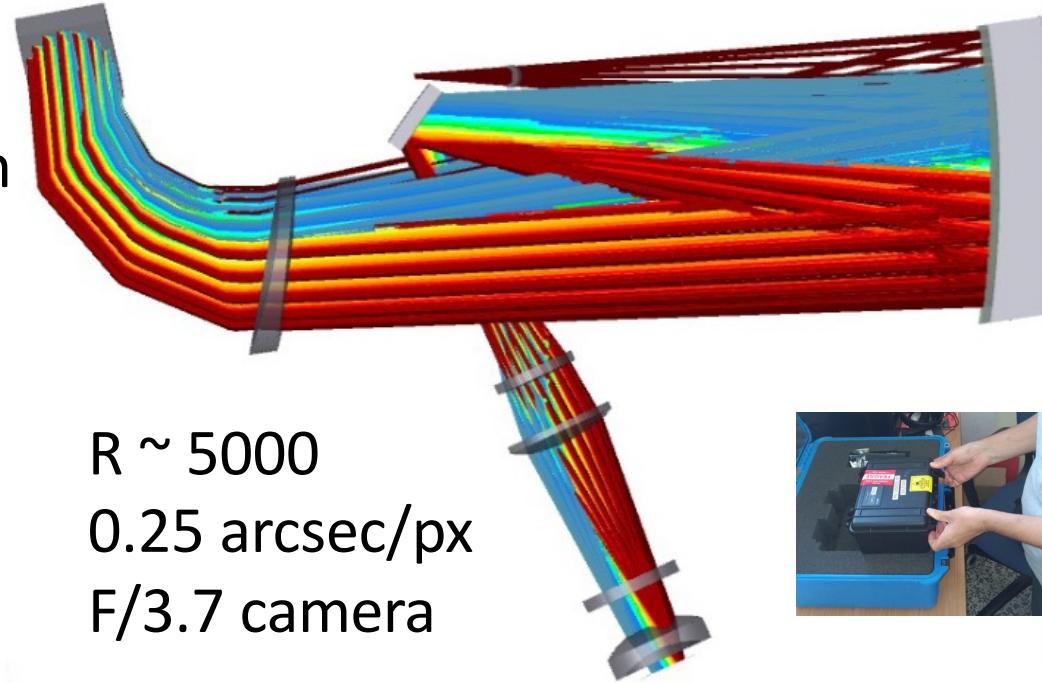


NIR Spectrograph

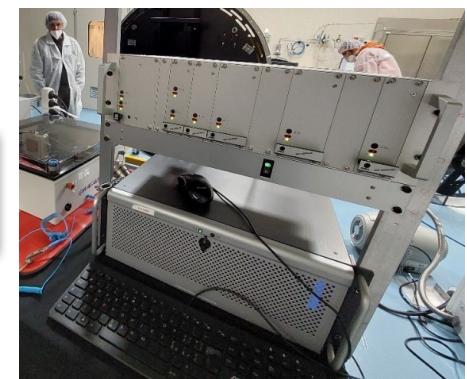
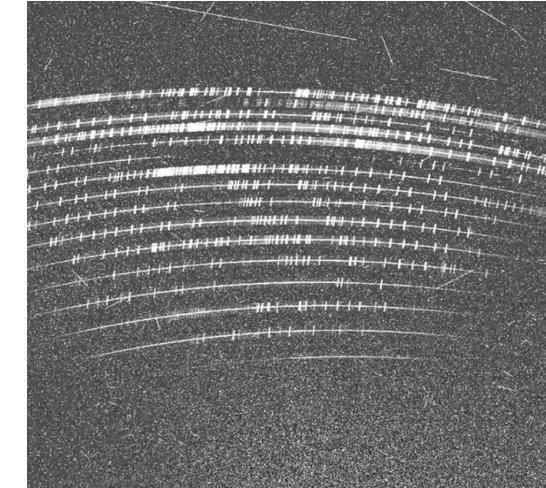
4C Design

