# Spectrometer gratings based on direct-write e-beam lithography

U.D. Zeitner, T. Flügel-Paul, T. Harzendorf, M. Heusinger, E.-B. Kley

Fraunhofer Institut für Angewandte Optik und Feinmechanik Jena, Germany

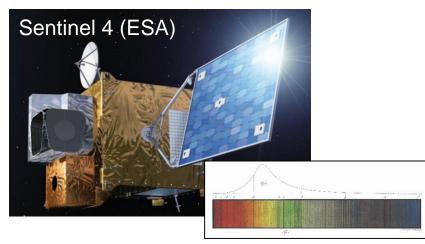
10. October 2017

Electron-beam lithography for grating fabrication

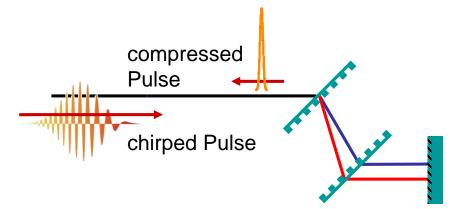
- Examples of astro-gratings:
  - CUBES UV-transmission grating
  - CarbonSat high-resolution gratings
  - > Sub- $\lambda$  structures for ultra-wide-band gratings

# **High Performance Applications of Gratings**

#### Spectrometers for Astronomy and Earth Observation



Manipulation/Compression of Ultra-Short Laser Pulses



#### relevant parameters:

- spectral dispersion
- bandwidth
- efficiency / polarization
- > wavefront
- straylight
- ➢ size, …

# often extreme demands to obtain required performance

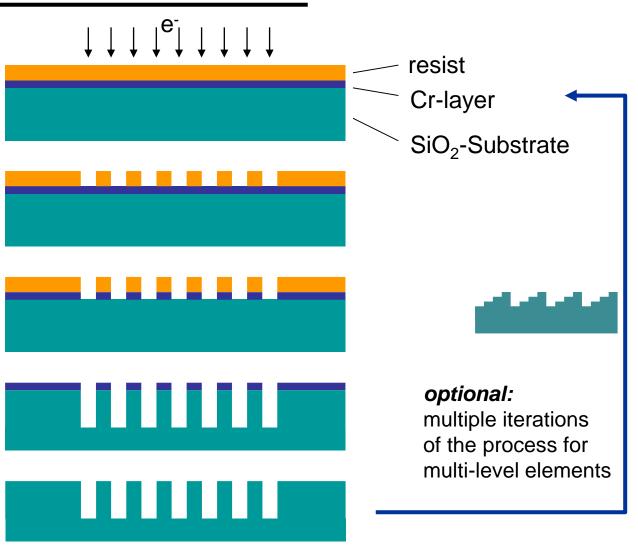




# Grating Technology at the IOF

- 1. Resist exposure with e-beam lithography
- 2. Resist development

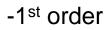
- 3. Chromium etching (RIE)
- 4. Deep etching into substrate (ICP)
- 5. Removal of Cr-layer

