

C I T
E X T
P O À

Rosso Marziano

l'alcool nello spazio

Anna Wolter

INAF – Osservatorio Astronomico di Brera

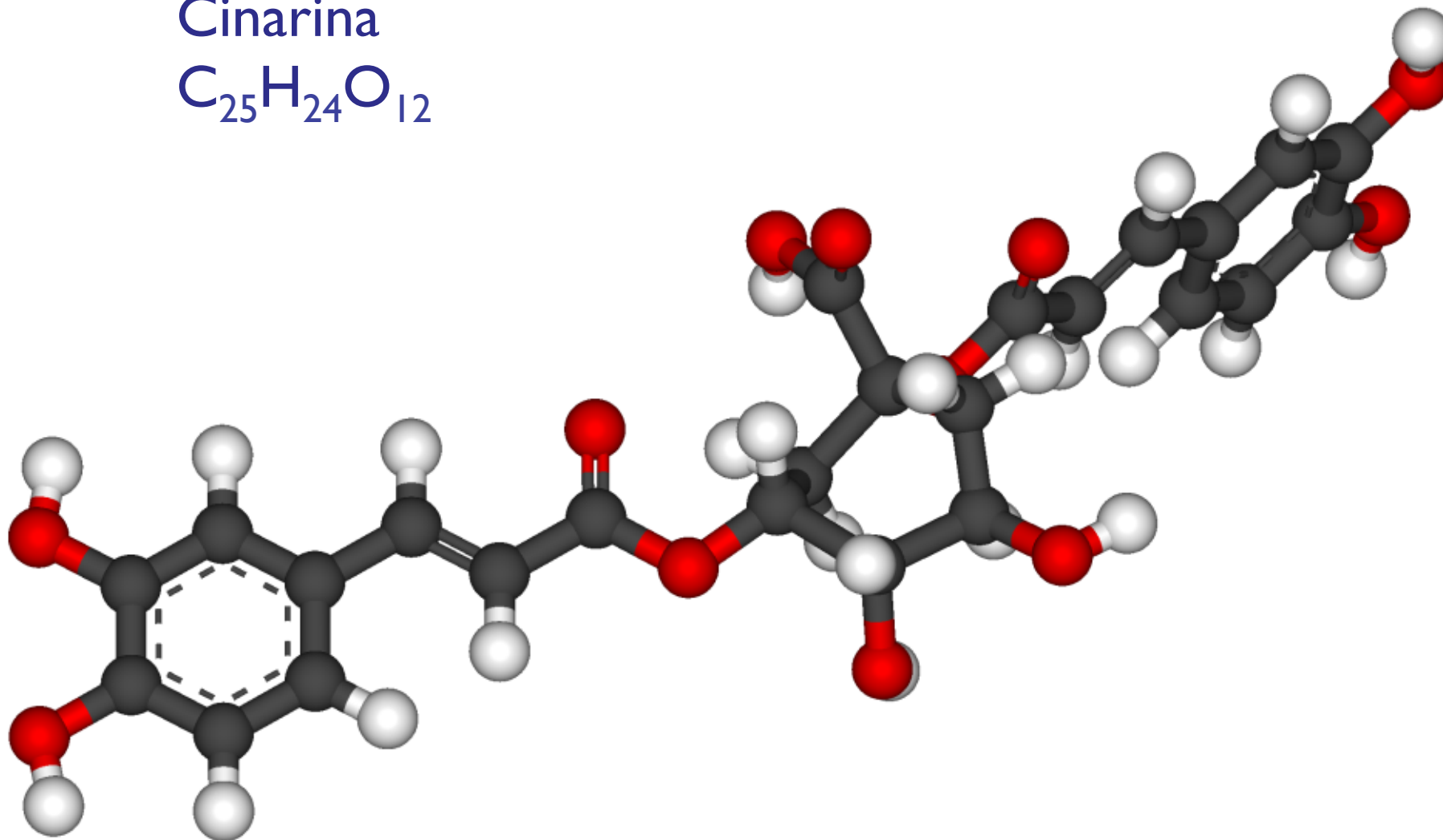
ESO Science Outreach Network

2015 **Milano**
a place
to **BE**

Il Vino



Cinarina

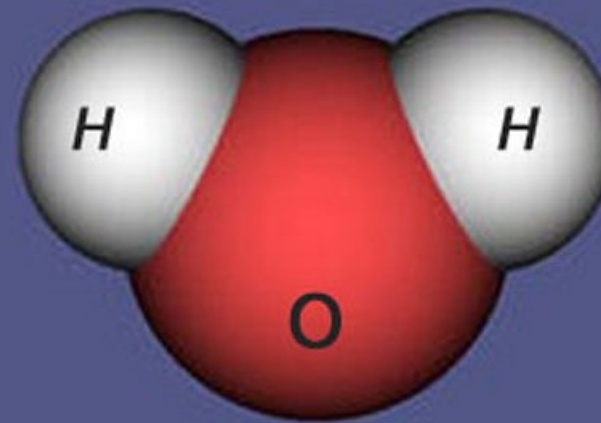


Acqua

- Elemento fondamentale: lo rivedremo

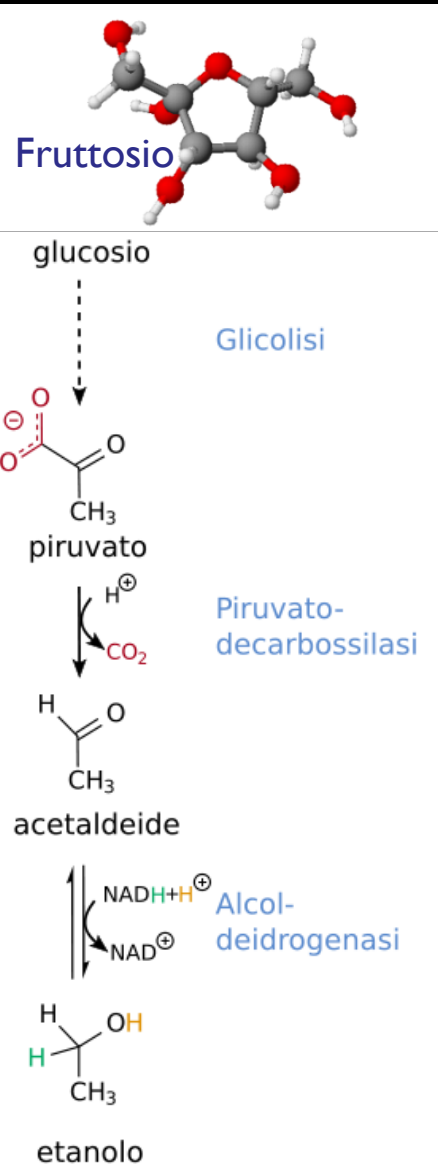


MOLECOLA D'ACQUA



H=IDROGENO
O=OSSIGENO

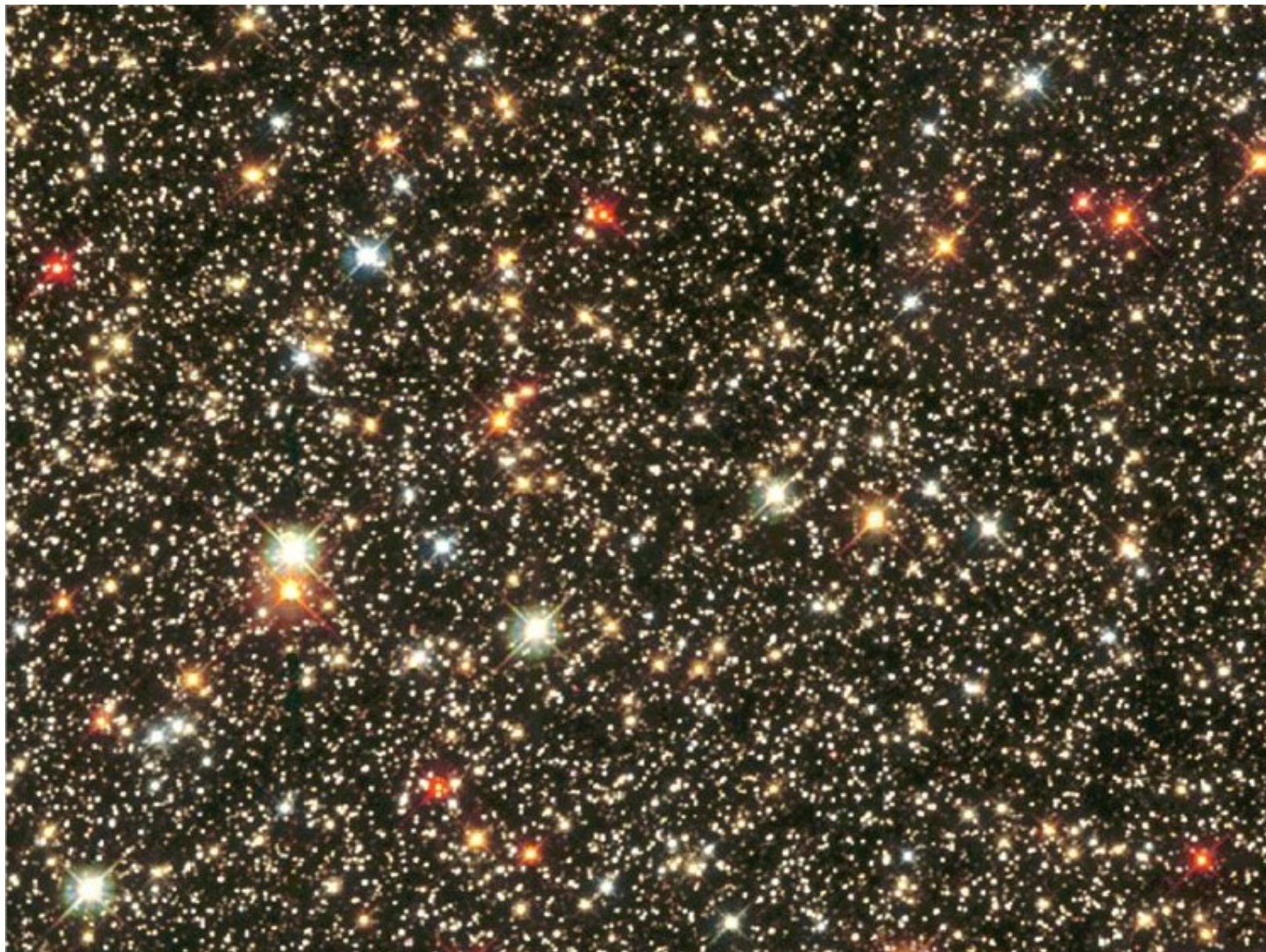
Alcol e zuccheri



Gli zuccheri si trasformano in alcoli per mezzo dell'azione di lieviti e batteri (saccaromiceti)





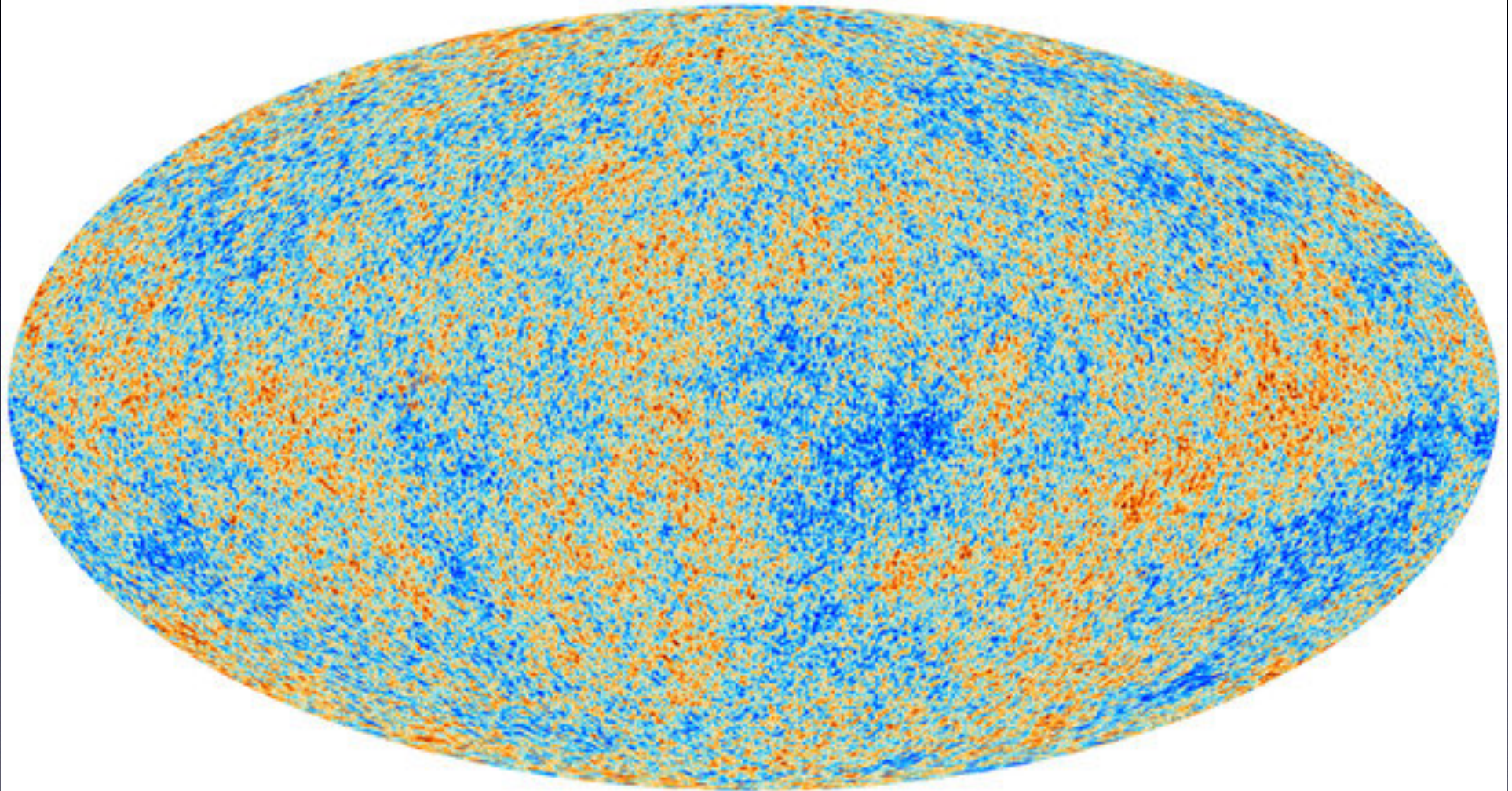




Ma come si formano gli elementi?

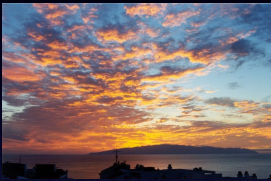


STEVEN WEINBERG
I PRIMI TRE MINUTI

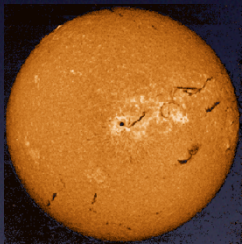


E il resto?

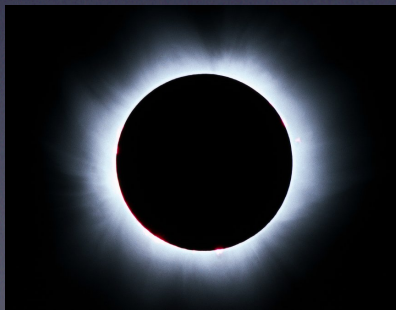
- Reazioni termonucleari:



luminoso



colorato



sferico



...e ha circa 4 miliardi e mezzo di anni!

Stella:

corpo celeste costituito da gas caldo tenuto insieme dalla gravità e che emette luce propria
Dizionario della lingua italiana (Devoto –Oli)

La luminosità del Sole

La quantità di energia emessa dal Sole al secondo è:

$$L_{sole} \approx 4 \times 10^{26} \text{ J / s} = 4 \times 10^{26} \text{ W}$$

Quella che raggiunge la Terra è circa 175 000 TW (TeraWatt)

Quella consumata in media sulla Terra è circa 13 TW (1TW = 10^{12} W)

Una famiglia media consuma circa 20 MW h all'anno (1MW = 10^6 W = 10^{-6} TW)

Qual è la sorgente di questa energia?

La contrazione no: vita del Sole troppo breve circa 20 milioni di anni

Ipotesi termonucleare: Eddington (ca 1920)



1957

REVIEWS OF MODERN PHYSICS

VOLUME 29, NUMBER 4

OCTOBER, 1957

Synthesis of the Elements in Stars*

E. MARGARET BURRIDGE, G. R. BURRIDGE, WILLIAM A. FOWLER, AND F. HOYLE

*Kellogg Radiation Laboratory, California Institute of Technology, and
Mount Wilson and Palomar Observatories, Carnegie Institution of Washington,
California Institute of Technology, Pasadena, California*

"It is the stars, The stars above us, govern our conditions";
(*King Lear*, Act IV, Scene 3)

but perhaps

"The fault, dear Brutus, is not in our stars, But in ourselves,"
(*Julius Caesar*, Act I, Scene 2)

TABLE OF CONTENTS

	Page
I. Introduction.....	548
A. Element Abundances and Nuclear Structure.....	548
B. Four Theories of the Origin of the Elements.....	550
C. General Features of Stellar Synthesis.....	550
II. Physical Processes Involved in Stellar Synthesis, Their Place of Occurrence, and the Time-Scales Associated with Them.....	551
A. Modes of Element Synthesis.....	551
B. Method of Assignment of Isotopes among Processes (i) to (viii).....	553
C. Abundances and Synthesis Assignments Given in the Appendix.....	555
D. Time-Scales for Different Modes of Synthesis.....	556
III. Hydrogen Burning, Helium Burning, the α Process, and Neutron Production.....	559
A. Cross-Section Factor and Reaction Rates.....	559
B. Pure Hydrogen Burning.....	562
C. Pure Helium Burning.....	565
D. α Process.....	567
E. Succession of Nuclear Fuels in an Evolving Star.....	568
F. Burning of Hydrogen and Helium with Mixtures of Other Elements; Stellar Neutron Sources.....	569
IV. ϵ Process.....	577
V. s and r Processes: General Considerations.....	580
A. "Shielded" and "Shielding" Isobars and the s , r , p Processes.....	580
B. Neutron-Capture Cross Sections.....	581
C. General Dynamics of the s and r Processes.....	583
VI. Details of the s Process.....	583

* Supported in part by the joint program of the Office of Naval Research and the U. S. Atomic Energy Commission.

547

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Sono le stelle! le stelle
al di sopra di noi che
governano la nostra
condizione!

Re Lear atto IV scena III

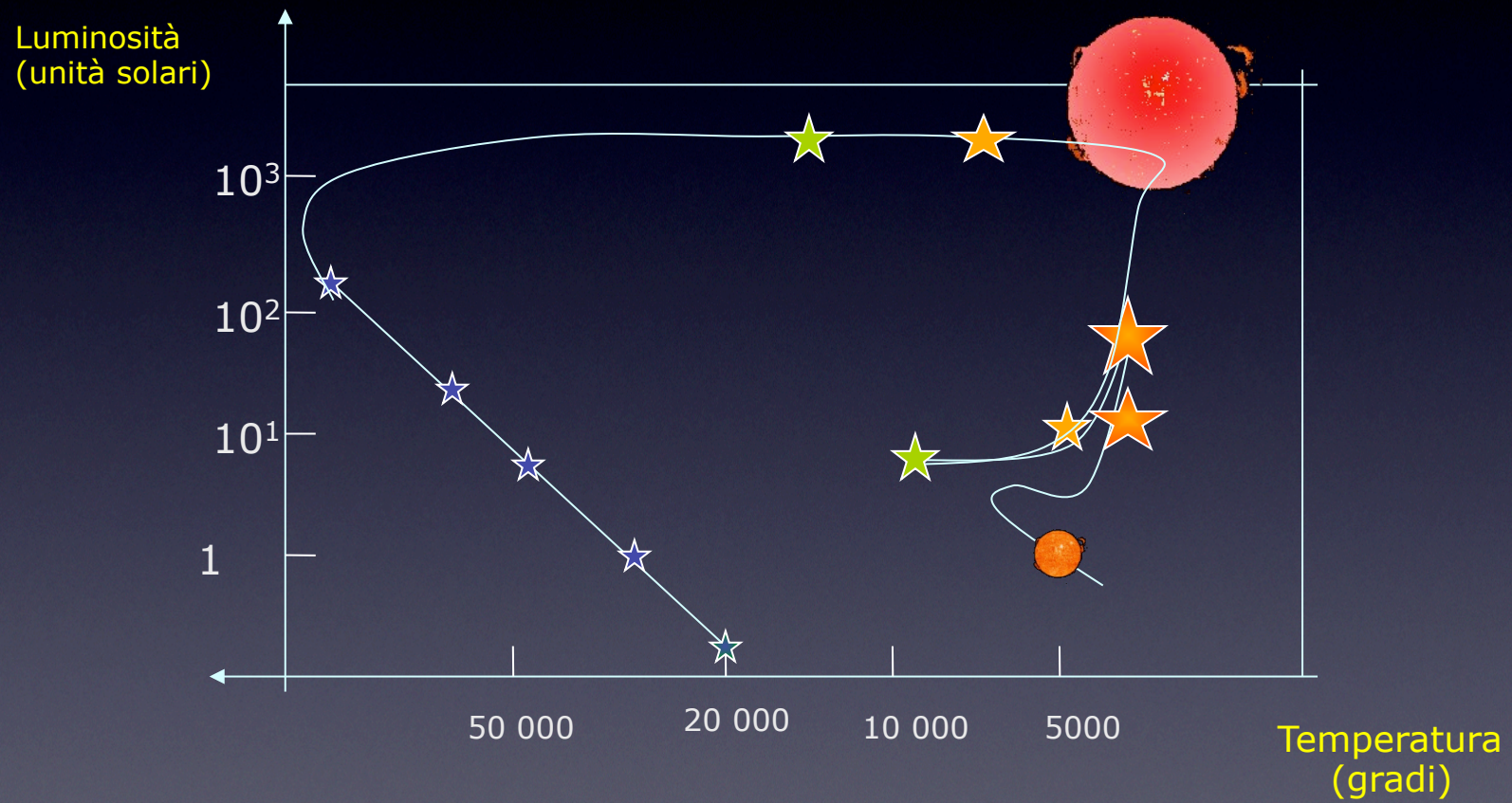


La catena protone protone



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L'evoluzione di una stella come il Sole