Spectrograph with
Collimator
Compensation of
Camera
Chromatism

Echelle
Cross-Dispersed



$R \sim 5000$
0.25 arcsec/px
F/3.7 camera



ESO NGC Controller





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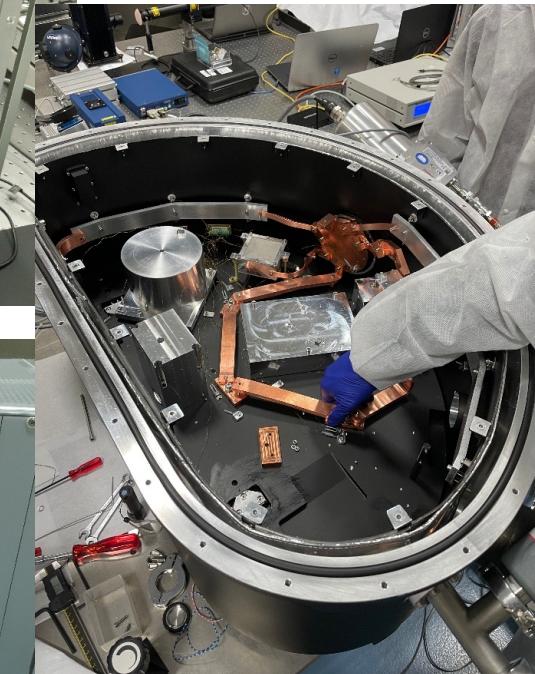
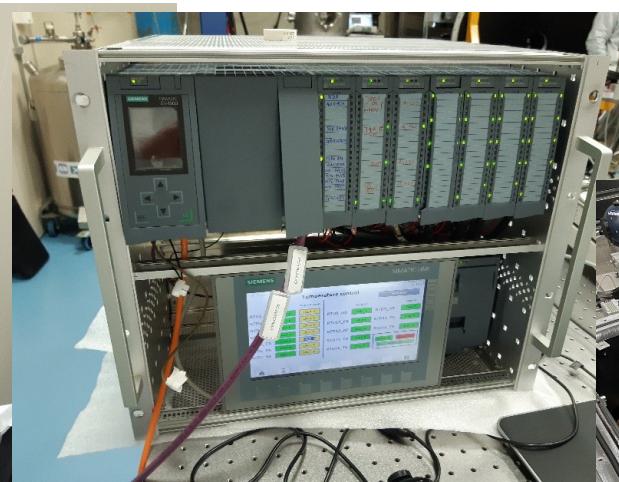