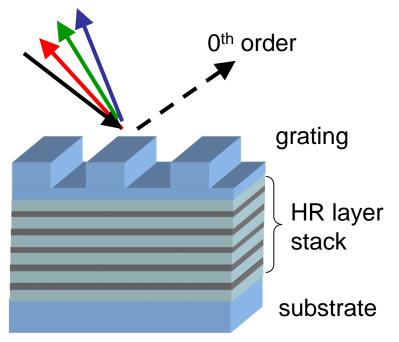


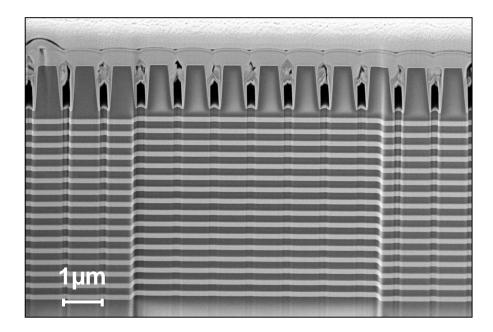




#### **Gratings on dielectric layer stacks**







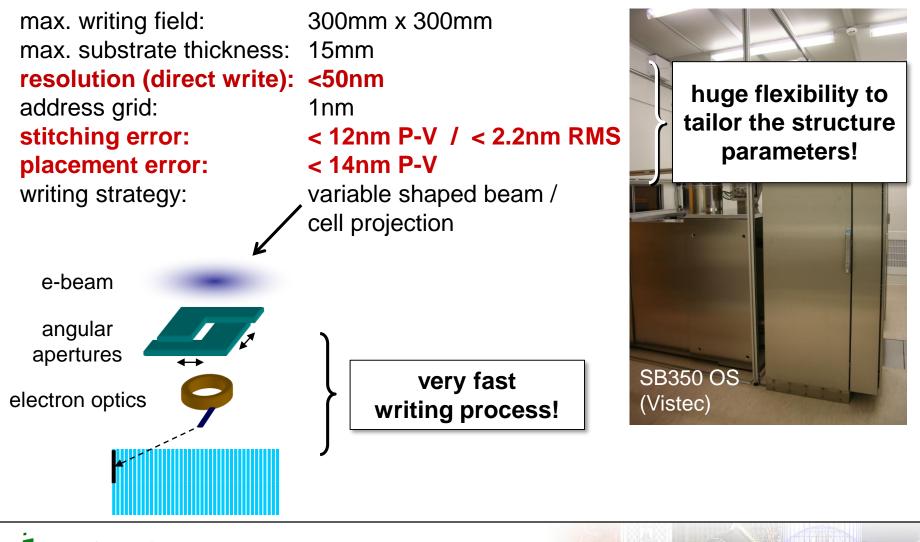
- highly efficient reflection gratings
- transmission gratings with tailored polarization properties