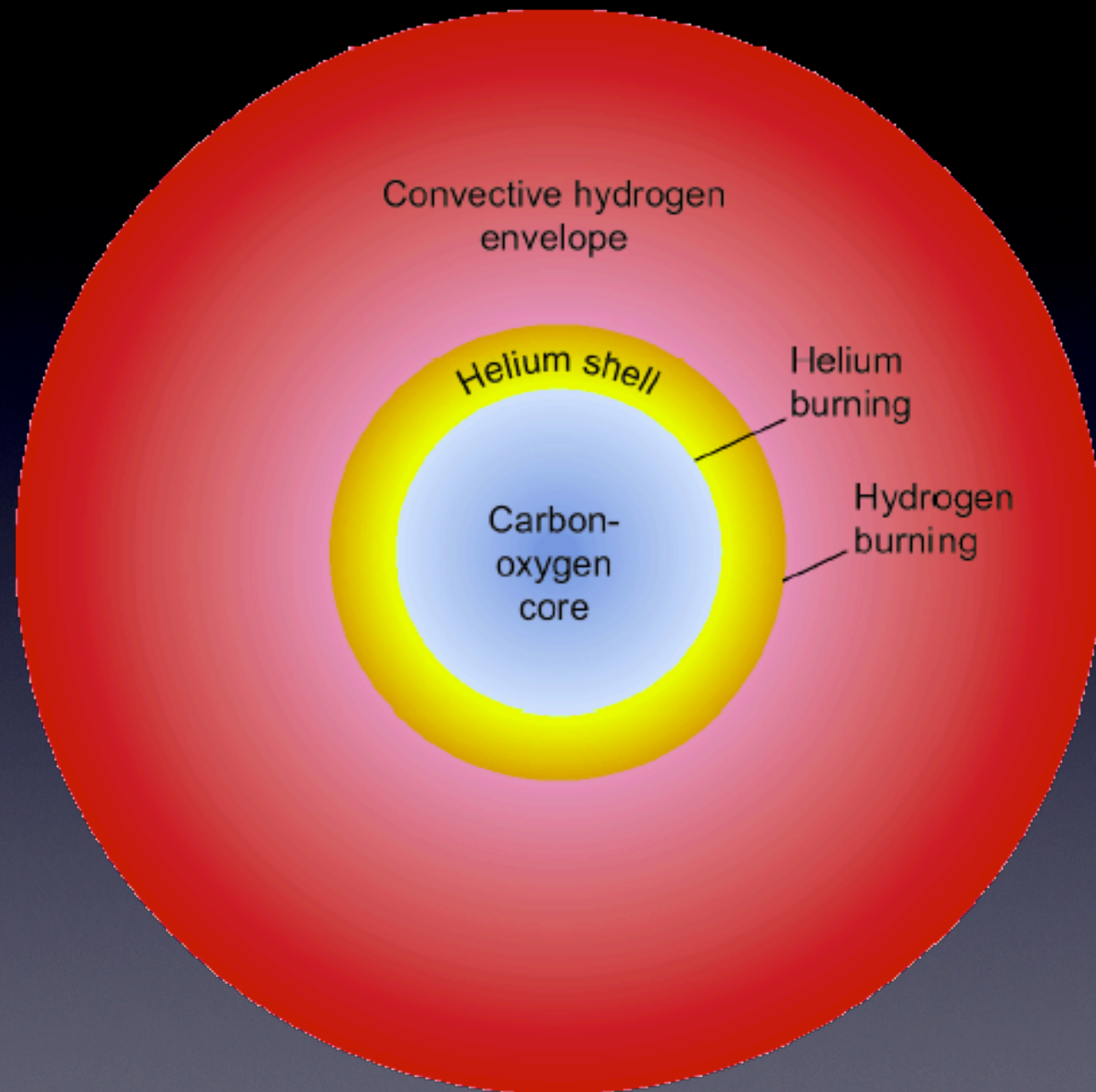


Gas diffuso



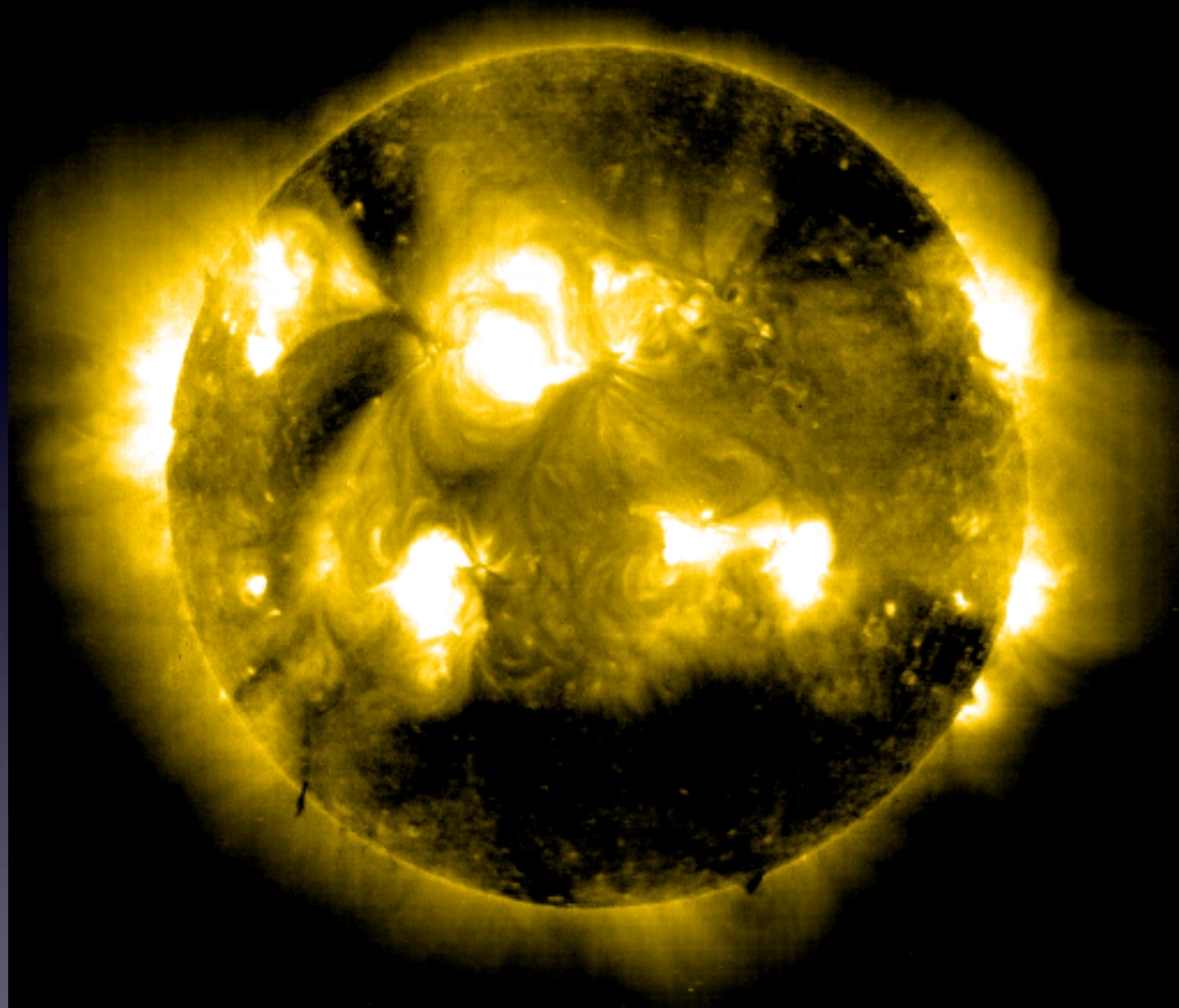
*Nubi
molecolari*



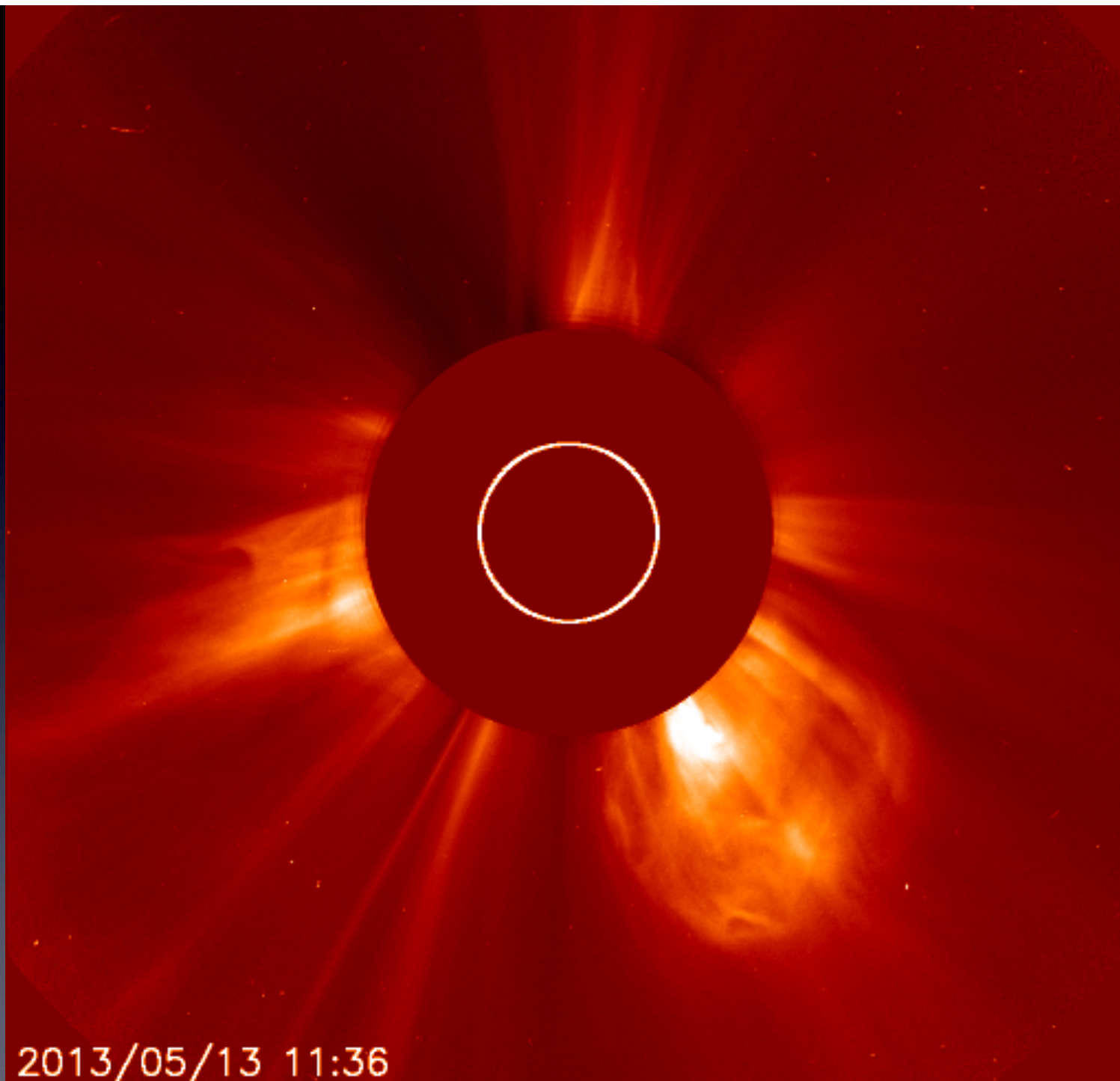


La combustione dell'elio



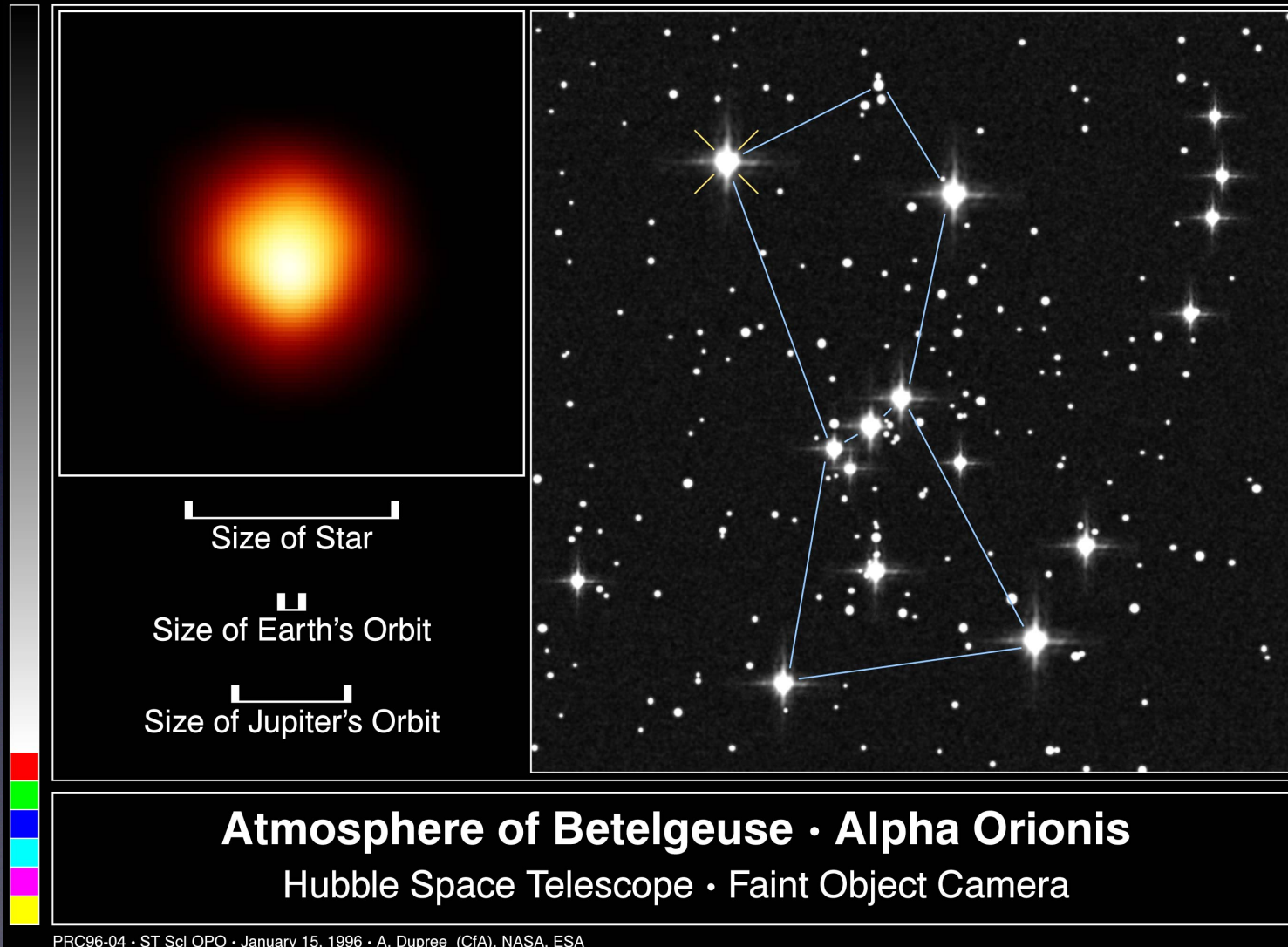


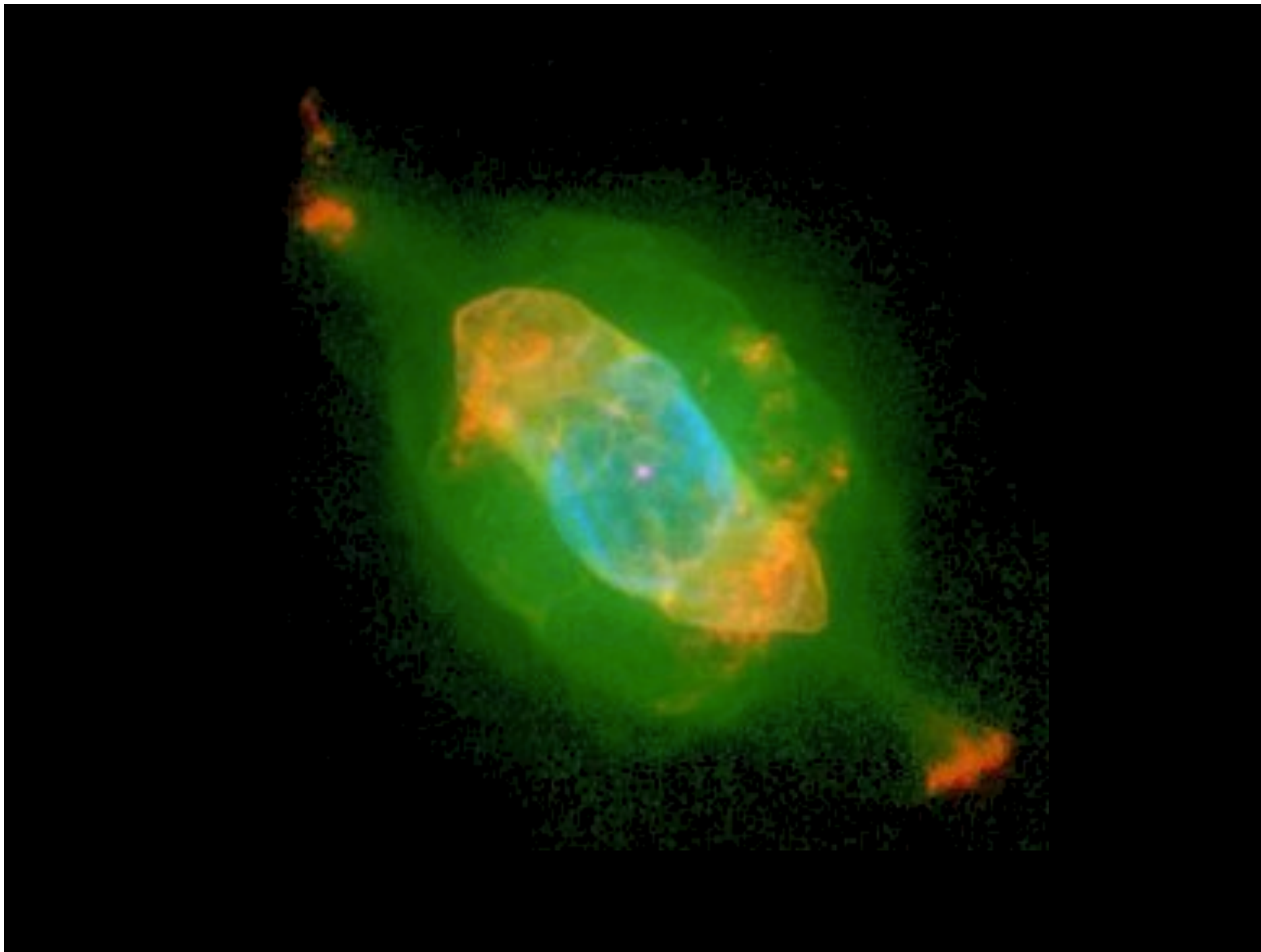
2013/03/08 13:06



2013/05/13 11:36

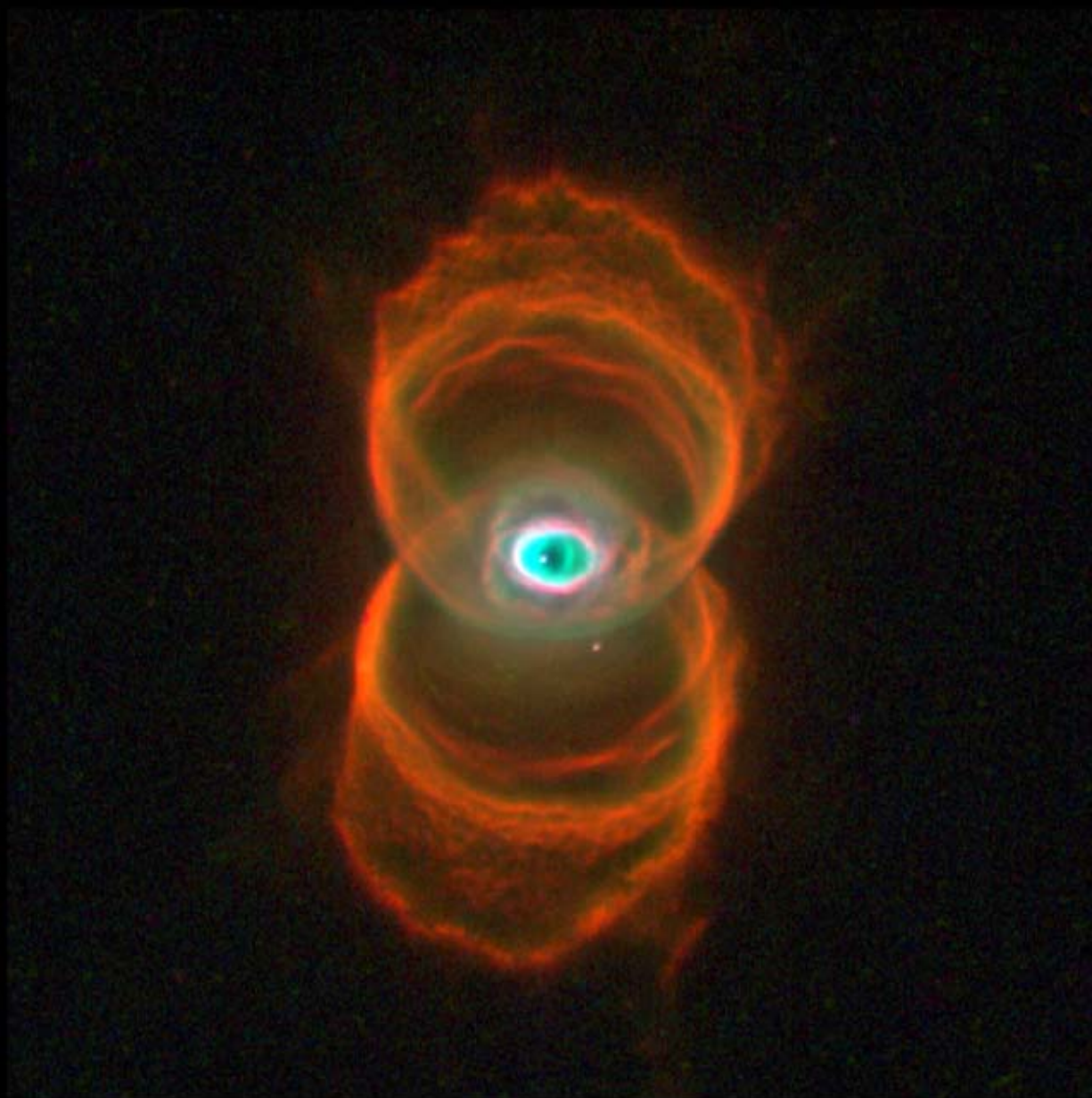
Una gigante rossa









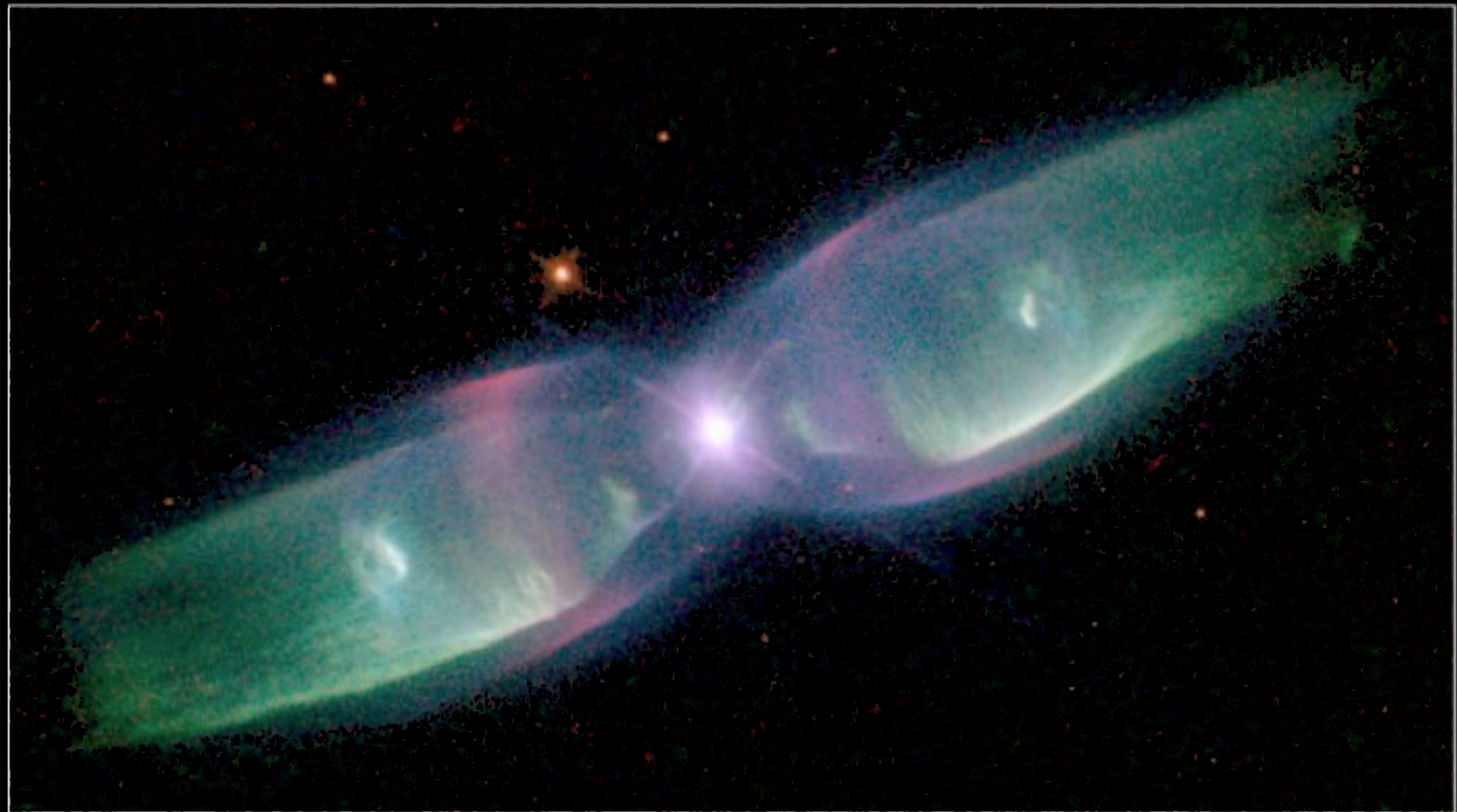


Hourglass Nebula · MyCn18

HST · WFPC2

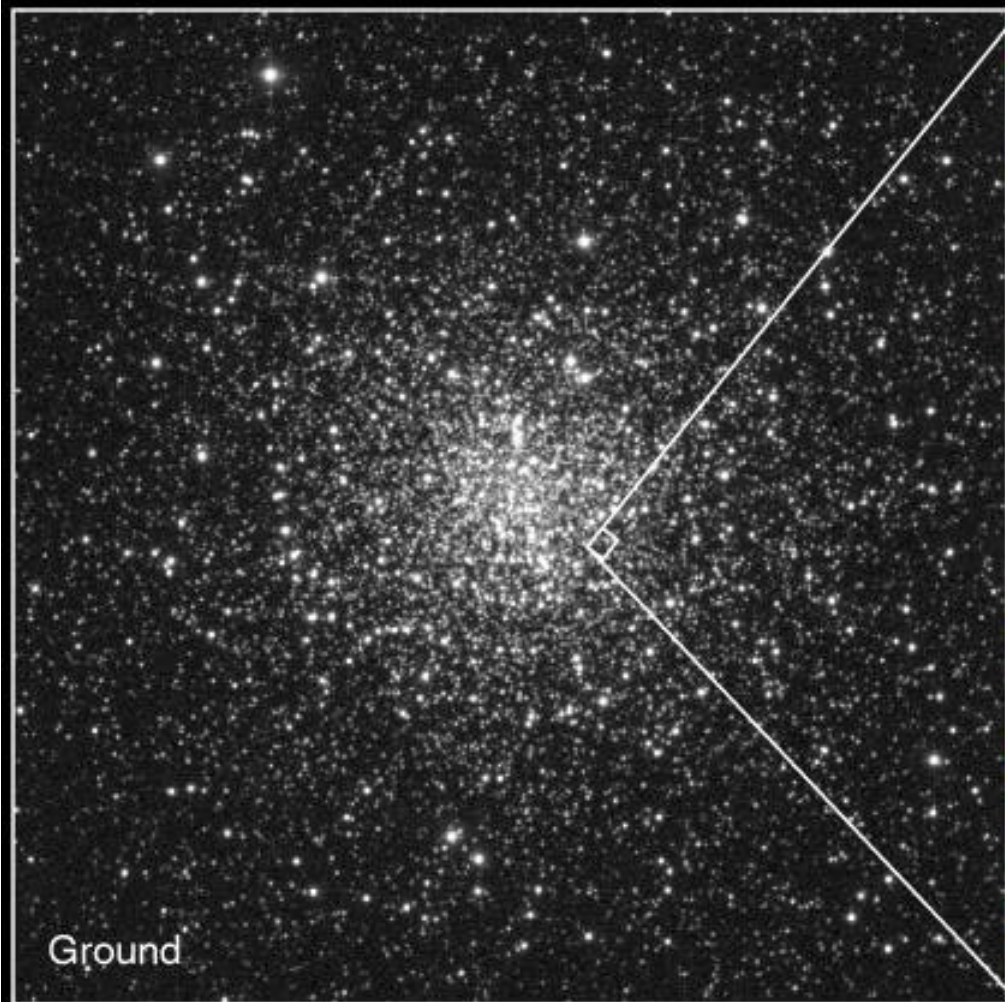
PRC96-07 · ST ScI OPO · January 16, 1996

R. Sahai and J. Trauger (JPL), the WFPC2 Science Team and NASA

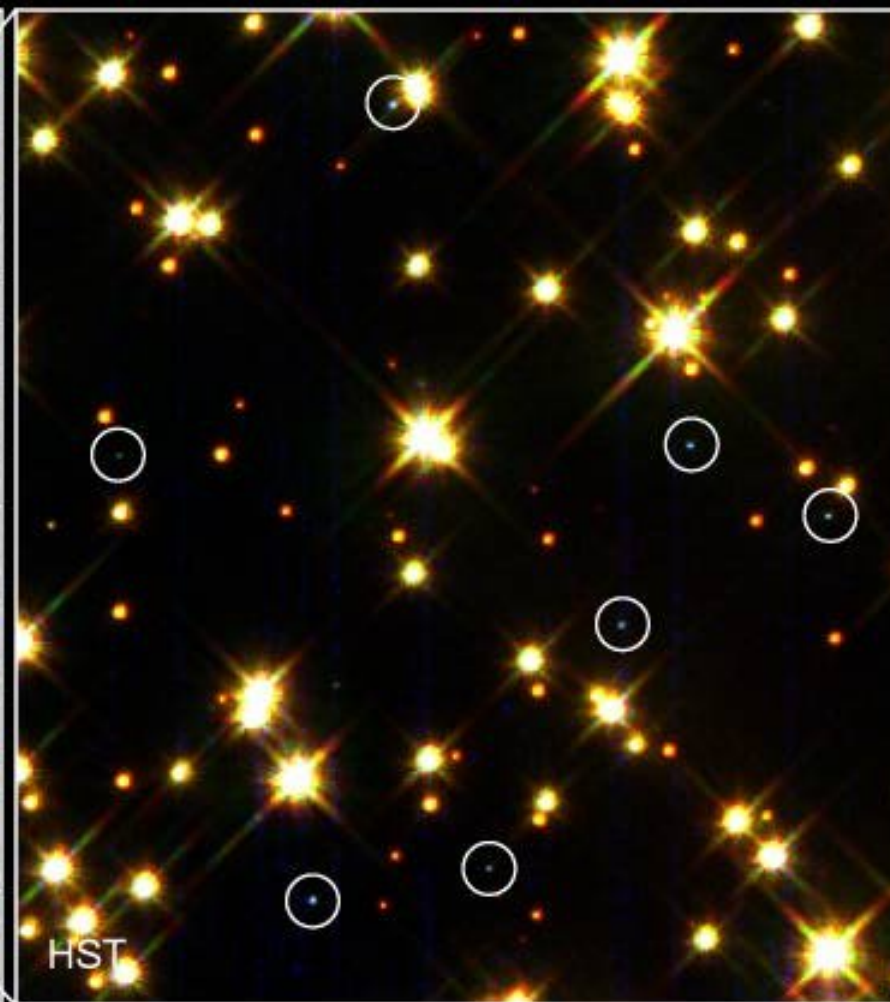


Planetary Nebula M2-9
Hubble Space Telescope • WFPC2

PRC97-36a • ST ScI OPO • December 17, 1997 • B. Balick (University of Washington) and NASA