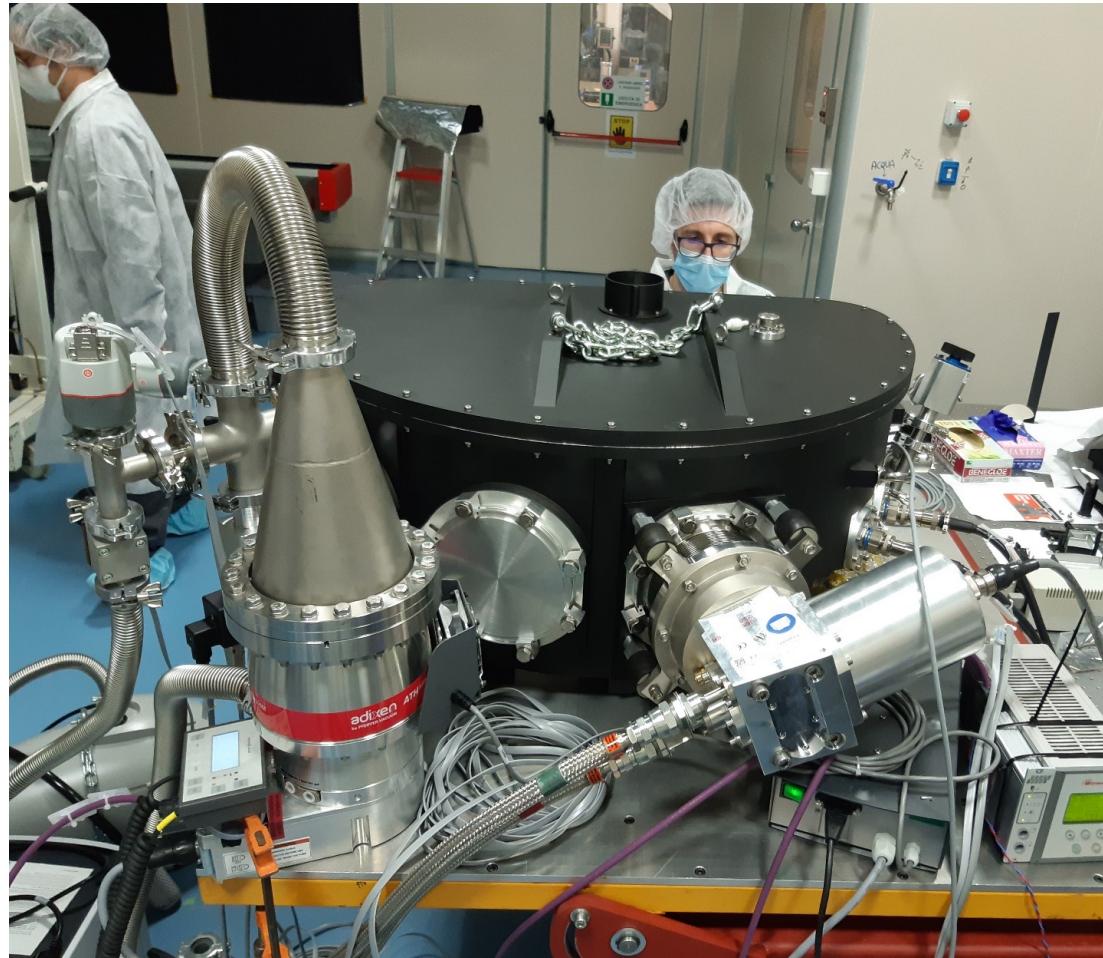
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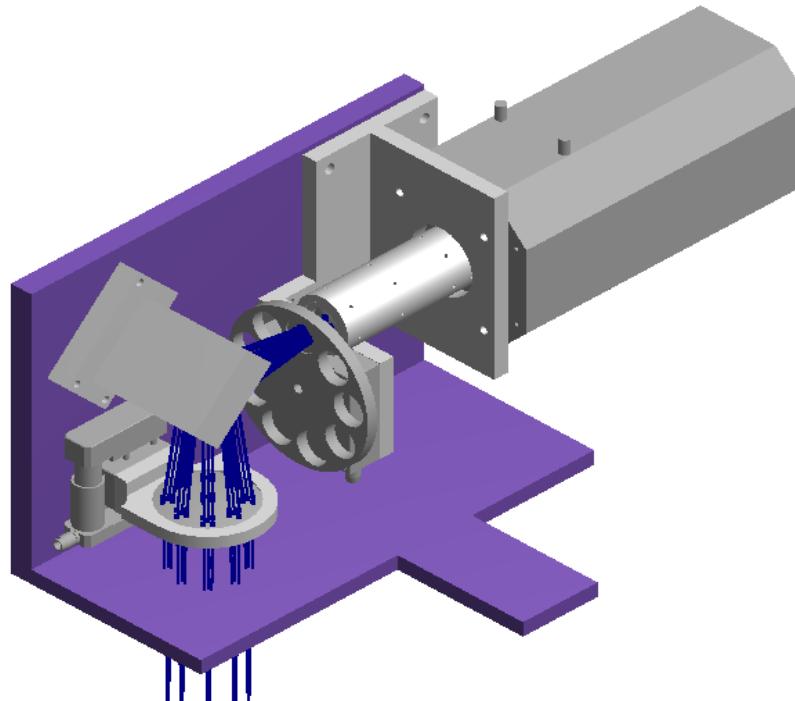
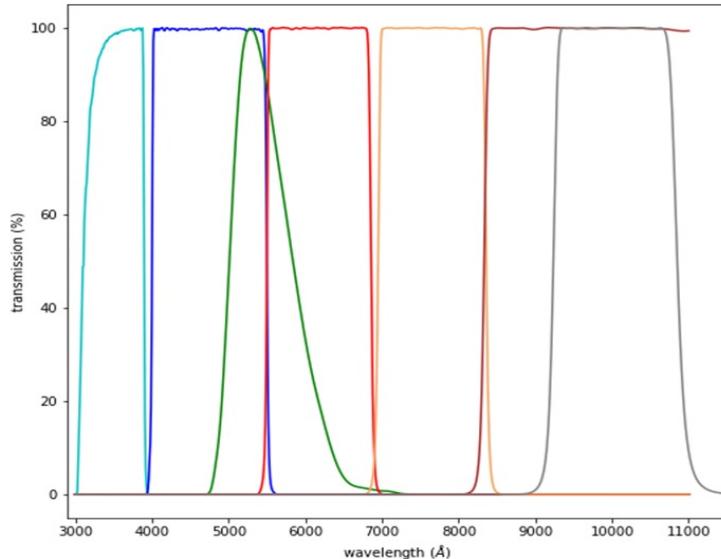
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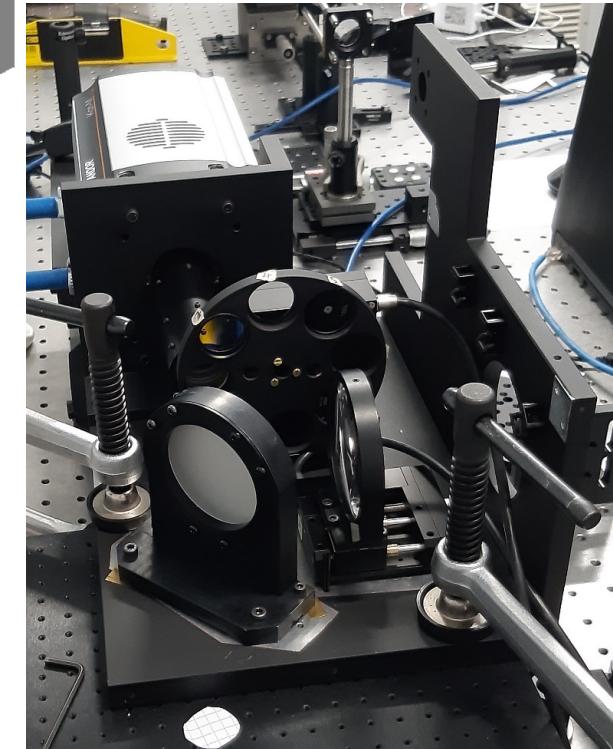