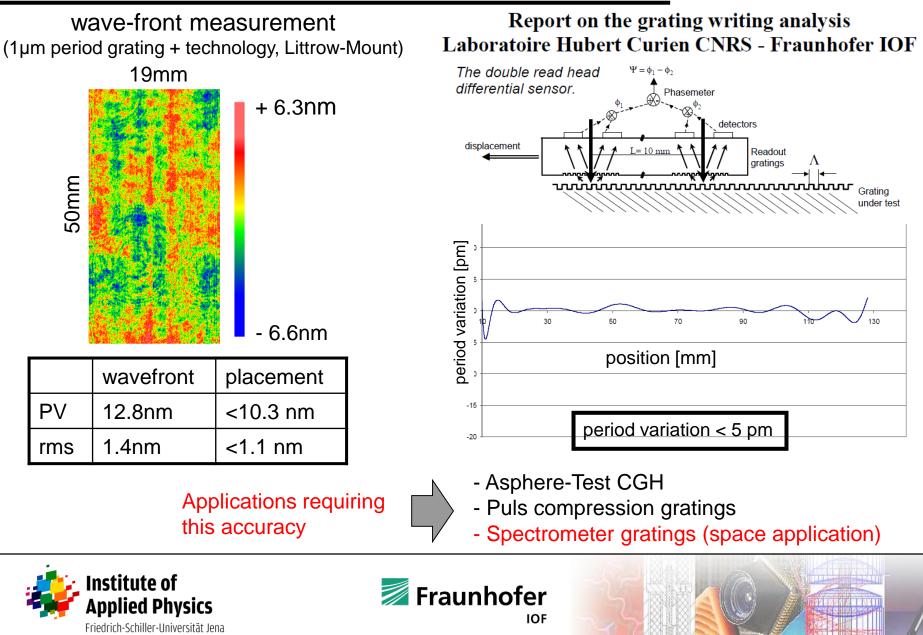
#### The Vistec SB350 OS e-beam writer





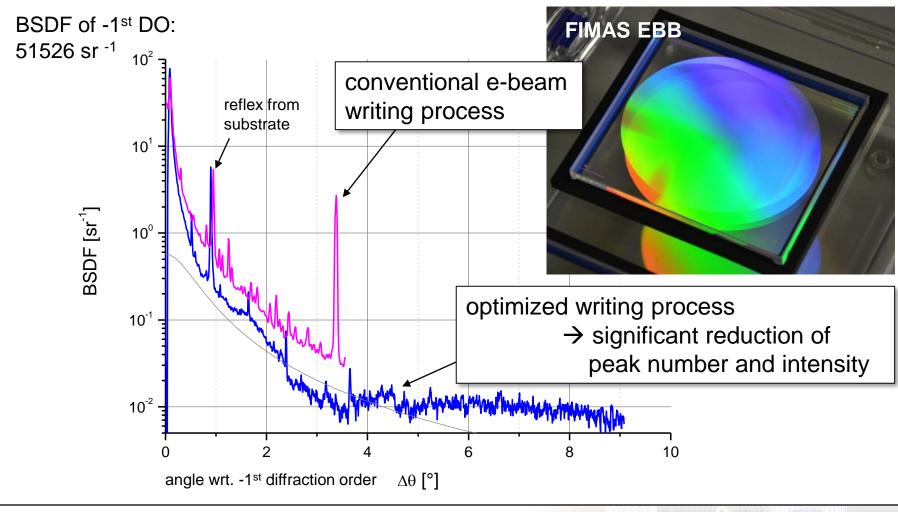


# **Key Performance: Writing Accuracy**



#### Accuracy of writing process: straylight

**Optimization of e-beam writing process** 







Examples of realized spectrometer gratings





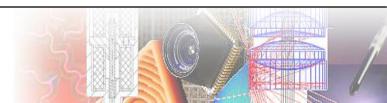


## **CUBES – UV Transmission Grating**

- CUBES (Cassegrain U-Band Brazilian ESO-Spectrograph)
- > Requirements:
  - spectral band: 300nm 400nm
  - line density: 3448 lines/mm  $\rightarrow$  p=290nm
  - AOI: 31°
  - grating size: 250 x 250 mm<sup>2</sup>; mosaic of 2x [250mm x 130mm]
- Challenges:
  - commercial VPH gratings difficult in the UV
- Solution:
  - Binary fused silica gratings