Ground



HST

White Dwarf Stars in M4

PRC95-32 · ST ScI OPO · August 28, 1995 · H. Bond (ST ScI), NASA

HST · WFPC2

L'evoluzione del Sole

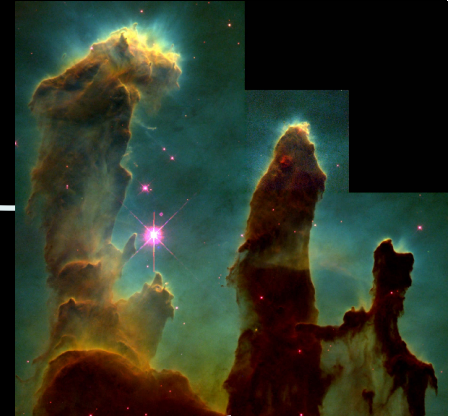
Gigante
rossa

Sequenza
principale

Vita media del processo
10 miliardi di anni = 70% età Universo

Nebulosa
planetaria

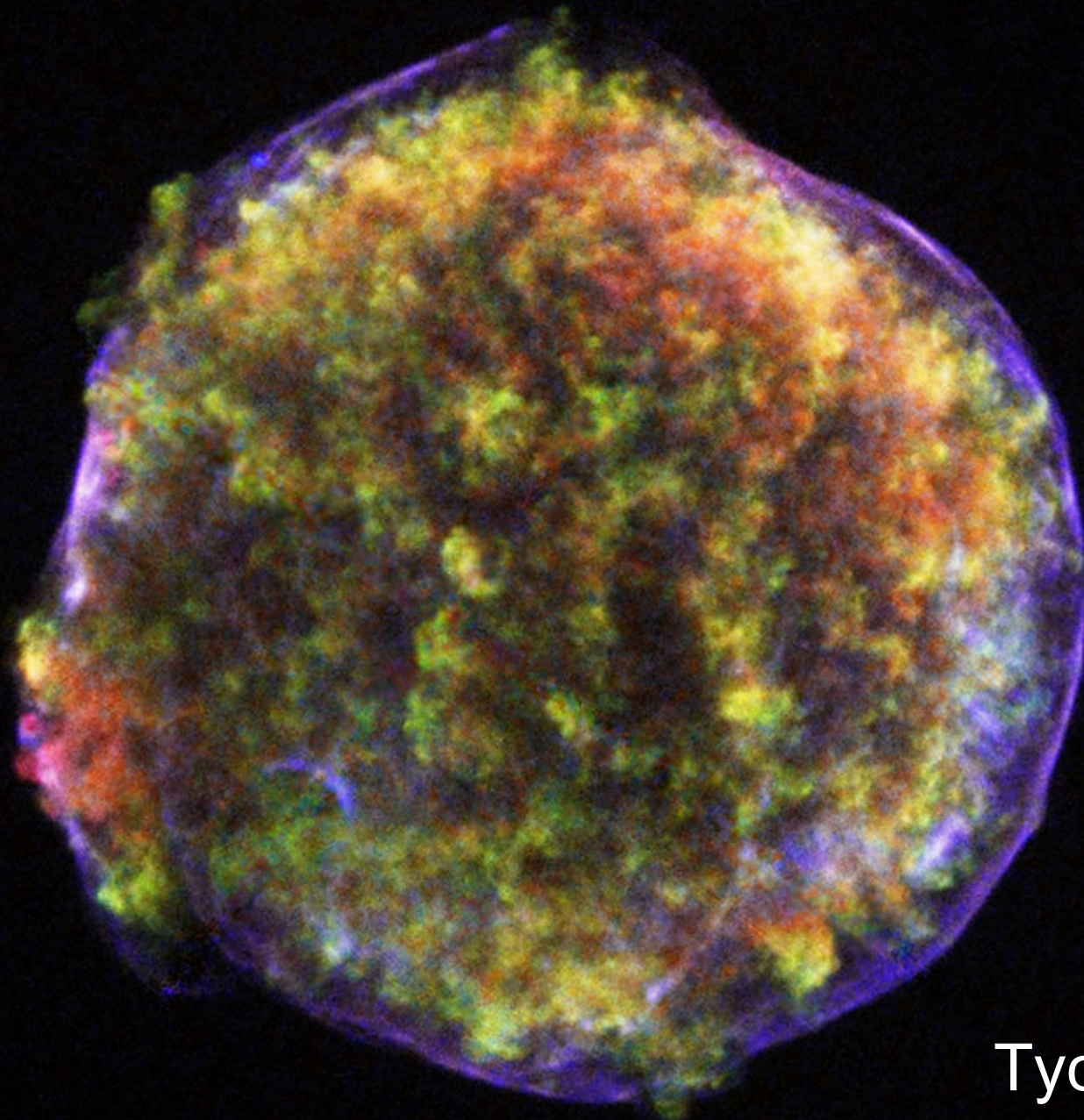
Nana
bianca





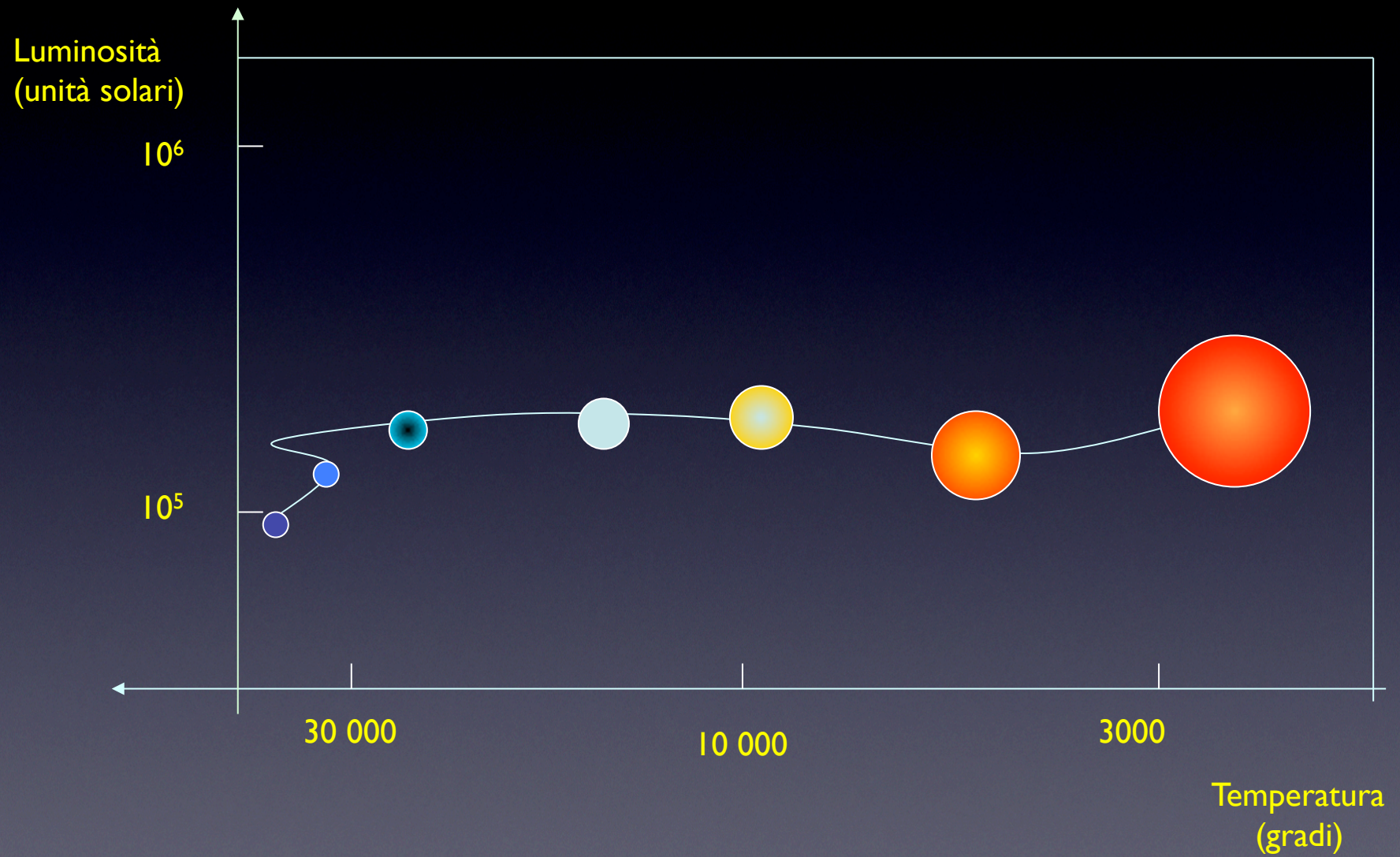
SN 185
3300 anni luce
direz Alfa Cen
SN Ia

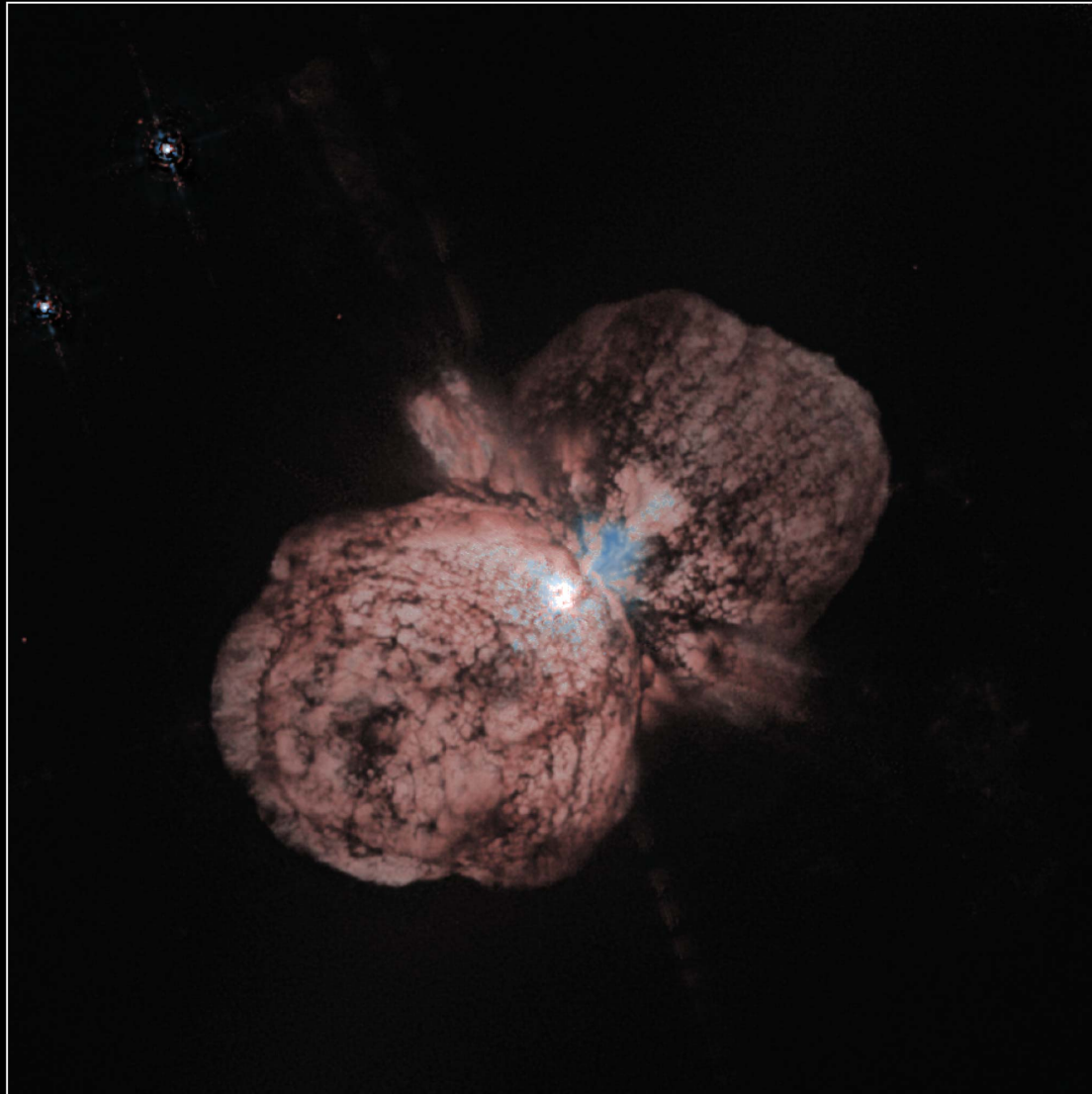




Tycho SNR

L'evoluzione di una stella di 25 masse solari

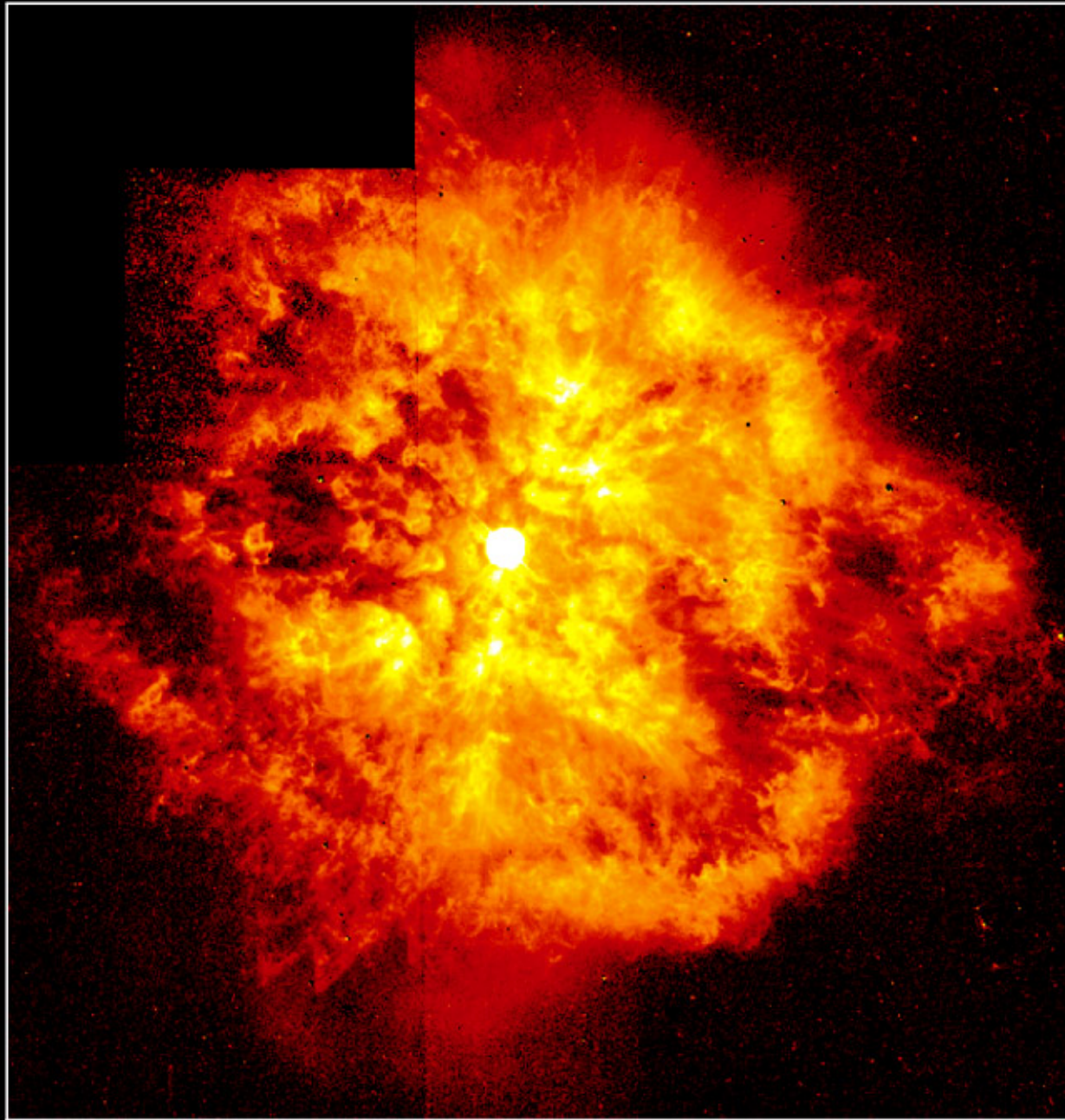




Eta Carinae
Hubble Space Telescope • WFPC2

PRC96-23a • ST ScI OPO • June 10, 1996 • J. Morse (U. CO), K. Davidson (U. MN) and NASA