A&G Camera

- Target Acquisition
- Secondary guiding
- Photometry



FoV: 3.5'x3.5'
Filters: ugrizY + V



ASAHI SPECTRA^{USA}

PECCHIOLI RESEARCH
OPTICS, IDEAS AND MORE



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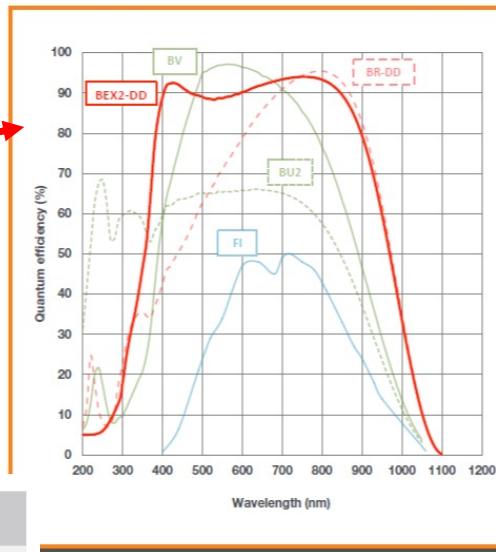
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A&G Camera

BEX2-DD => High
QE in a broad
wavelength range



Active pixels	1024 x 1024
Sensor size	13.3 x 13.3 mm
Pixel size (W x H)	13 μm x 13 μm
Active area pixel well depth	100,000 e ⁻ (130,000 e ⁻ for BR-DD and BEX2-DD models)
Pixel readout rates (MHz)	5, 3, 1, 0.05
Read noise	2.9 e ⁻
Maximum cooling	-100°C
Frame rate	4.4 fps (full frame)