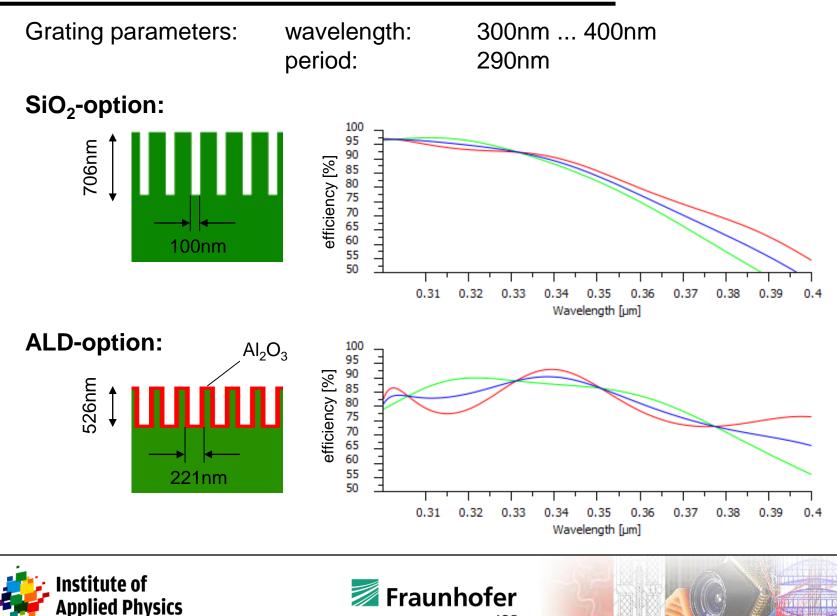




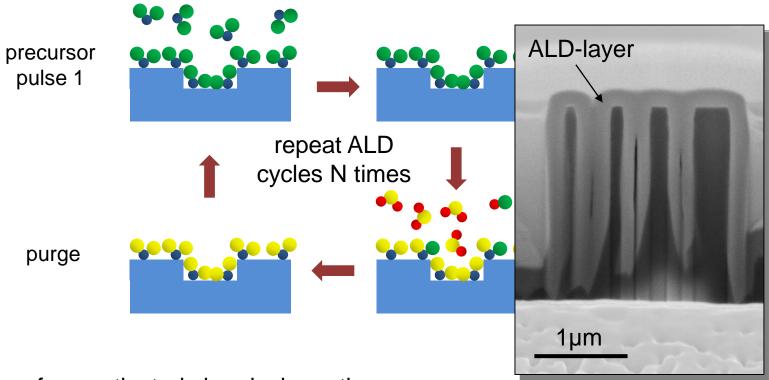


### **ESO Cubes Spectrometer**

Friedrich-Schiller-Universität Jena



### **Atomic-Layer-Deposition (ALD)**

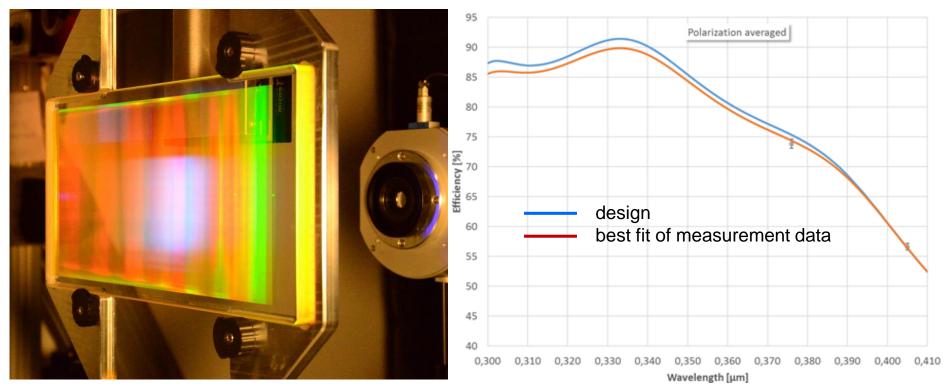


- surface activated chemical reactions
- conformal overcoating of surface reliefs
- large number of materials possible, e.g.  $TiO_2$ ,  $Ta_2O_5$ ,  $Al_2O_3$ ,  $HfO_2$  ...





### **CUBES – UV Transmission Grating**



realized grating during efficiency measurement

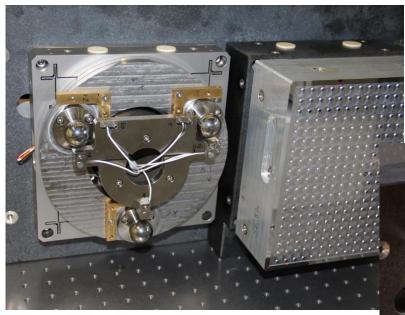
#### grating size: 250mm x 130mm







### **Tiling for Larger Gratings**



active alignment for wave-front optimization

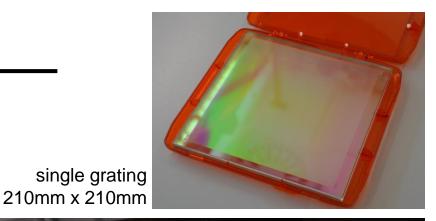
→ also possible for transmission gratings

arrangement of 2 reflection gratings (420mm x 210mm)