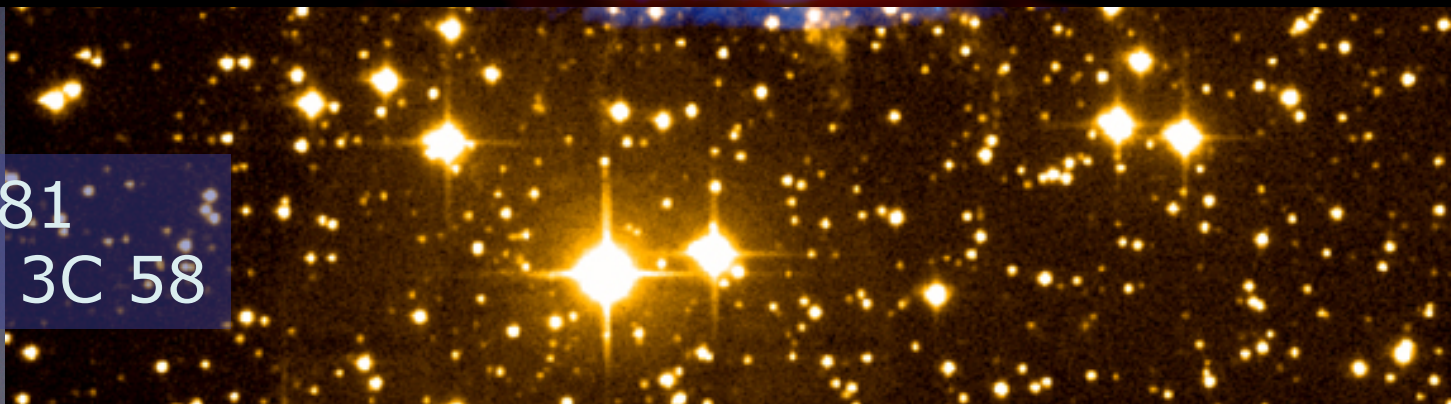
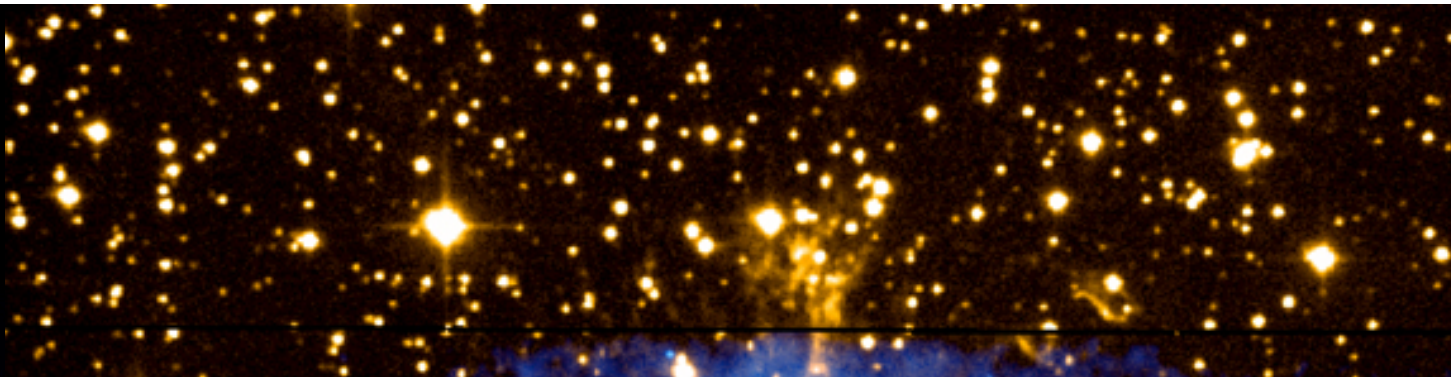




Nebula M1-67 around Star Wr224

HST • WFPC2

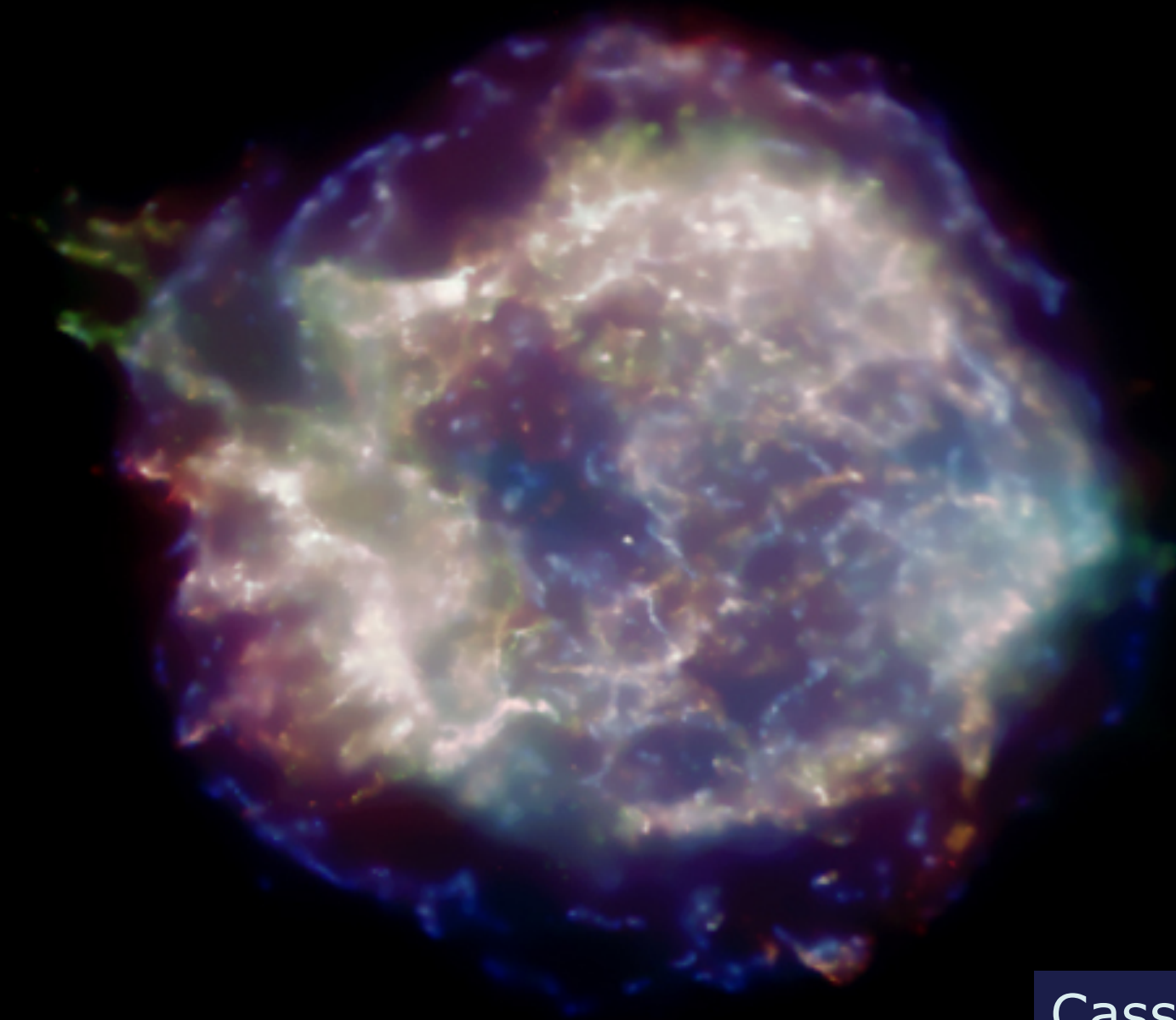
PRC98-38 • STScI OPO • November 3, 1998 • NASA



SN 1181
Pulsar 3C 58

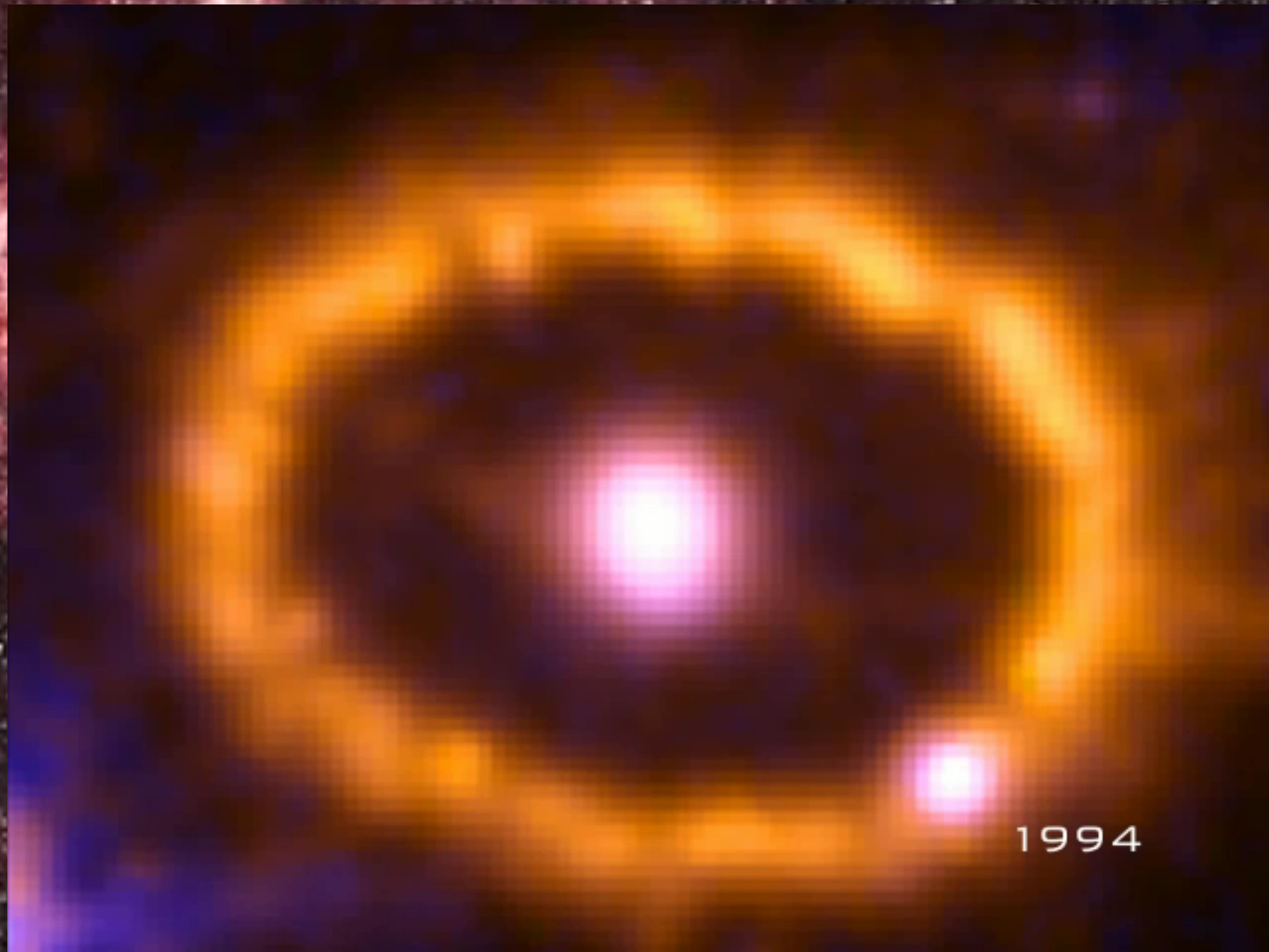
A deep-field astronomical image of the Crab Nebula (SN 1054). The nebula is a complex, multi-colored structure of ionized gas. It features a bright, blue-white central core surrounded by a dense, swirling ring of green and yellow. The outer edges are composed of wispy, reddish-pink filaments. Numerous small, bright stars are scattered throughout the dark background of space.

SN 1054
Nebulosa Granchio

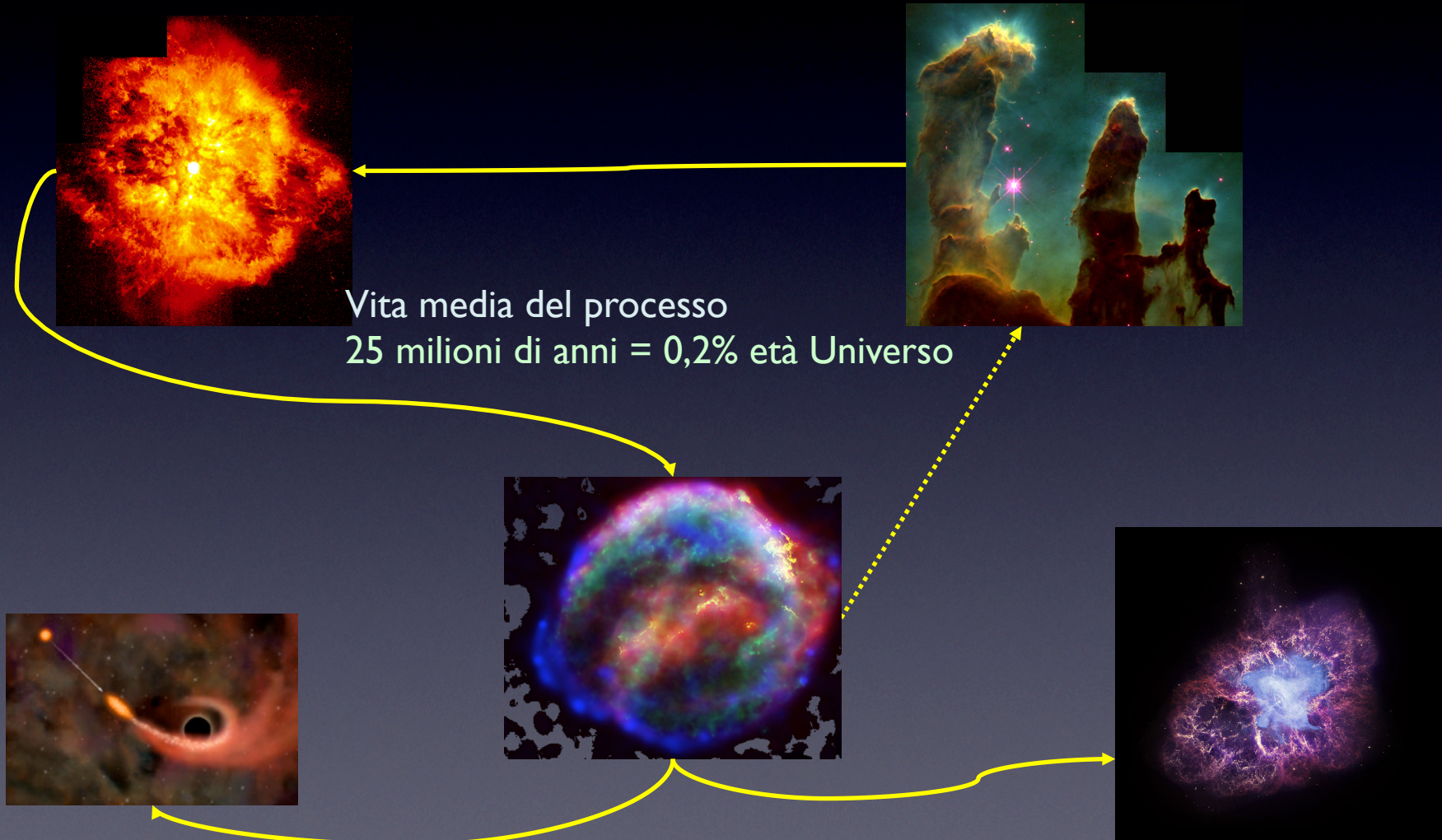


Cassiopea A



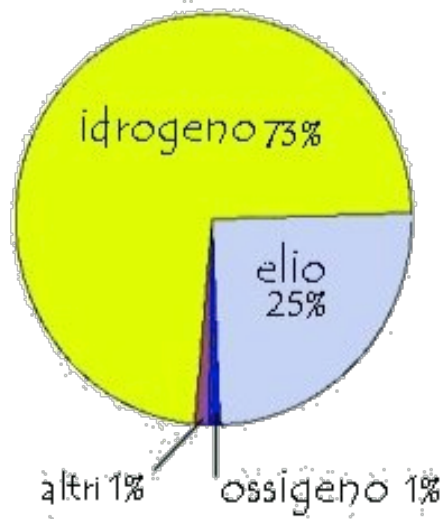


L'evoluzione di una stella di massa pari a 25 soli



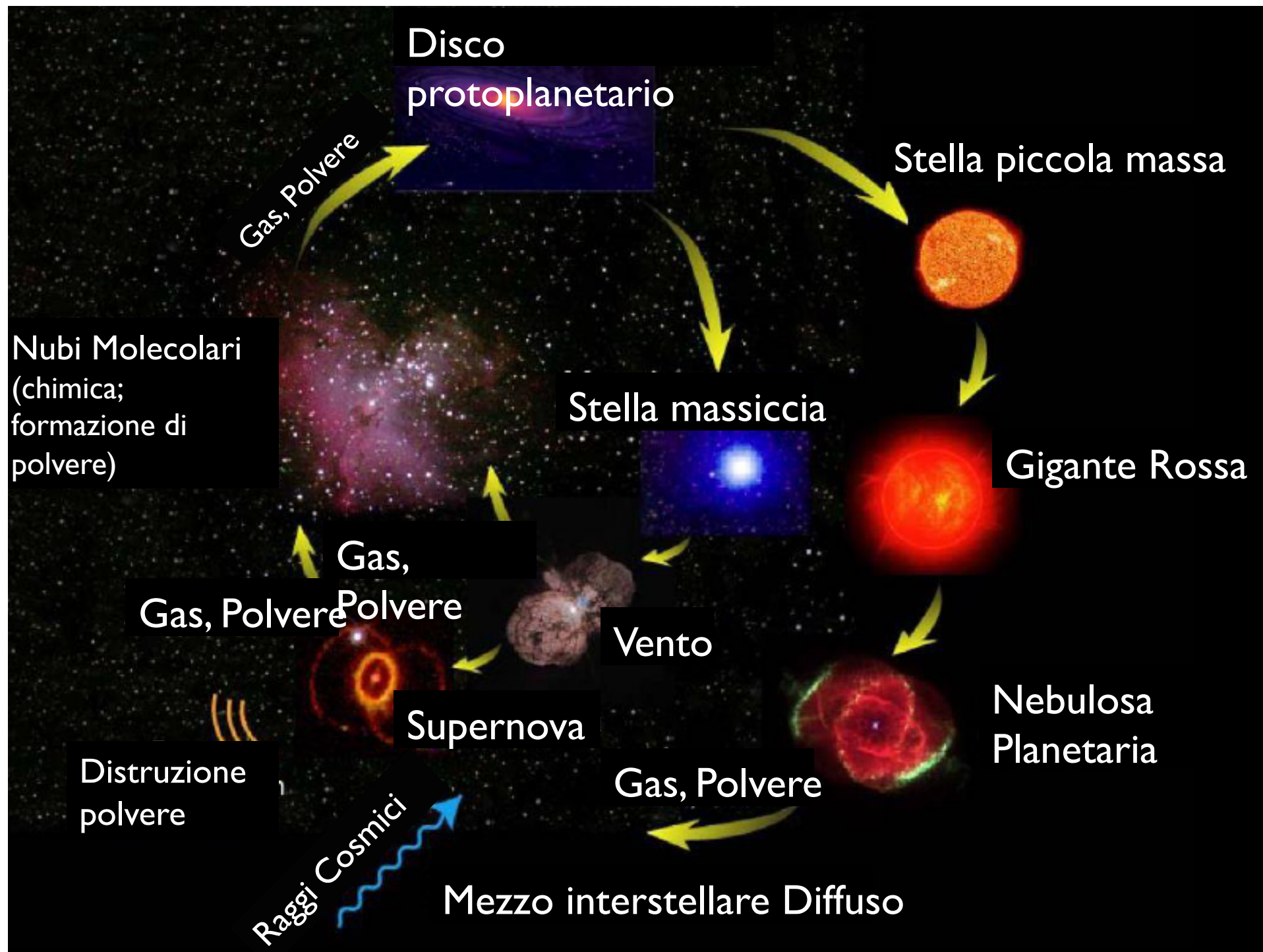


Universo



Acqua ~ 90%





B Big Bang L Large stars s Super-novae C Cosmic rays s Small stars M Man-made																		B C N O F Ne										
H B	Li C		Be C															Al s L						Si s L	P L	S s L	Cl L	Ar L
Na L	Mg L	K L	Ca L	Sc L	Ti s L	V s L	Cr L	Mn L	Fe s L	Co s	Ni s	Cu L	Zn L	Ga s	Ge s	As L	Se s	Br s	Kr s									
Rb s	Sr L	Y L	Zr L	Nb L	Mo s L	Tc L	Ru s L	Rh s	Pd s L	Ag s L	Cd s L	In s L	Sn s L	Sb s	Te s	I s	Xe s											
Cs s	Ba L	Hf s L		Ta s L	W s L	Re s	Os s	Ir s	Pt s	Au s	Hg s L	Tl s L	Pb s	Bi s	Po s	At s	Rn s											
Fr s	Ra s	La L		Ce L	Pr s L	Nd s L	Pm s L	Sm s L	Eu s	Gd s	Tb s	Dy s	Ho s	Er s	Tm s	Yb s L	Lu s											
		Ac s		Th s	Pa s	U s	Np s	Pu s	Am M	Cm M	Bk M	Cf M	Es M	Fm M	Md M	No M	Lr M											

by Cmglee - Own work

Così si determina la
presenza delle
molecole

I “codici a barre”
degli elementi e
delle molecole

Spectra of ethylene oxide and acetaldehyde

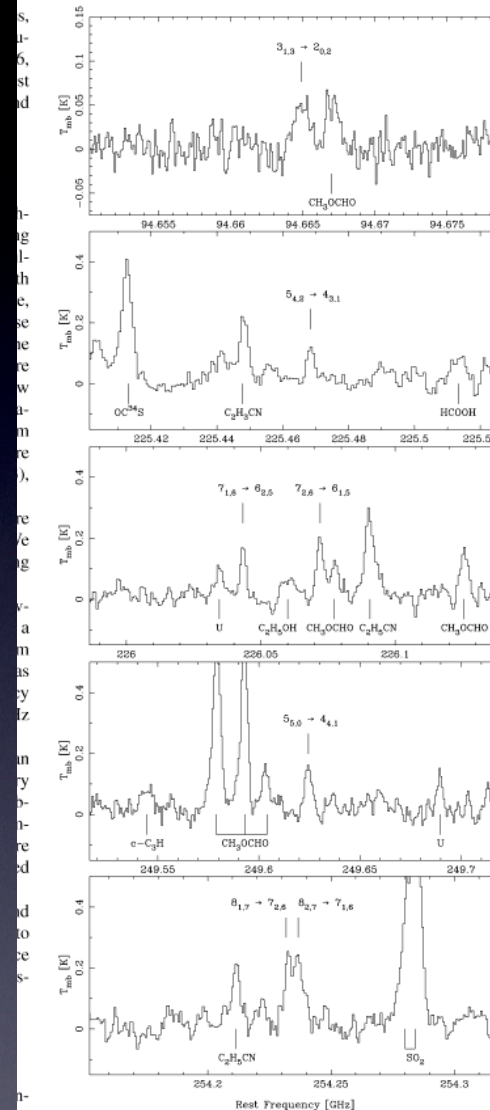


Fig. 1. Spectra towards NGC 6334F, with all seven detected $c\text{-C}_2\text{H}_4\text{O}$ transitions indicated with their quantum numbers. Spectral lines of other molecules are labelled with the relevant species. The rest frequency scale is based on $V_{\text{LSR}} = -7 \text{ km s}^{-1}$.