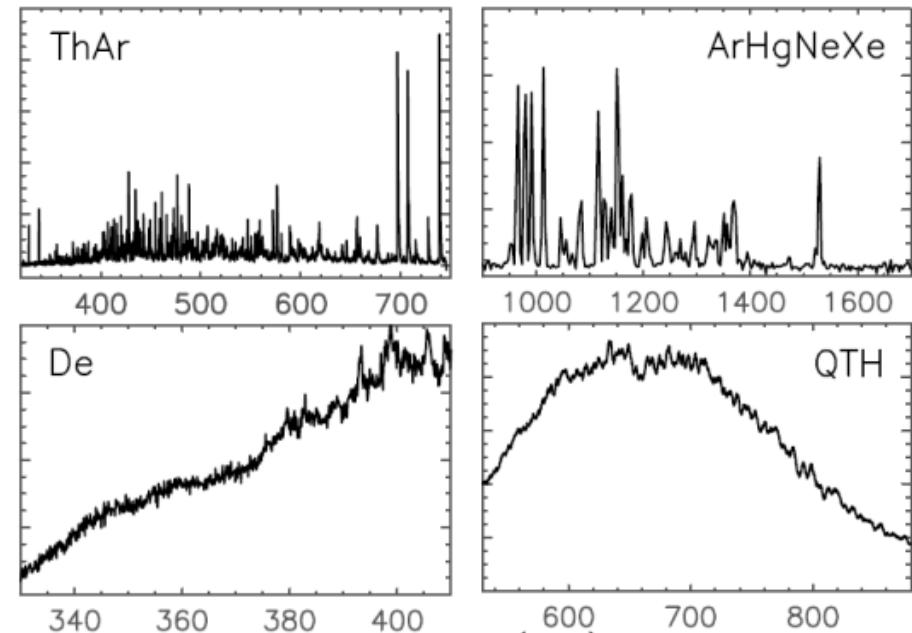
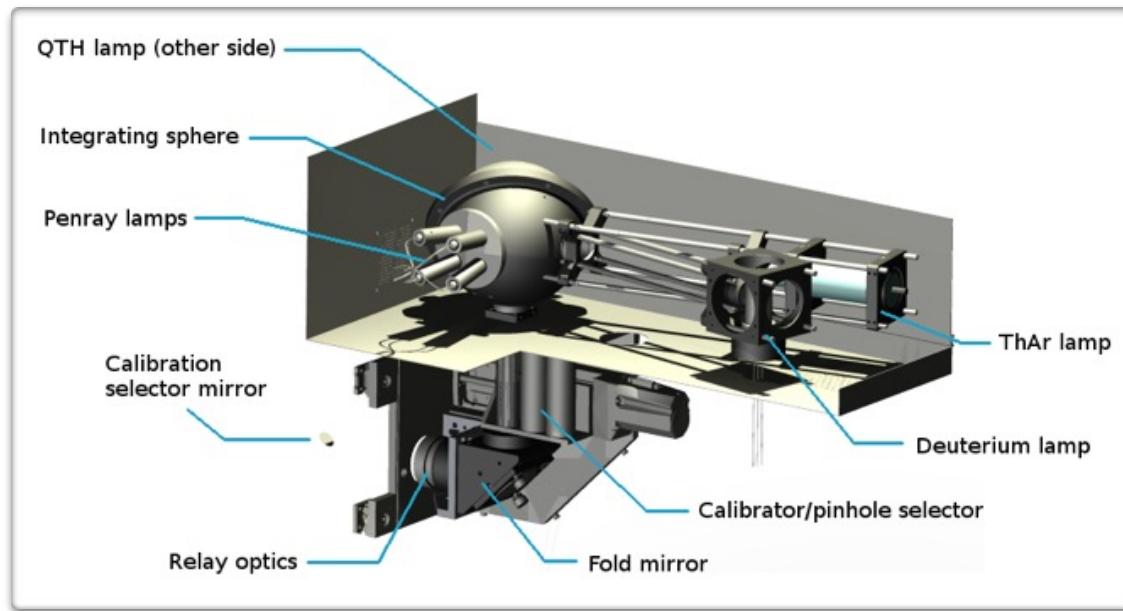
Andor iKon M934
1024x1024
13μm/px 0.205 “/px