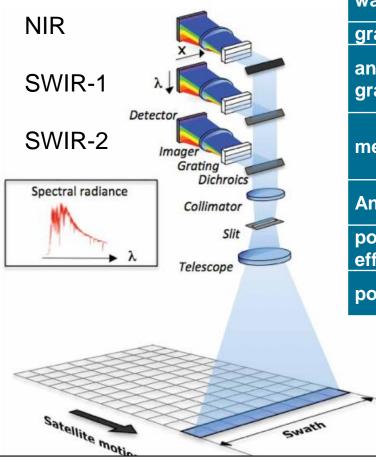






### **Carbon Monitoring Satellite (CarbonSat)**

#### instrument concept:



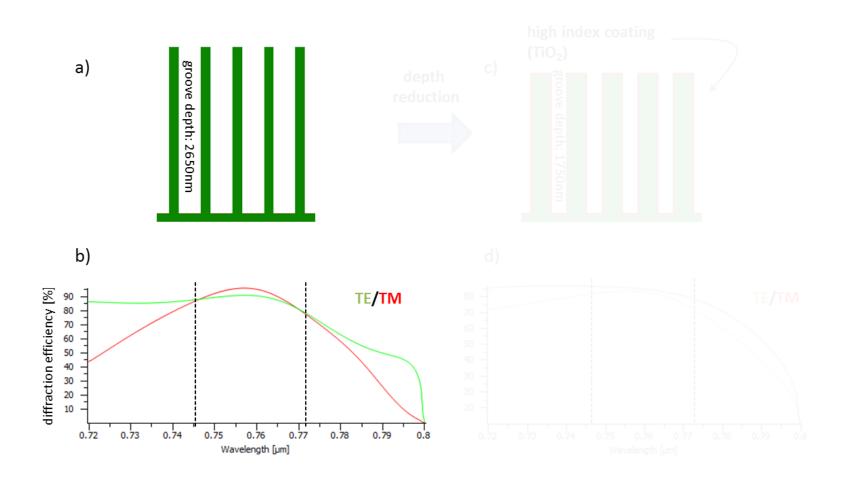
Parameter	NIR	SWIR-1
wavelength	747nm 773nm	1590nm 1675nm
grating period	423nm	991nm
angle of incidence to the grating (equivalent in air)	<b>63.6°</b>	55.5°
mean angle of diffraction	Transmission Gratings in -1. order Littrow configuration	
Angular dispersion	0.3° / nm	0.1°/nm
polarization avg. efficiency	>70%	>70%
polarization sensitivity	<10%	<10%







## **NIR – High Resolution Transmission Grating**

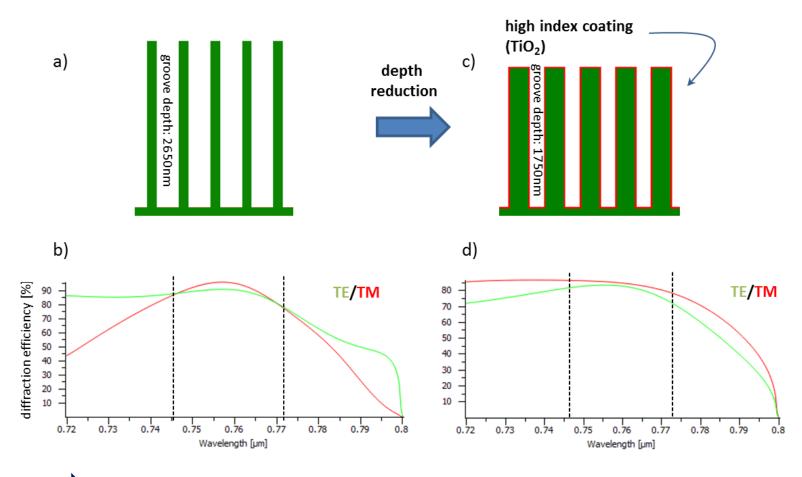








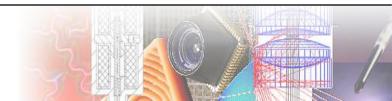
# NIR – High Resolution Transmission Grating