provided by the NASA Astrophysics Data System

Nummelin et al 1998

Il telescopio ALMA

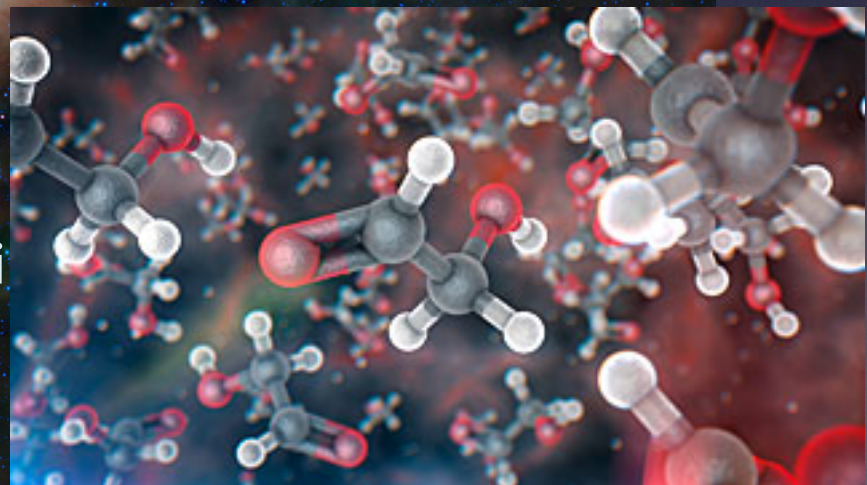


Un “dolce” risultato per ALMA

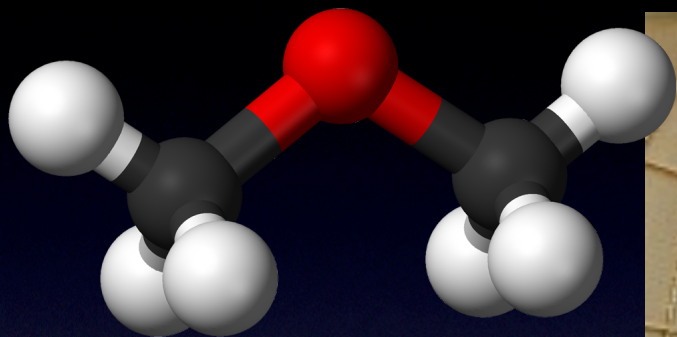


uno degli ingredienti per la formazione dell'RNA

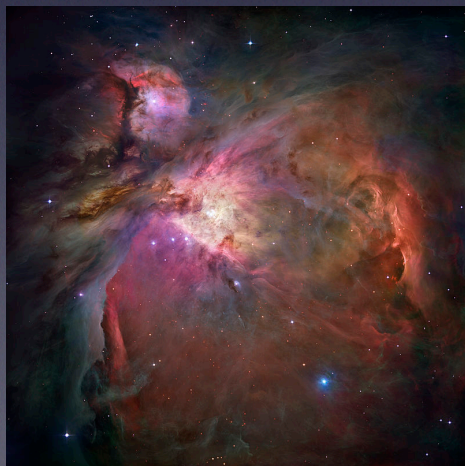
Glicolaldeide in Rho Ophiuchi



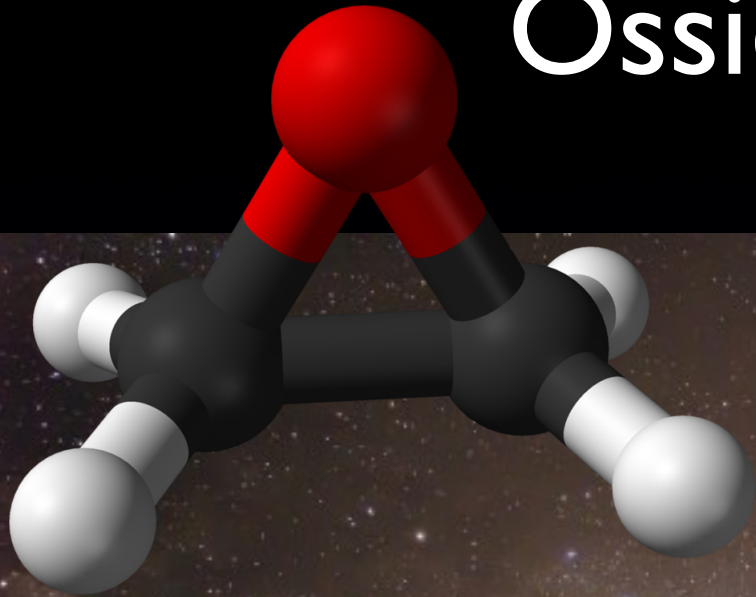
Etere dimetilico



Prima osservazione
Da Kitt Peak (AZ) 1974
nella Nebulosa di Orione



Ossido di etilene SgrB

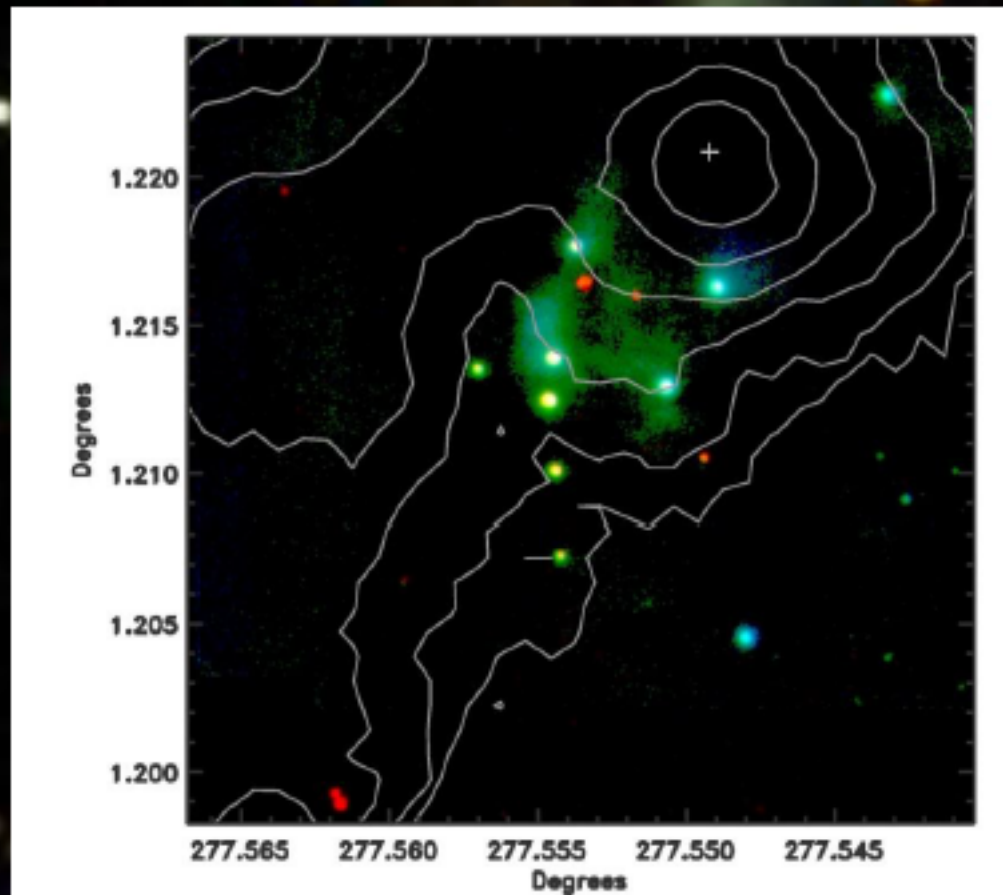


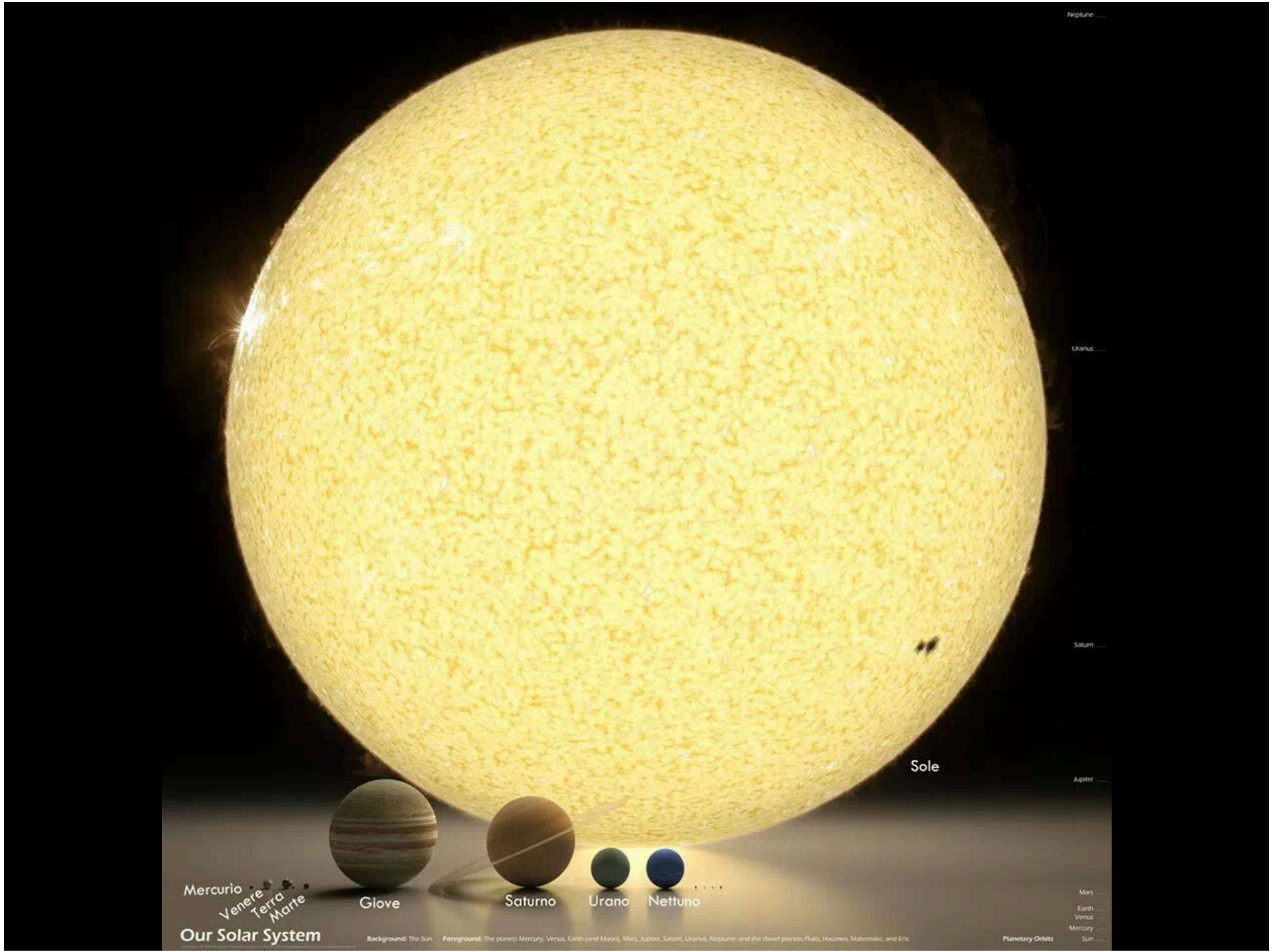
E l'acqua??

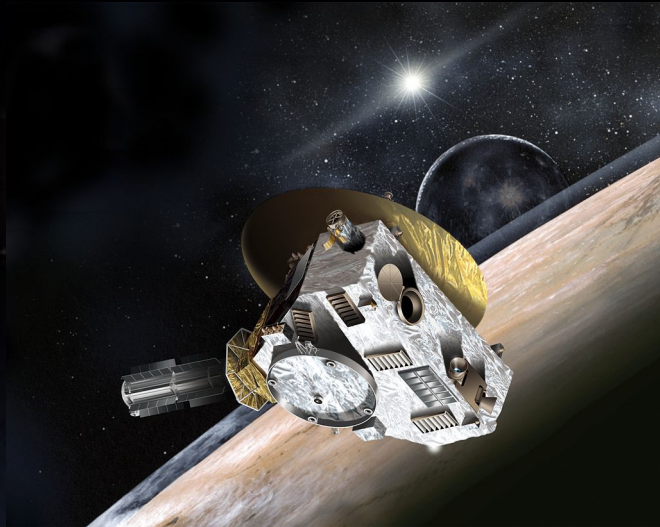
Grandi quantità di ghiaccio – mappa fino a 1000 A

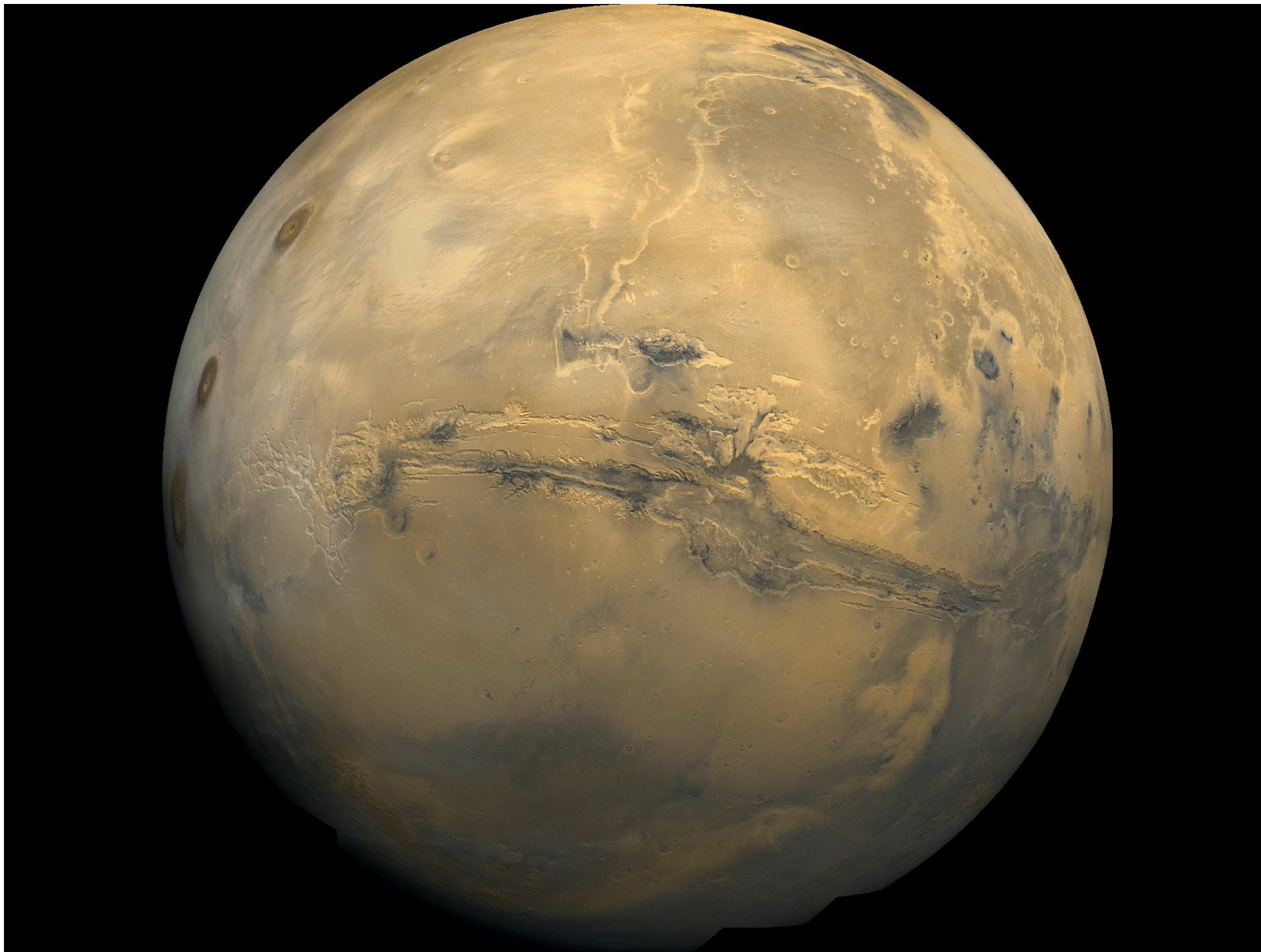
Metà dell'O è congelato – acqua ghiacciata

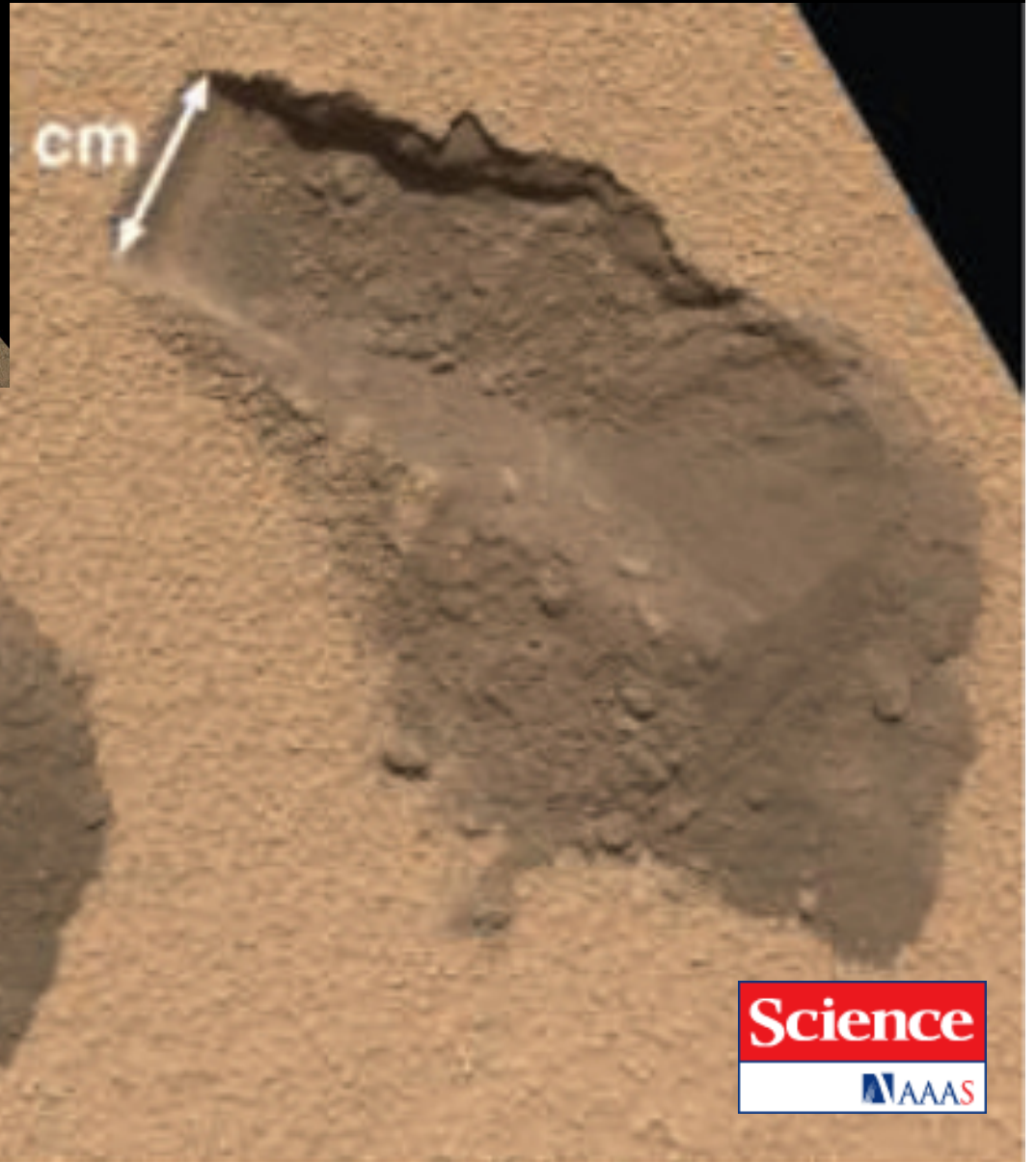
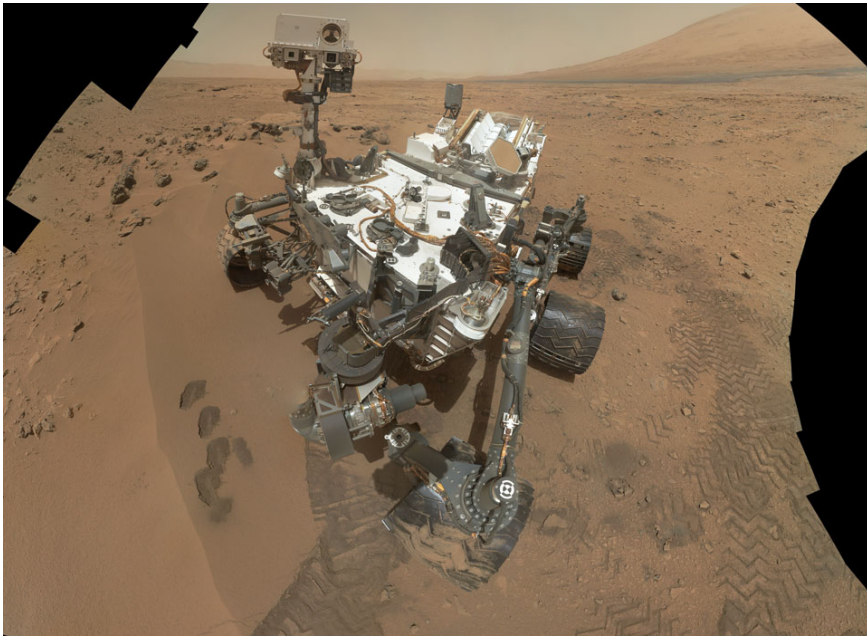
Serpens core