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Calibration Unit

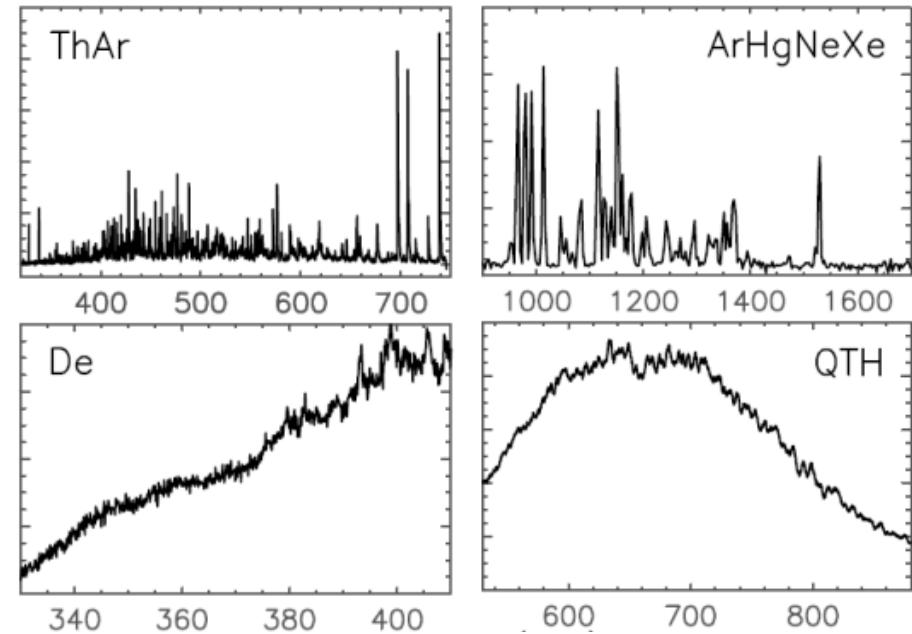


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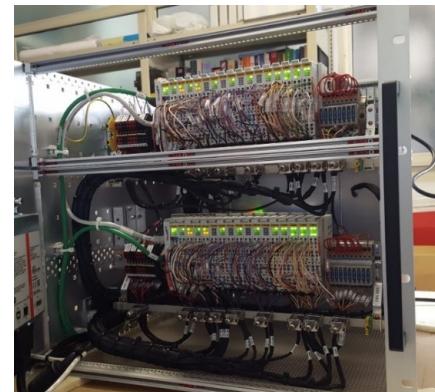
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Electronics





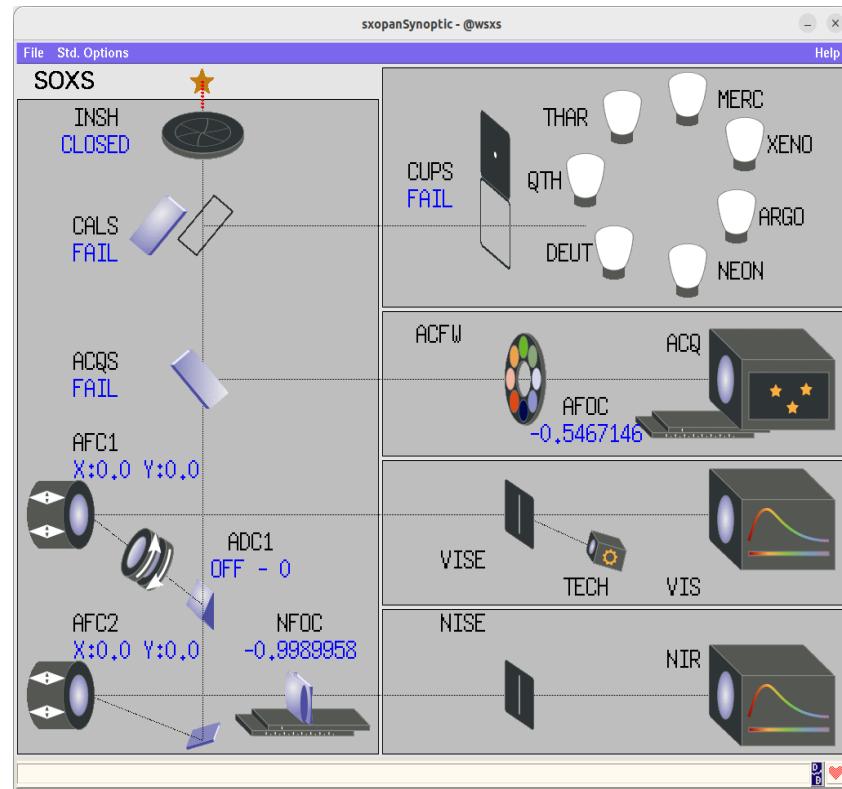
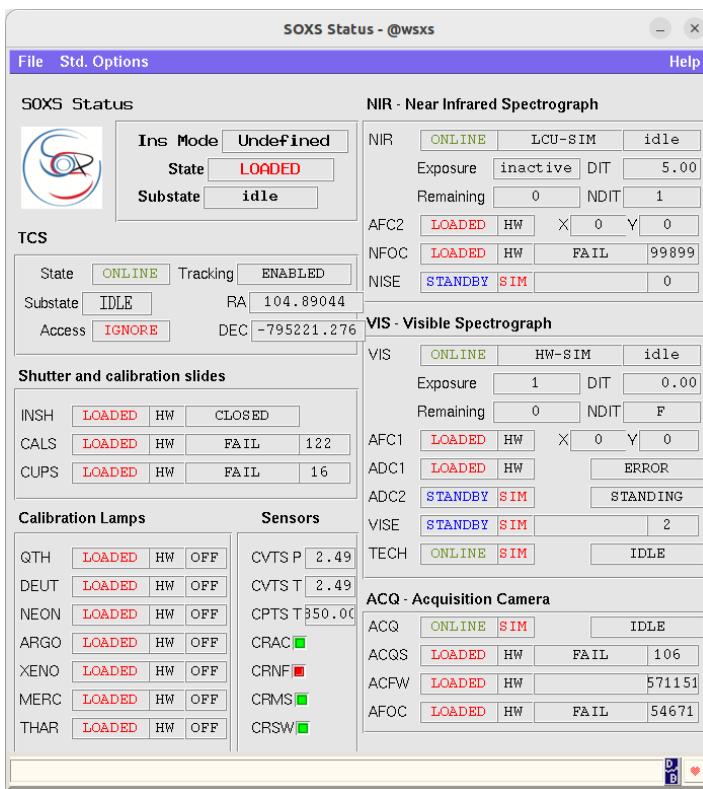
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Instrument Software