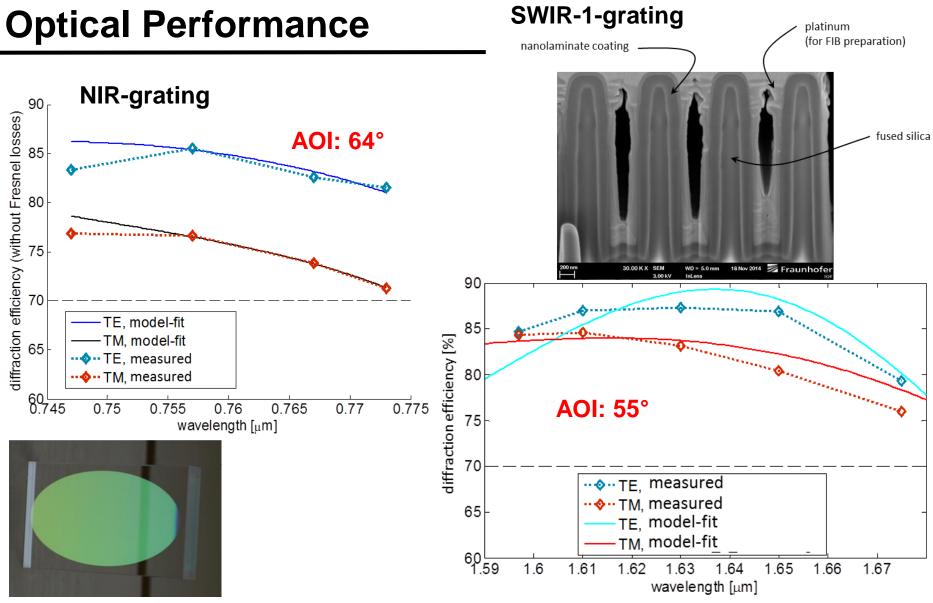


use high-refractive-index (dielectric) coating to reduce depth









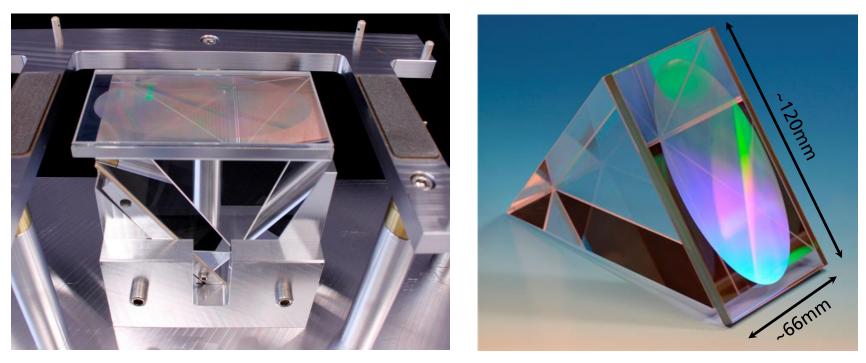






#### **Direct Glass-to-Glass Bonding**

# Advantages: adhesive free glass-to-glass connection no additional optical interface



- achieved alignment accuracy: 0.25mrad (< 1 arcmin)
- bond strength up to 2/3 of bulk fused silica
- current TRL: 6







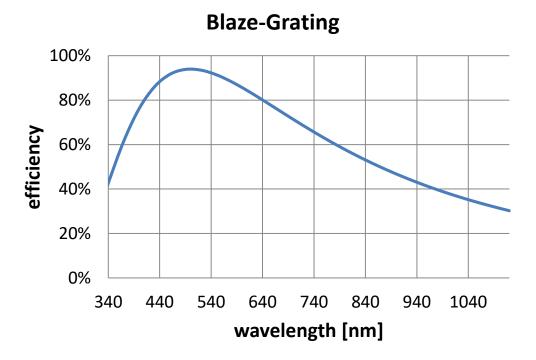
#### **Wide-Band Gratings**

- typical requirements for a low-resolution, broad-band disperser
  - spectral range: several 100nm
  - AOI: near-perpendicular incidence
  - period: few µm



 blazed-grating in low order (saw-tooth profile)