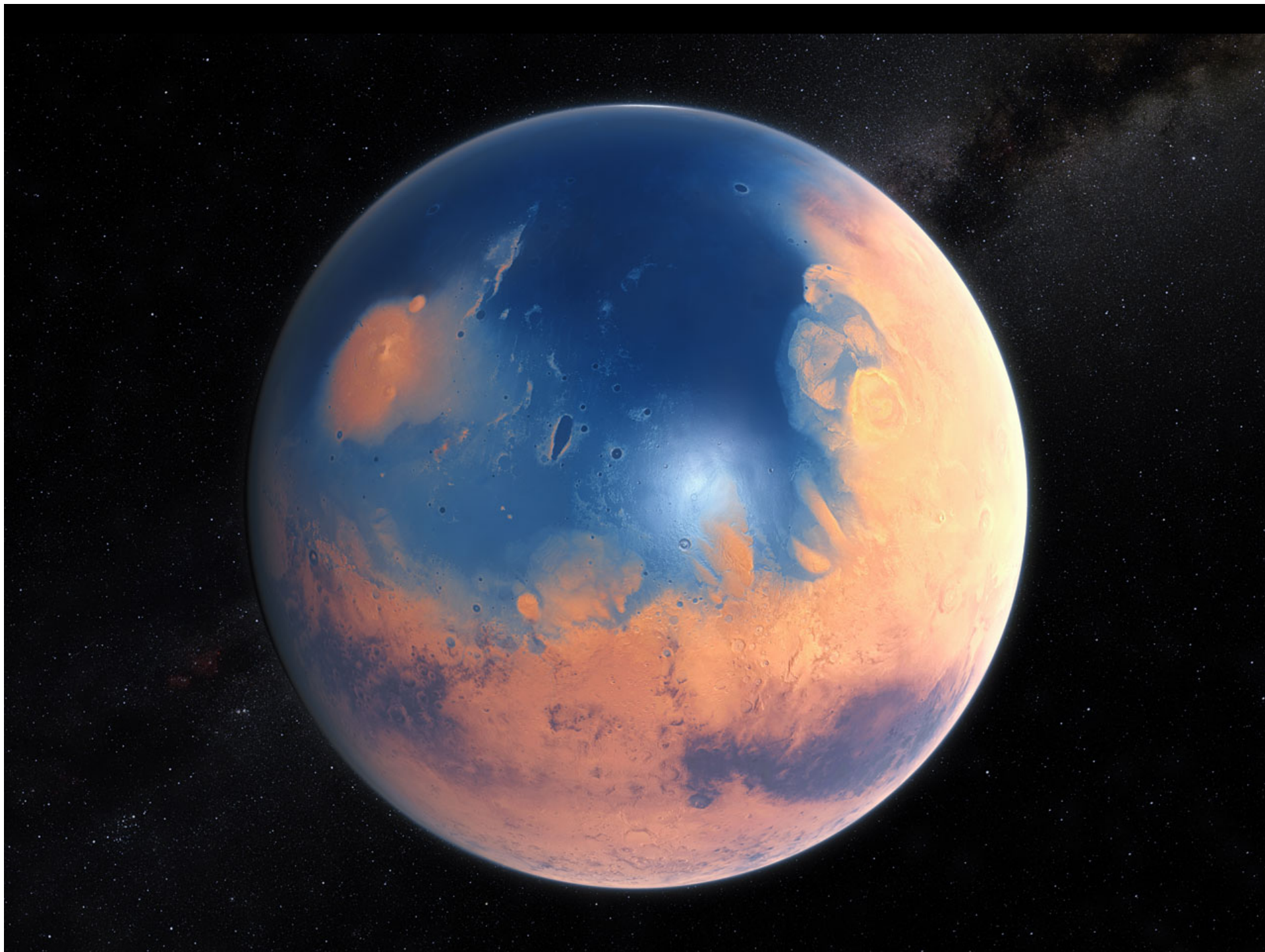






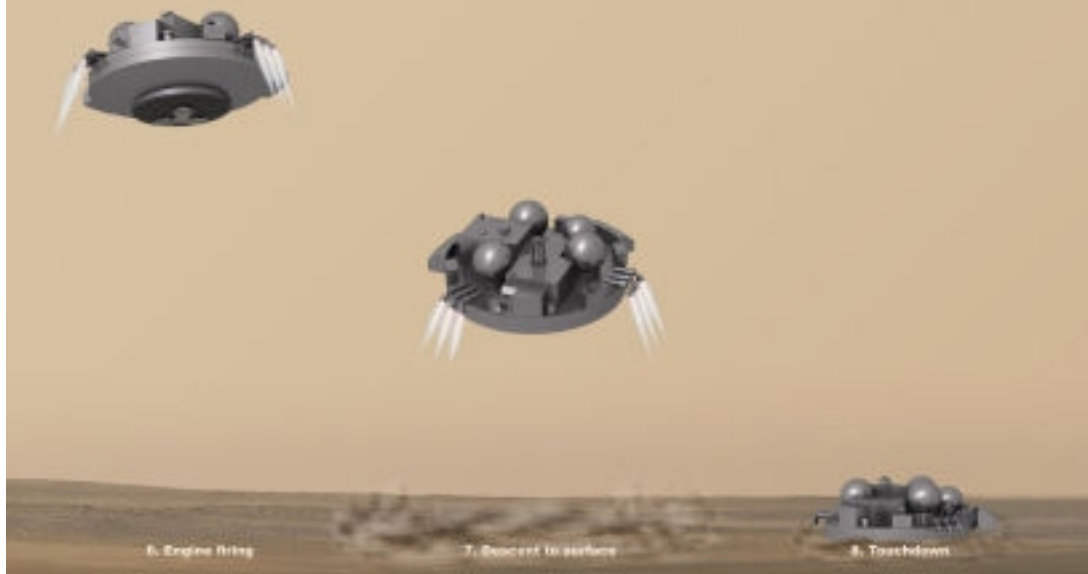
J P Grotzinger Science 2013;341:1475







www.eso.org

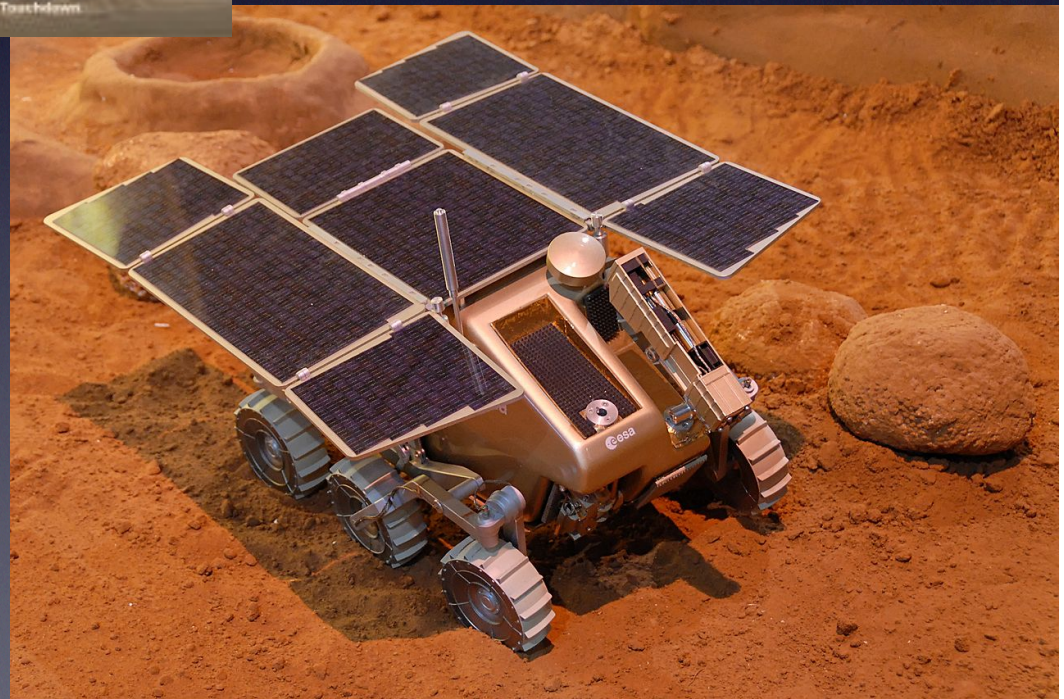


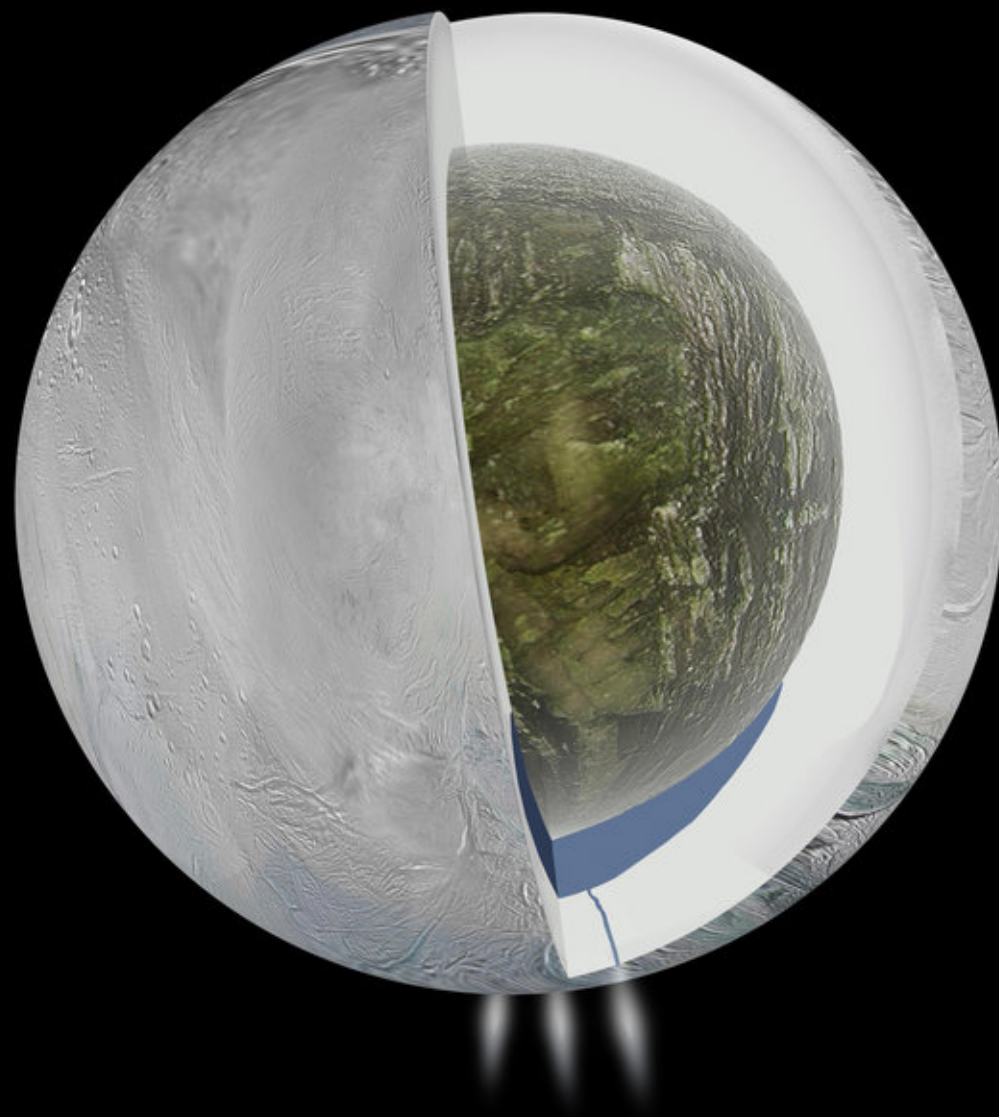
Missione **Aurora** dell'ESA

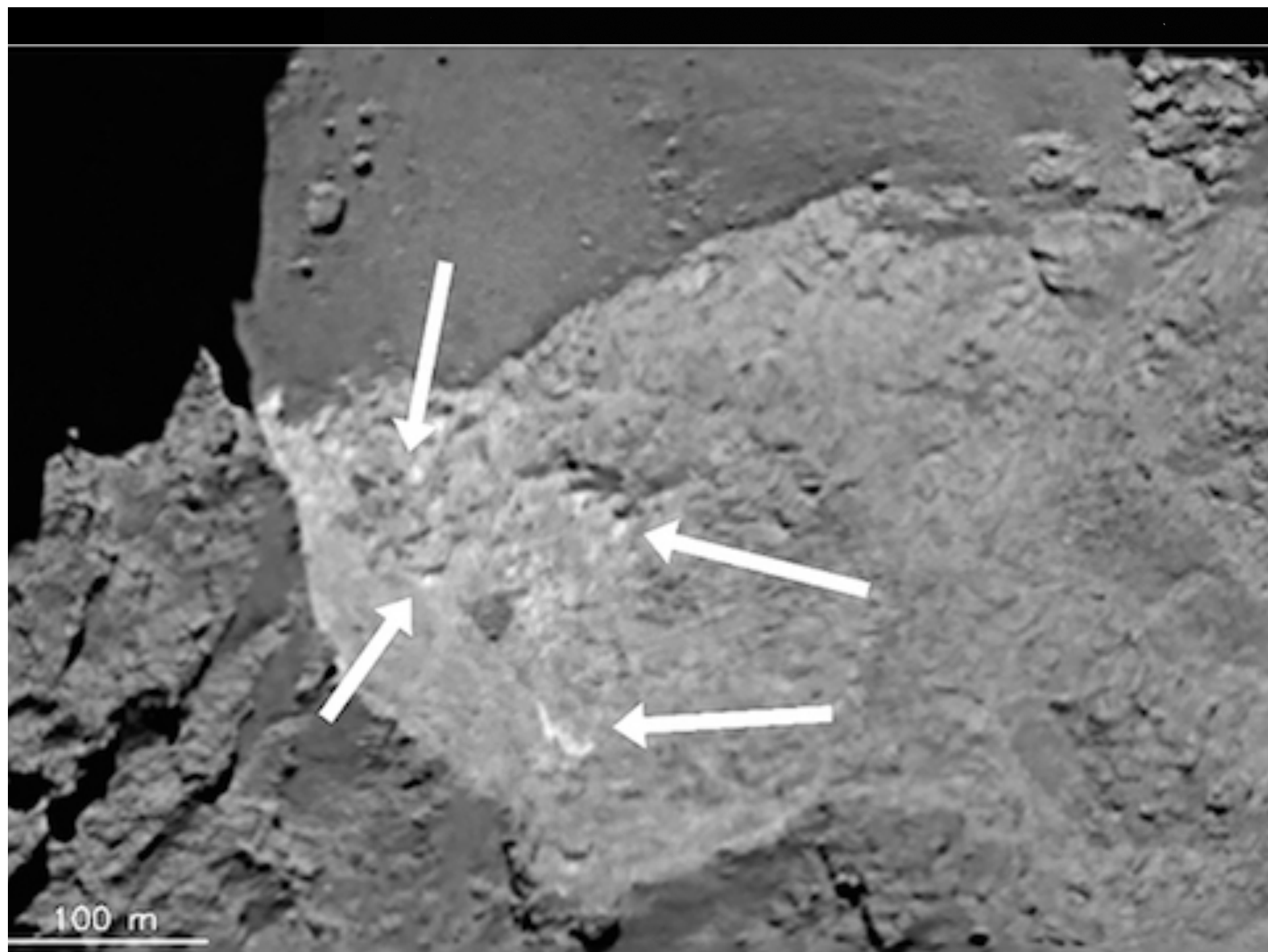
2016 **Schiaparelli:**
ExoMars Entry,
Descent and Landing
Demonstrator Module

Alla ricerca di ematite

2018: **lander per esperimenti**
di esobiologia e geochimica









Giordano Bruno

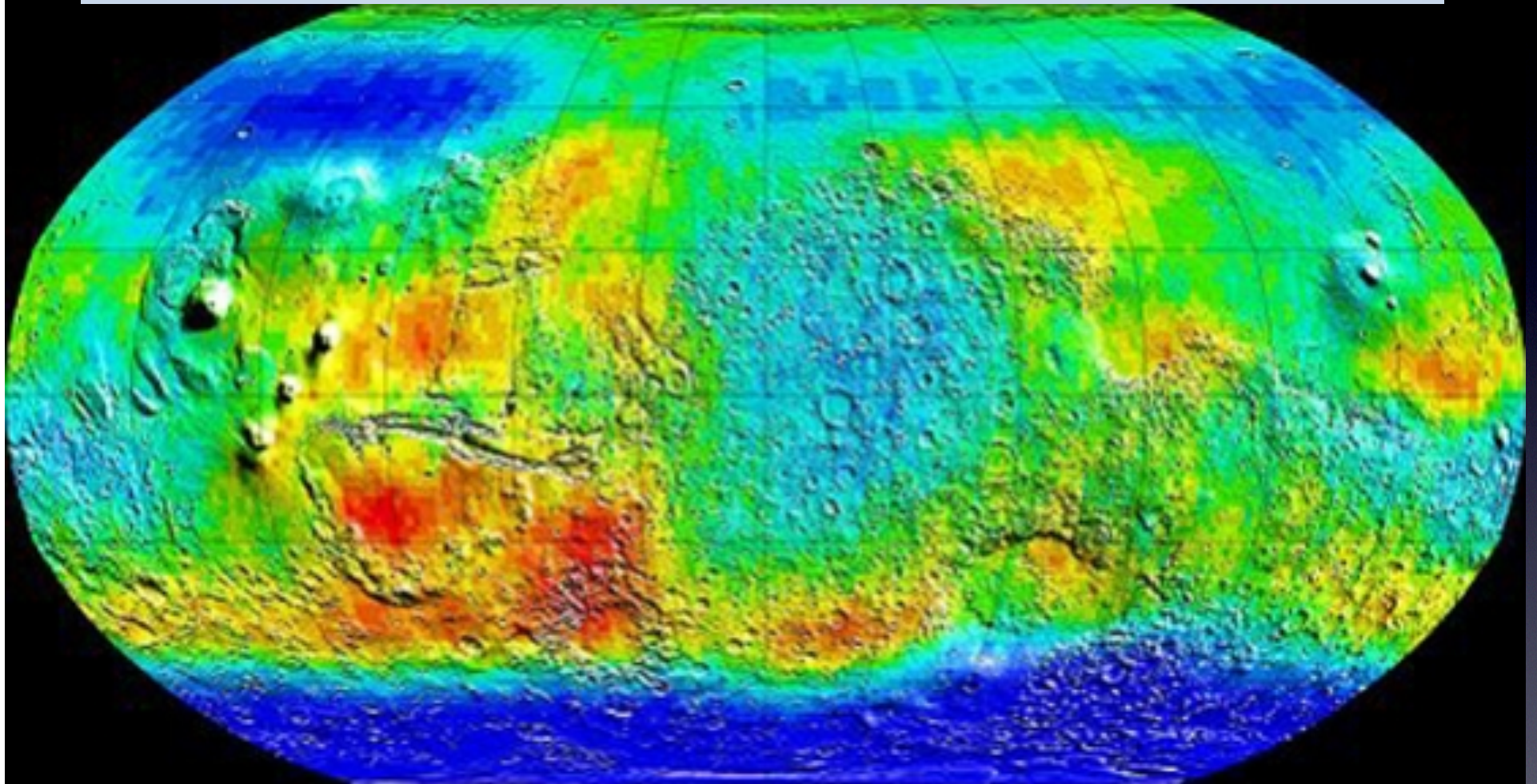
.... Sono dunque **soli innumerabili, sono terre infinite**, che similmente circuiscono quei soli; come veggiamo questi sette circuire questo sole a noi vicino.

De l'infinito, universo e mondi – 3° dialogo (1584)

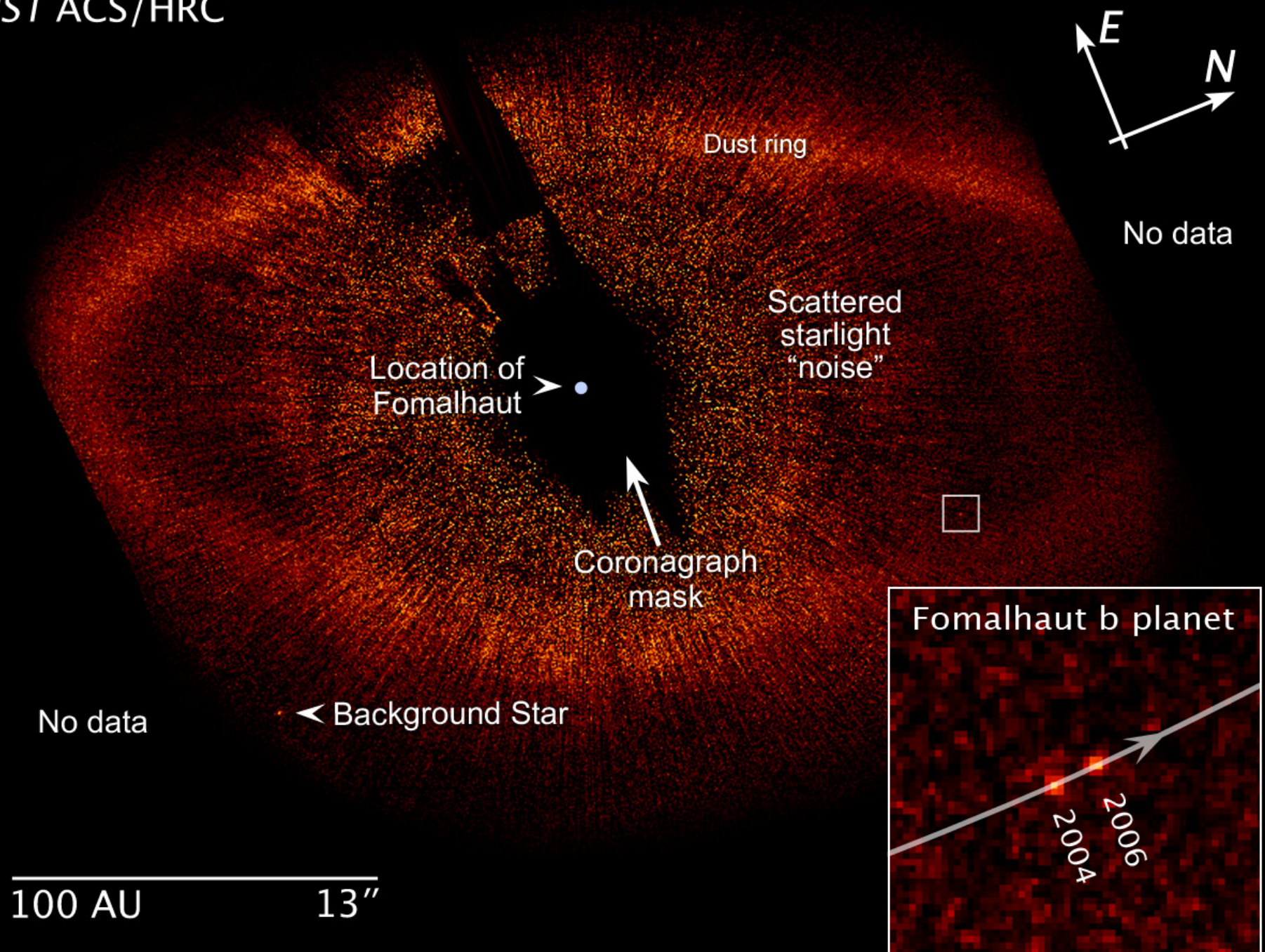
New Mathematical Hall **Praga.**
Clementinum - affresco



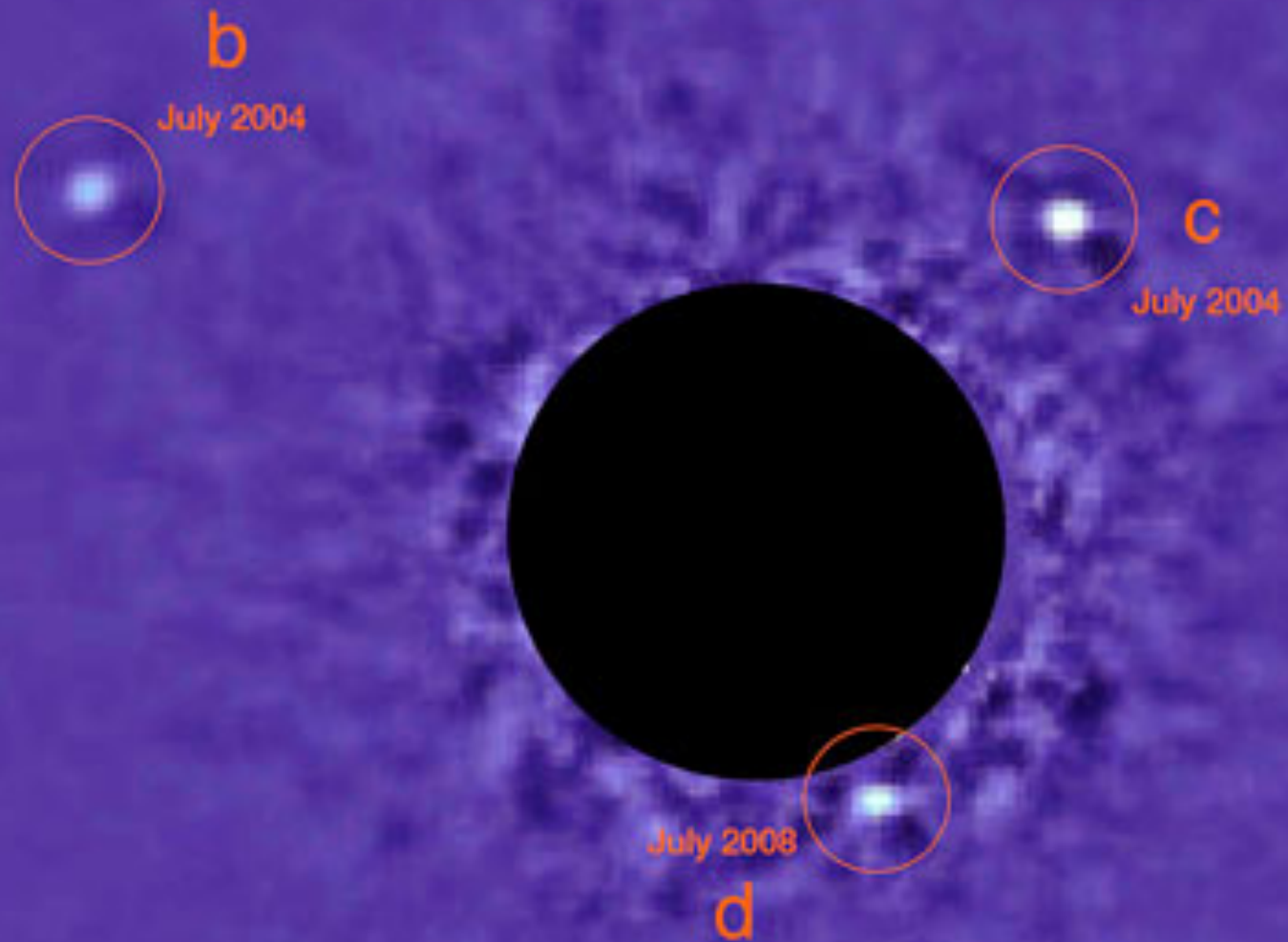
Photo: I. Král, National Library, Prague