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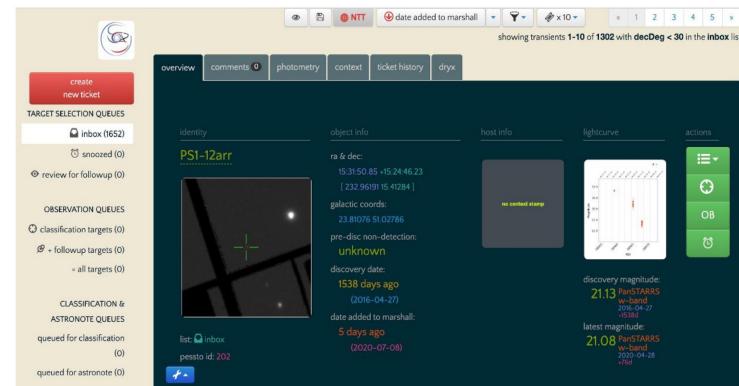
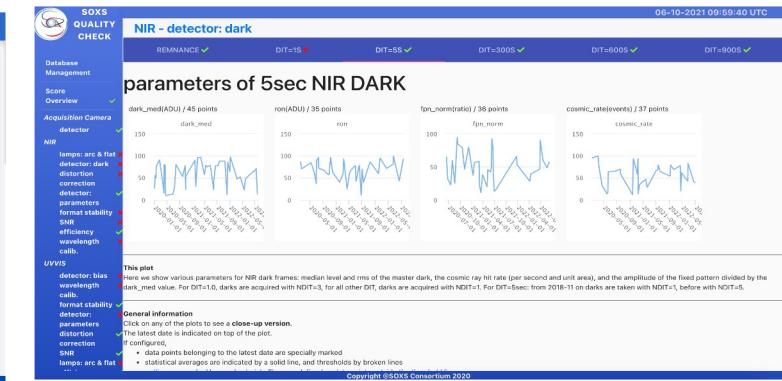
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Operations

- ETC
- Pipeline
- Scheduler
- Quality Control
- Remote Monitoring
- Help desk

The screenshot shows the SOXS Scheduler interface with a table titled "List of OBs". The columns include AID, OB Type, Target Name, Ra., Dec., Magnitude, Exp. Time, and Actions. There are 30 items per page. The table lists various observations such as PKS 1553-11Bris, AT2018bjq, and SN2018fuk.