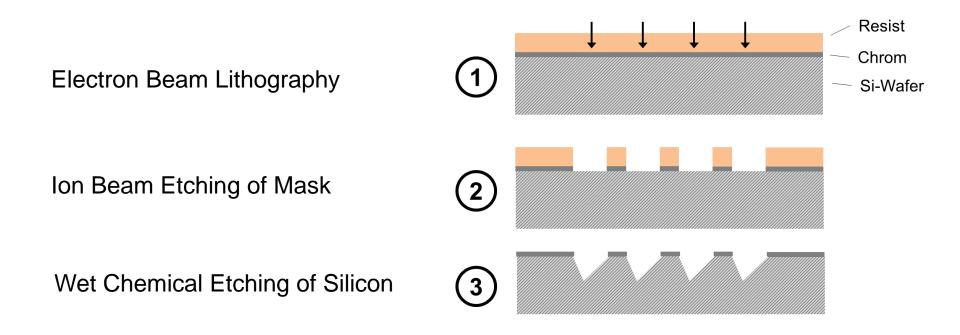








#### **Echelle or Echellette Structures**



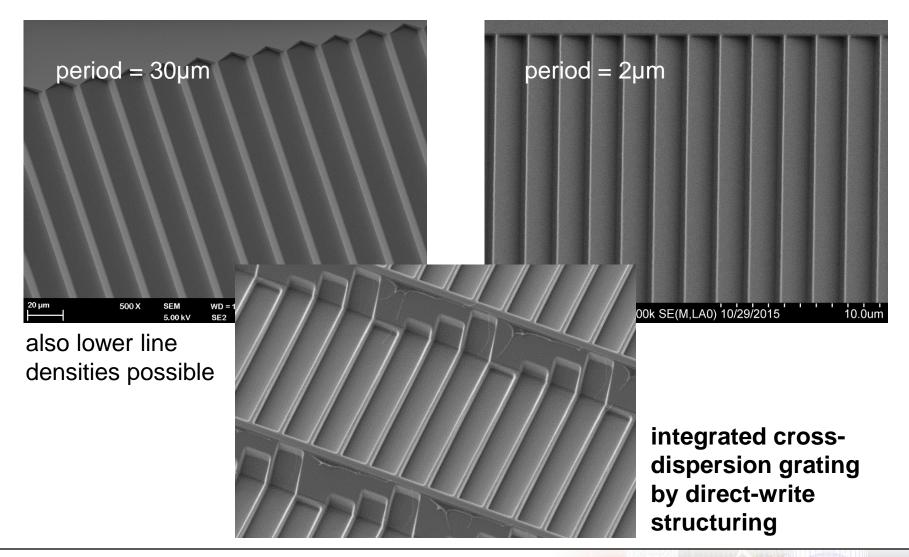


"Blaze Angle" can be adjusted by crystalline orientation of Silicon substrate





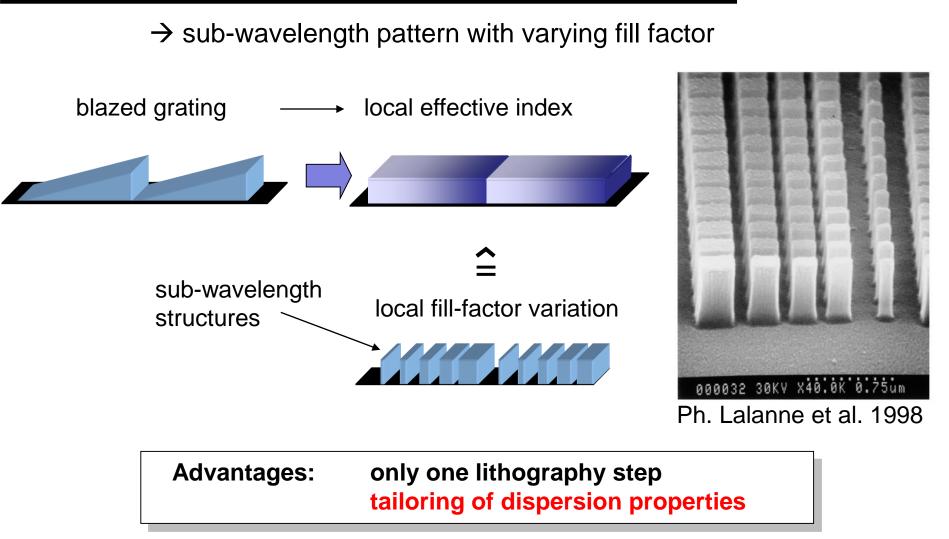
#### **Echelle or Echellette Structures**