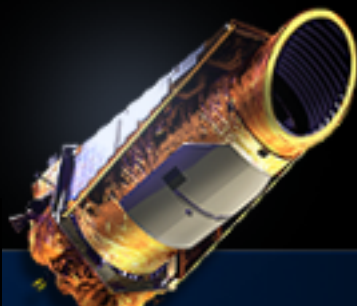
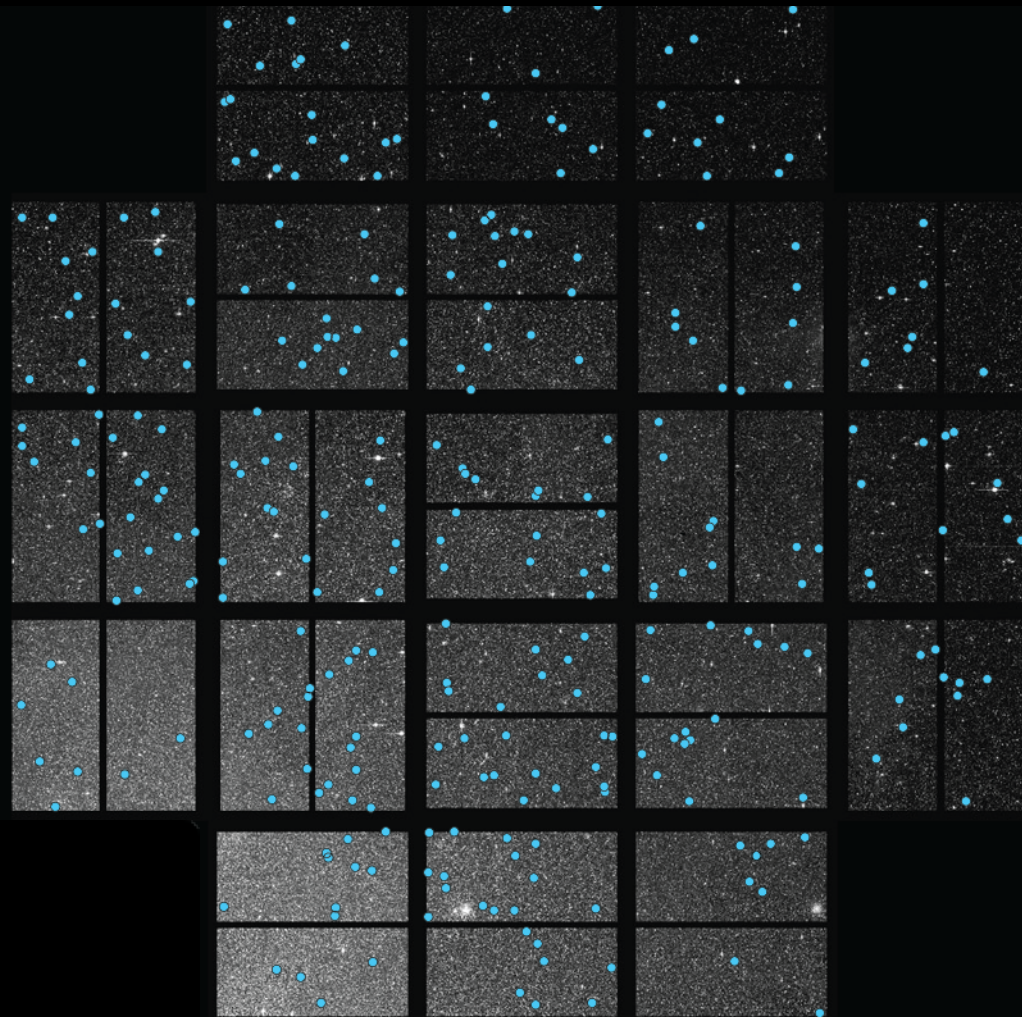
Fomalhaut
HST ACS/HRC



HR 8799

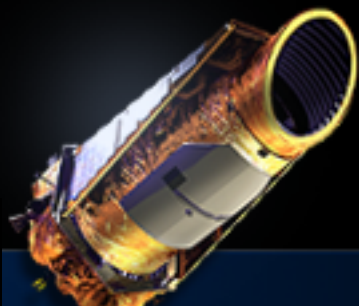
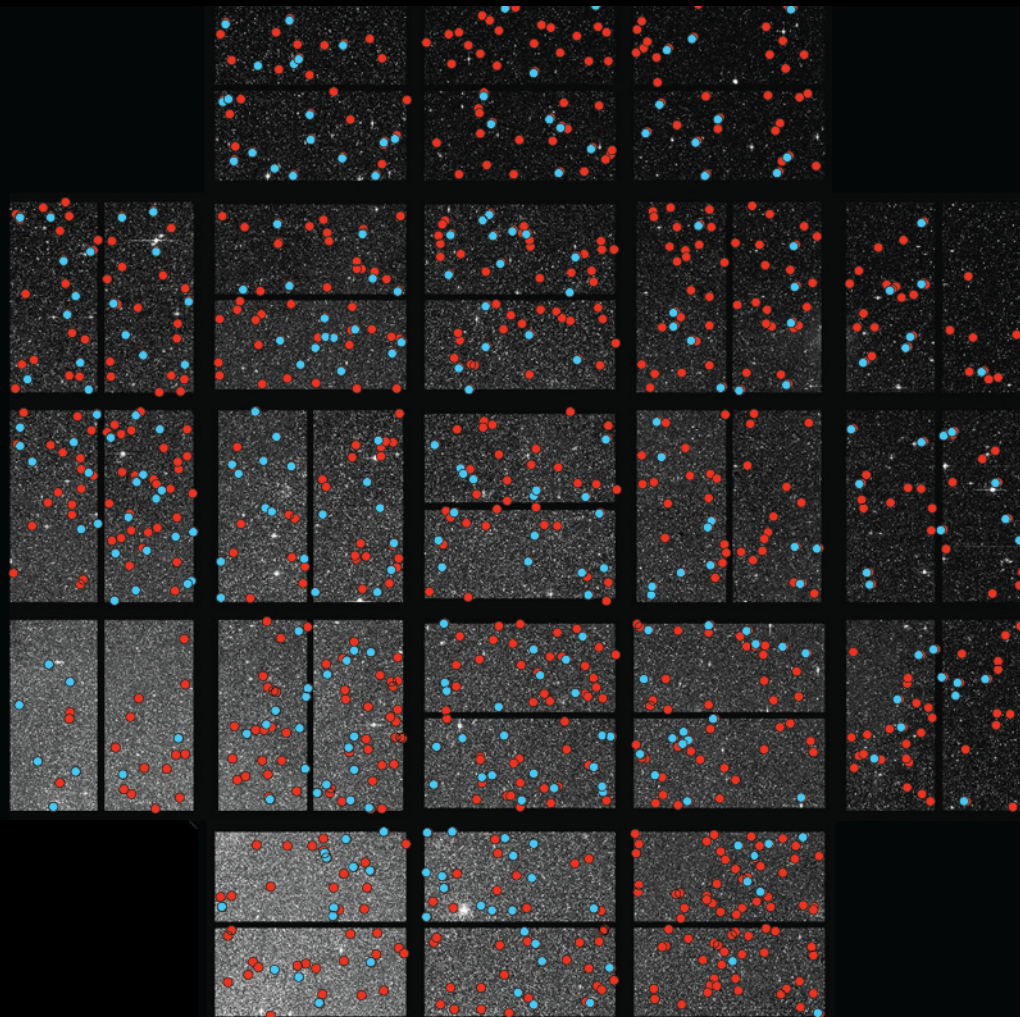


● June 2010
Catalog Release

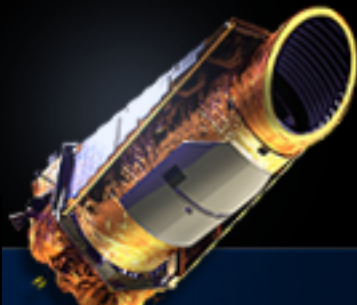
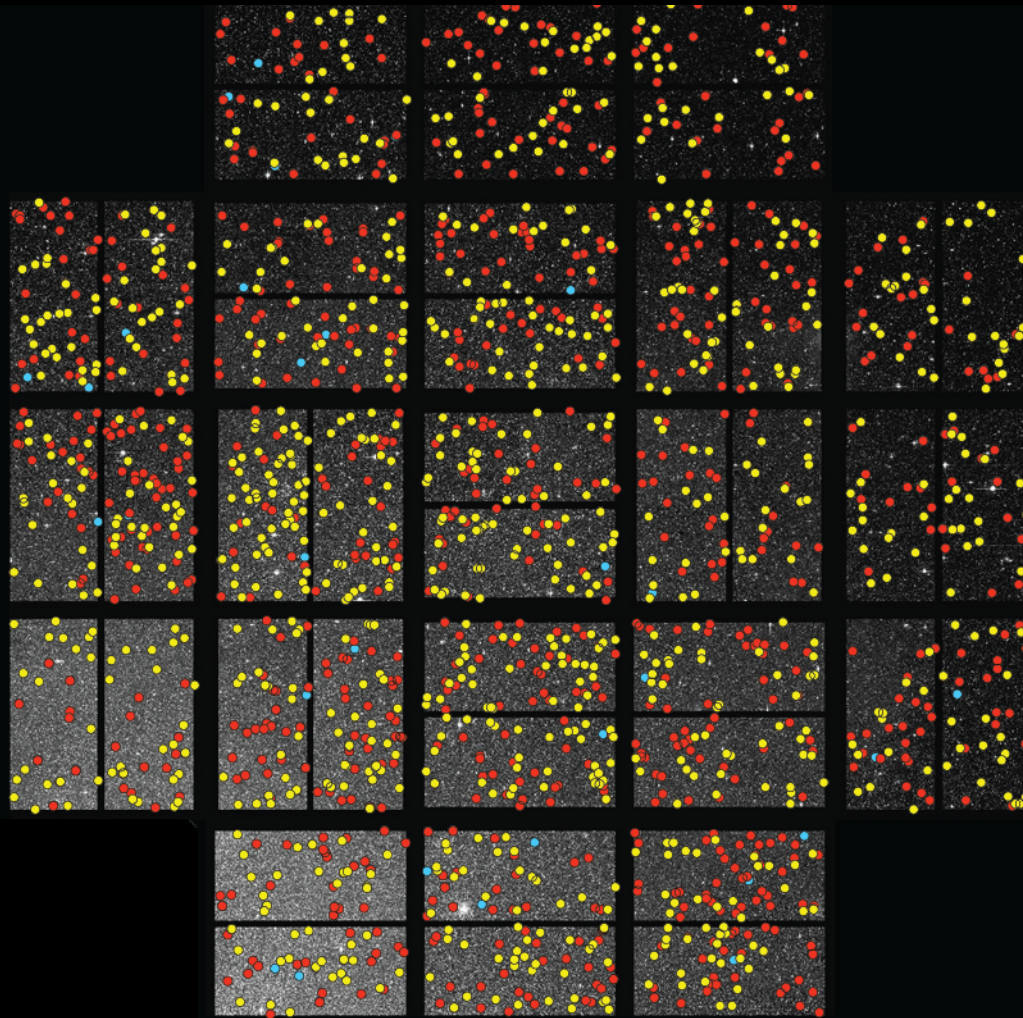


● June 2010
Catalog Release

● February 2011
Catalog Release



- June 2010
Catalog Release
- February 2011
Catalog Release
- February 2012
Catalog Release



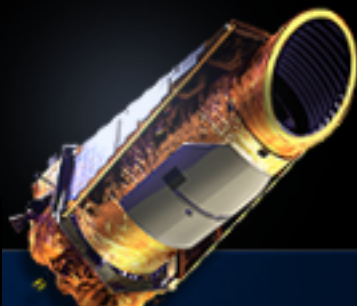
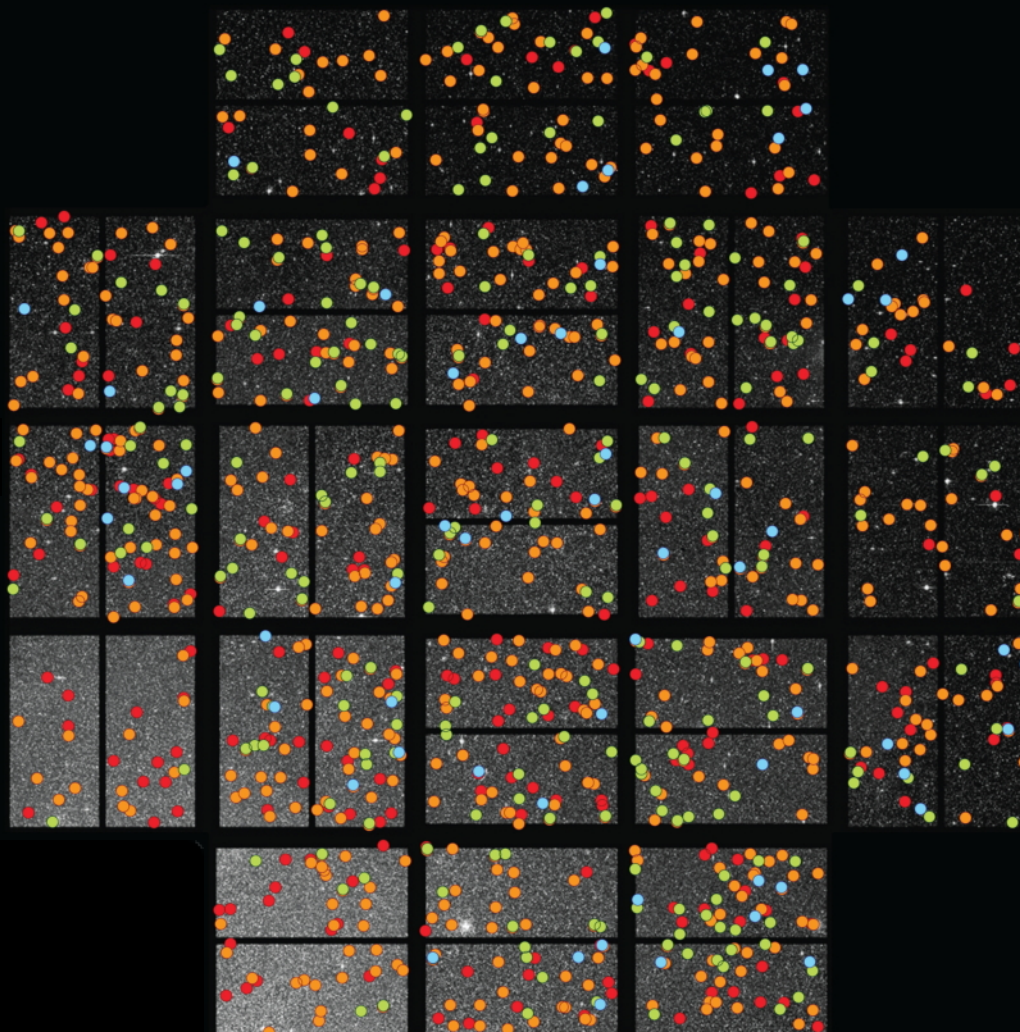
Candidati Pianeti