5 years



SOXS

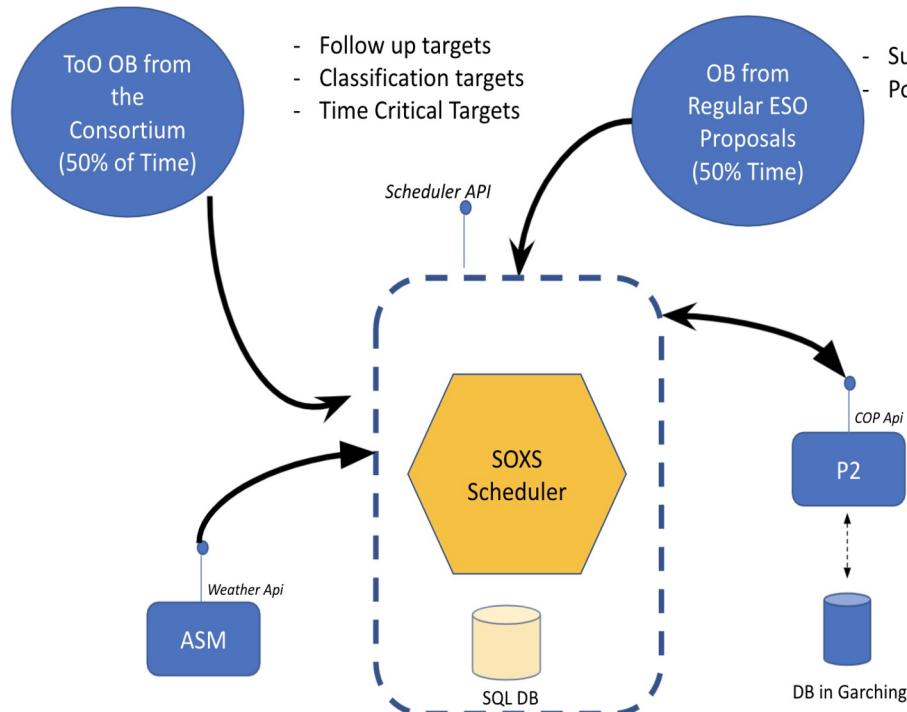
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Scheduler

- ❑ Schedule is updated daily
- ❑ Telescope operator on site



- ❑ P2 system, vOT interface with ESO
- ❑ Web based app

The screenshot shows a web-based application titled "SOXS SCHEDULER". On the left, a sidebar includes options like "Switch to VES", "New OB", "New Urgent OB", "Refresh", and "Show Logs". The main area is titled "List of OBs" and displays a table of 12 rows. Each row contains information about an observing block, including its ID, OB Type, Target Name, Right Ascension (Ra.), Declination (Dec.), Magnitude, Exposure Time (Exp. Time), and Actions (represented by icons). The table has columns for ID, OB Type, Target Name, Ra., Dec., Magnitude, Exp. Time, and Actions. The first few rows show entries such as "Undefined PKS 1553+113trs", "Classification AT2018qsi", and "Proposal ESO AT2018tn".

- ## Marshall Feeders:
- ❑ ZTF, ATLAS, PanStarrs, LSST-Lasair, etc.
 - ❑ TNS, Atel, GCN, etc.

The screenshot shows a software interface titled "Marshall". It features a top navigation bar with tabs like "create new ticket", "inbox (1652)", "comments", "photometry", "context", "ticket history", and "dryx". Below this is a search bar with the query "showing transients 1-10 of 1302 with decDeg < 30 in the inbox list". The main area is divided into several sections: "TARGET SELECTION QUEUES" (with "inbox (1652)" and "review for followup (0)" listed), "OBSERVATION QUEUES" (with "classification targets (0)" and "+ followup targets (0) -> all targets (0)" listed), "CLASSIFICATION & ASTRONOTE QUEUES" (with "queued for classification (0)" and "queued for astronote (0)" listed), and "identity" (showing an image of a star and coordinates "15:31:50.85 +15:24:46.23" and "galactic coords: 233.96919 15.41284"). Other sections include "object.info", "host.info", "lightcurve", and "actions".

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SPIE. ASTRONOMICAL TELESCOPES + INSTRUMENTATION

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- M. Landoni et al., "The SOXS Scheduling system", Paper No. AS110-29
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