## **Alternative: Effective Index Gratings**



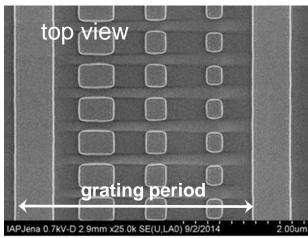


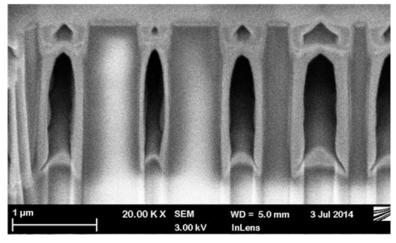




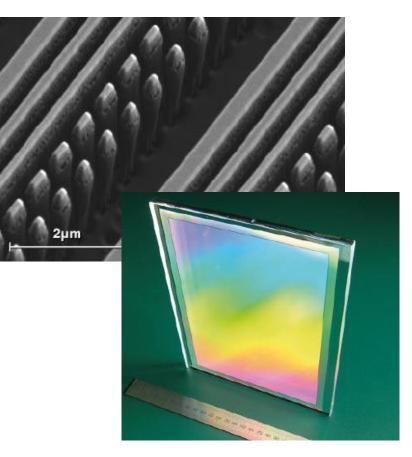
### **Effective Medium Gratings**

#### FLEX (fluorescence explorer); [500nm – 800nm]





# GAIA (global astrometic interferometer for astrophysics); [750nm – 800nm]





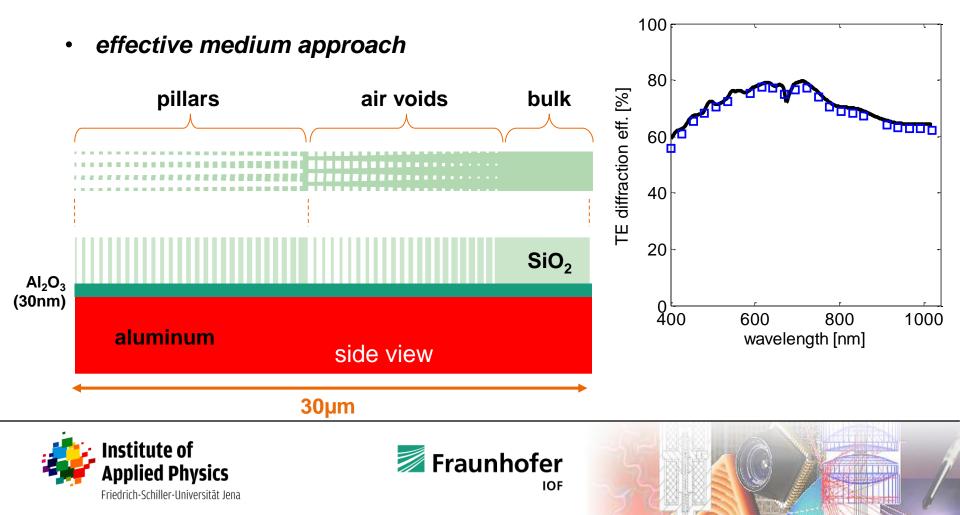




## **Wide-Band Reflection Grating**

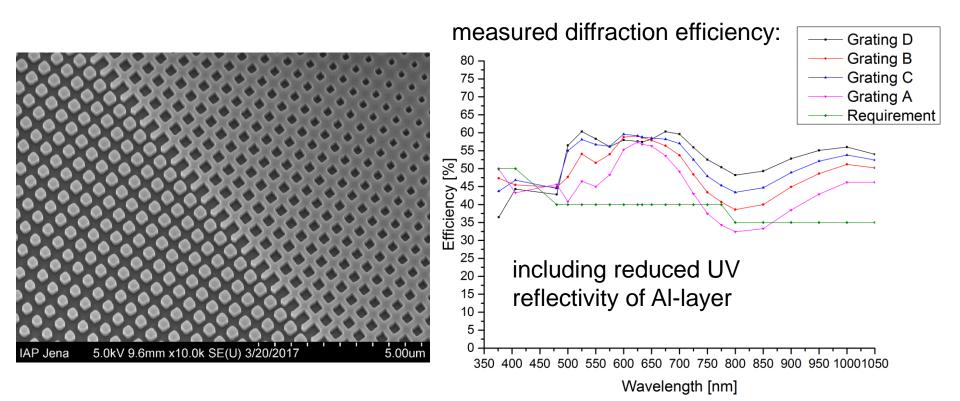
- typical requirements
  - based on a concave grating
  - spectral range: **340nm 1050nm**

- AOI: 0.5°
- period: 30µm



### Wide-Band Reflection Grating ...

... realized by E-beam lithography



#### $\rightarrow$ very weak spectral dependency of diffraction efficiency



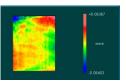


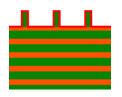


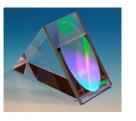
## Summary



• Direct write electron-beam lithography has a huge potential for the realization of high-performance gratings







Sub-period engineering by combining E-Beam lithography and Atomic-Layer-Deposition

To make use of the large flexibility and the advantageous optical properties requires talking with the grating manufacturer already during the design of the instrument !!! (not after PDR...)

- Examples are:
  - high resolution gratings with low polarization sensitivity
  - echelle-type gratings with integrated cross-disperse
  - ultra-wide-band gratings for lower resolution spectrometers