● Terrestri

● SuperTerre

● ~ Nettuno

● Giganti

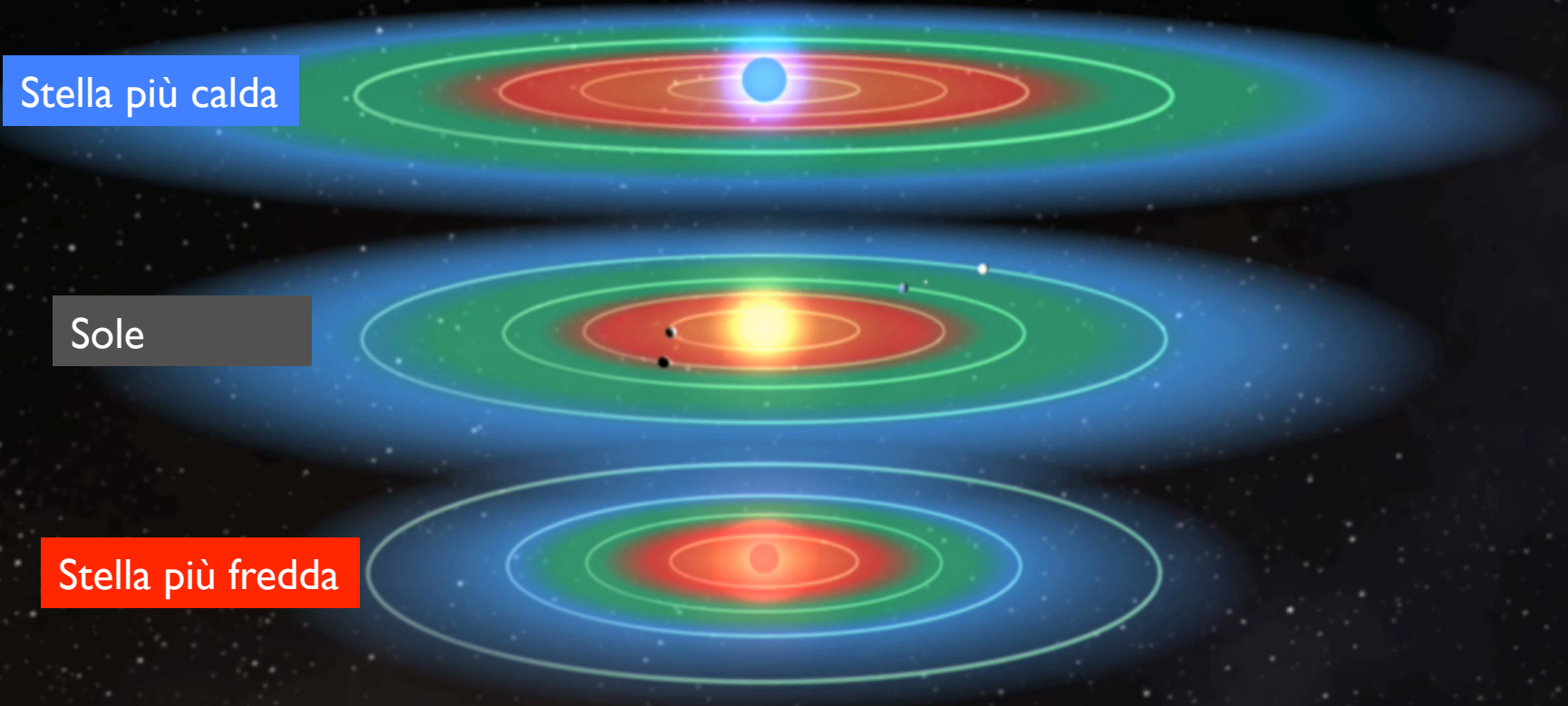


Zona abitabile

Stella più calda

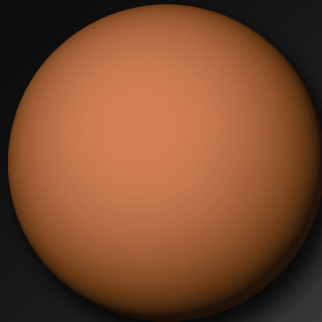
Sole

Stella più fredda



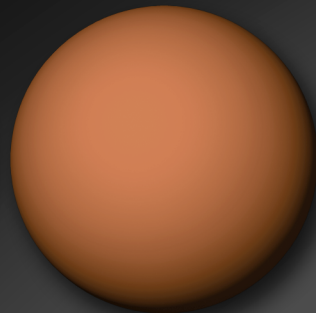
Dimensione dei pianeti

Kepler-7b



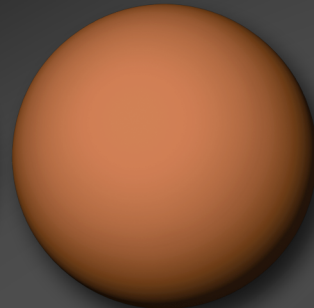
16.52 R_E

Kepler-5b



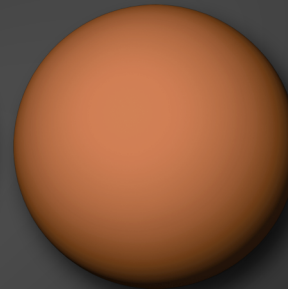
16.00 R_E

Kepler-8b



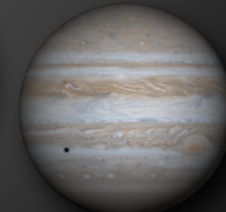
15.86 R_E

Kepler-6b



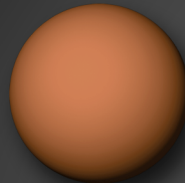
14.79 R_E

Jupiter



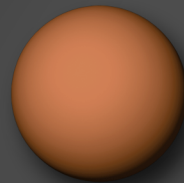
11.2 R_E

Kepler-9b



9.4 R_E

Kepler-9c



9.2 R_E

Kepler-4b



3.99 R_E

Kepler-9d



1.64 R_E

Kepler-10b



1.42 R_E

Earth



Kepler-11b



1.97 R_E

Kepler-11c



3.15 R_E

Kepler-11d



3.43 R_E

Kepler-11e



4.52 R_E

Kepler-11f



2.61 R_E

Kepler-11g



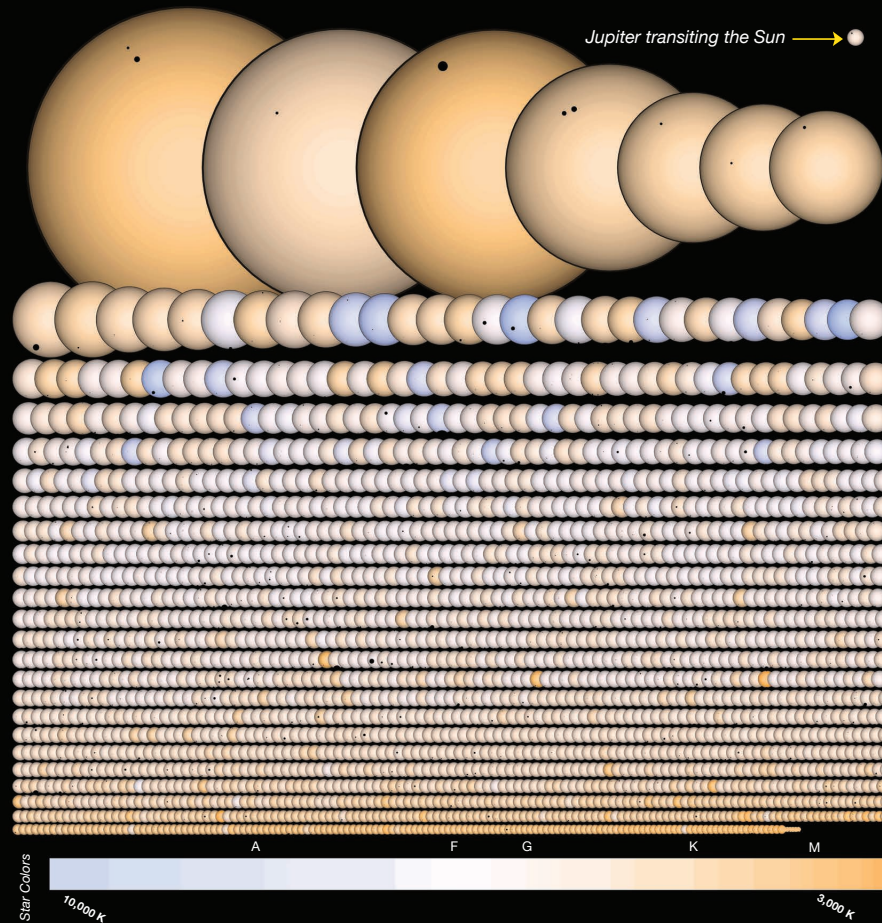
3.66 R_E

National Aeronautics and Space Administration



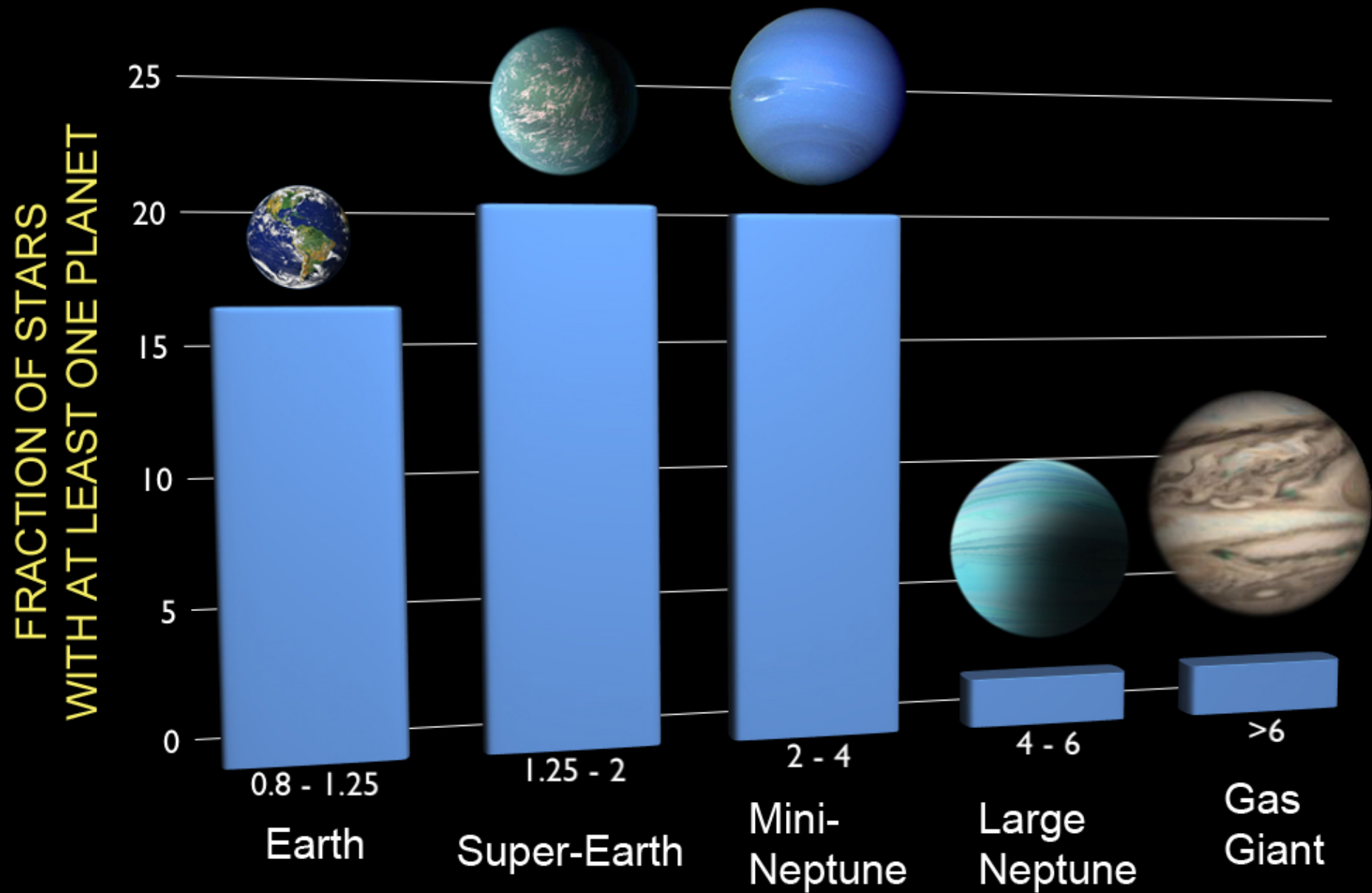
KEPLER'S PLANET CANDIDATES

2,740 AS OF JANUARY, 2013



Using NASA's planet-hunting Kepler spacecraft, astronomers have discovered 2,740 planet candidates orbiting 2,036 other suns in a search for Earth-size worlds. The search began in 2009. Kepler monitors a rich star field for planetary transits, which cause a slight dimming of starlight when a planet crosses the face of its star. In "Kepler's Planet Candidates," the systems are ordered by star diameter. The star's color represents its temperature as shown in the lower scale, and the letters (A, F, G, K, M) designate star types. The simulated stellar disks and the planet silhouettes are shown at the same scale, with saturated star colors. Look carefully: some systems have multiple planets. For reference, Jupiter is shown transiting the Sun. Higher resolutions of this graphic are available at <http://Kepler.NASA.gov/images/graphics>

www.nasa.gov



Fressin et al 2013

PLANET SIZE (relative to Earth)

Jupiter

Neptune

Super Earth

Radius

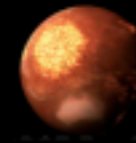
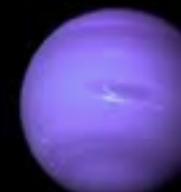
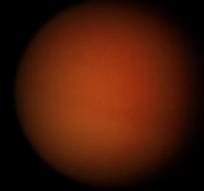
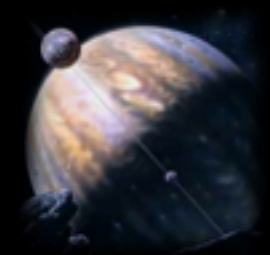
'Hot'
(850-1500K)

'Warm'
(500-800K)

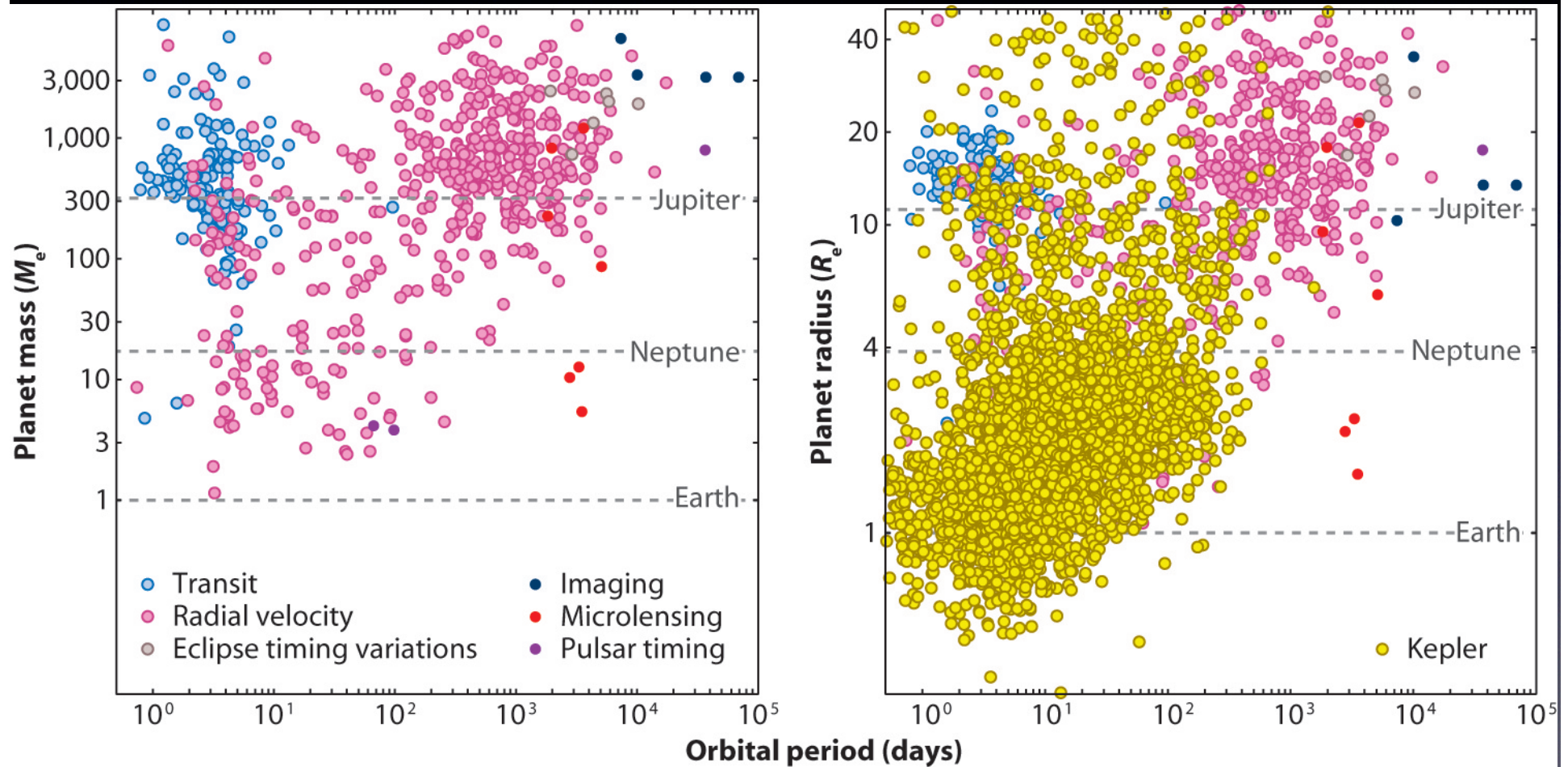
'Temperate'
(250-350K)

Temperature

Difficulty

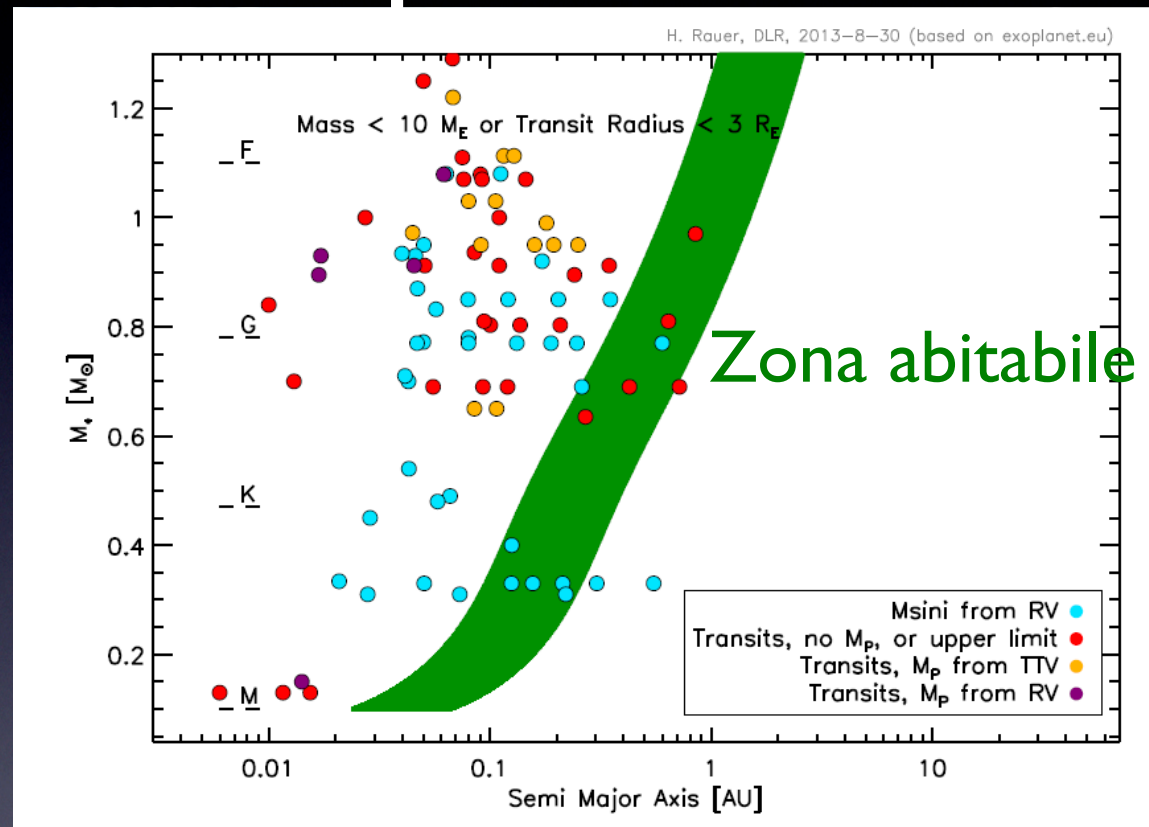


Dimensioni



Super Terre

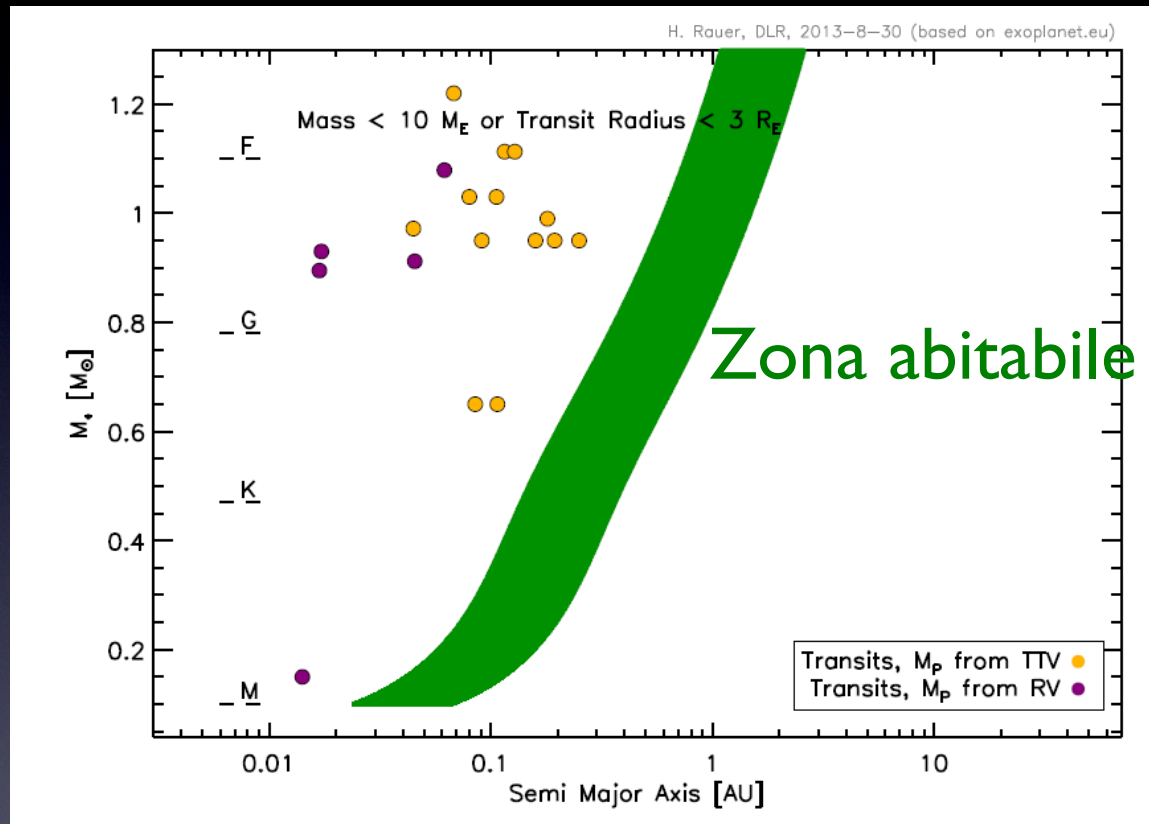
Massa



Semi asse maggiore dell'orbita

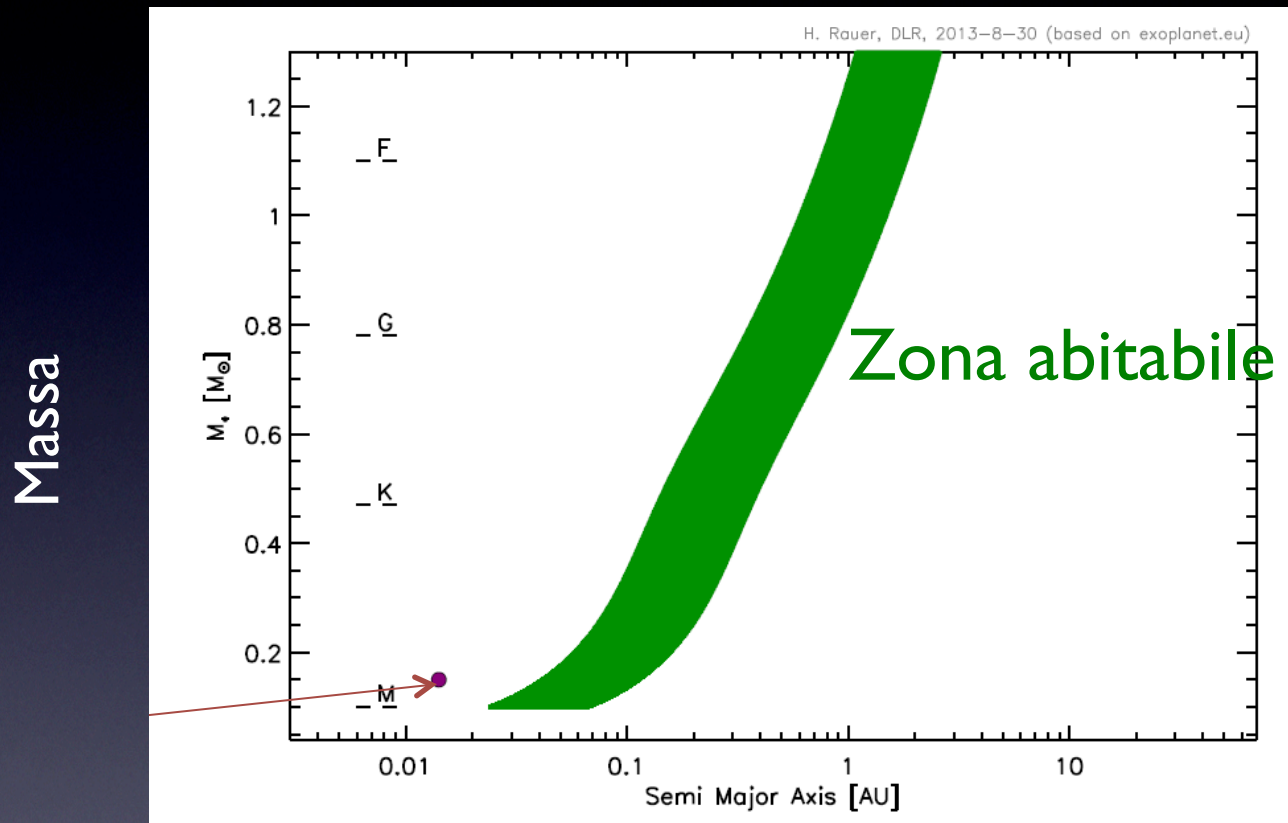
Massa e raggio noti

Massa

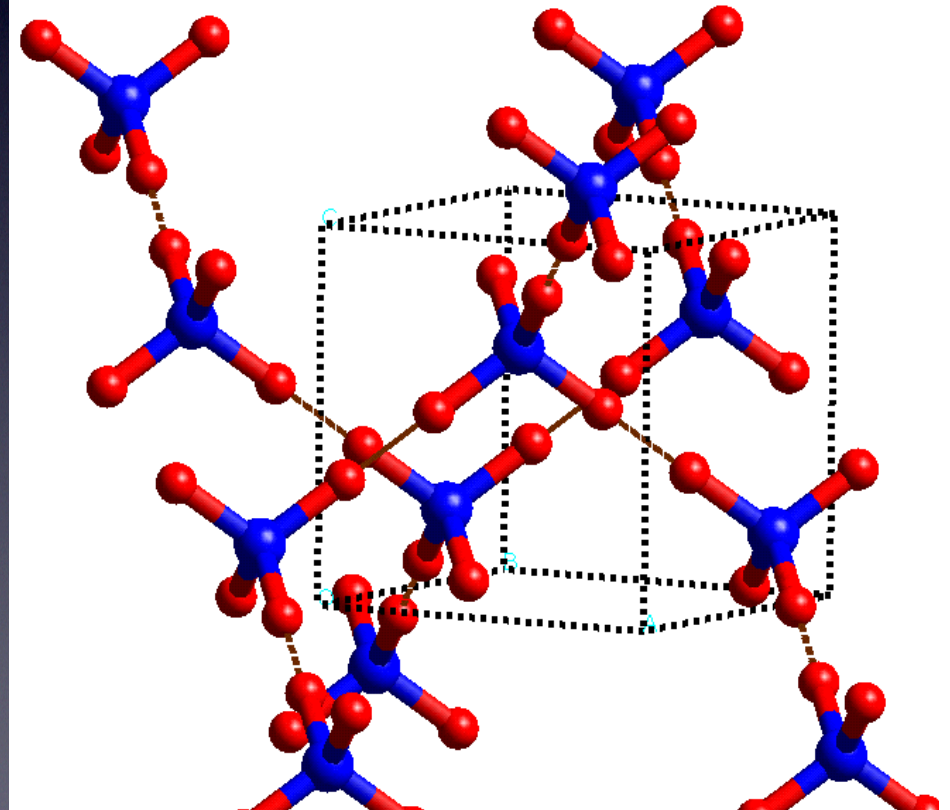
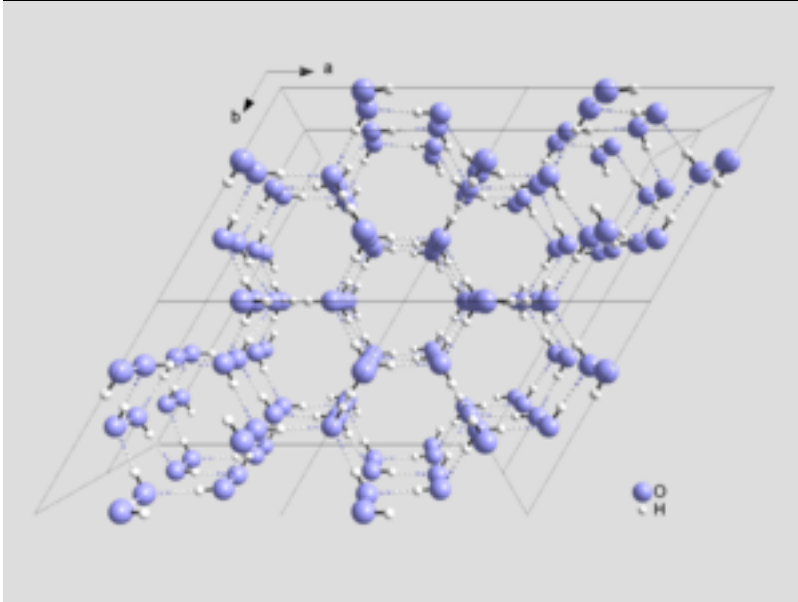


Semi asse maggiore dell'orbita

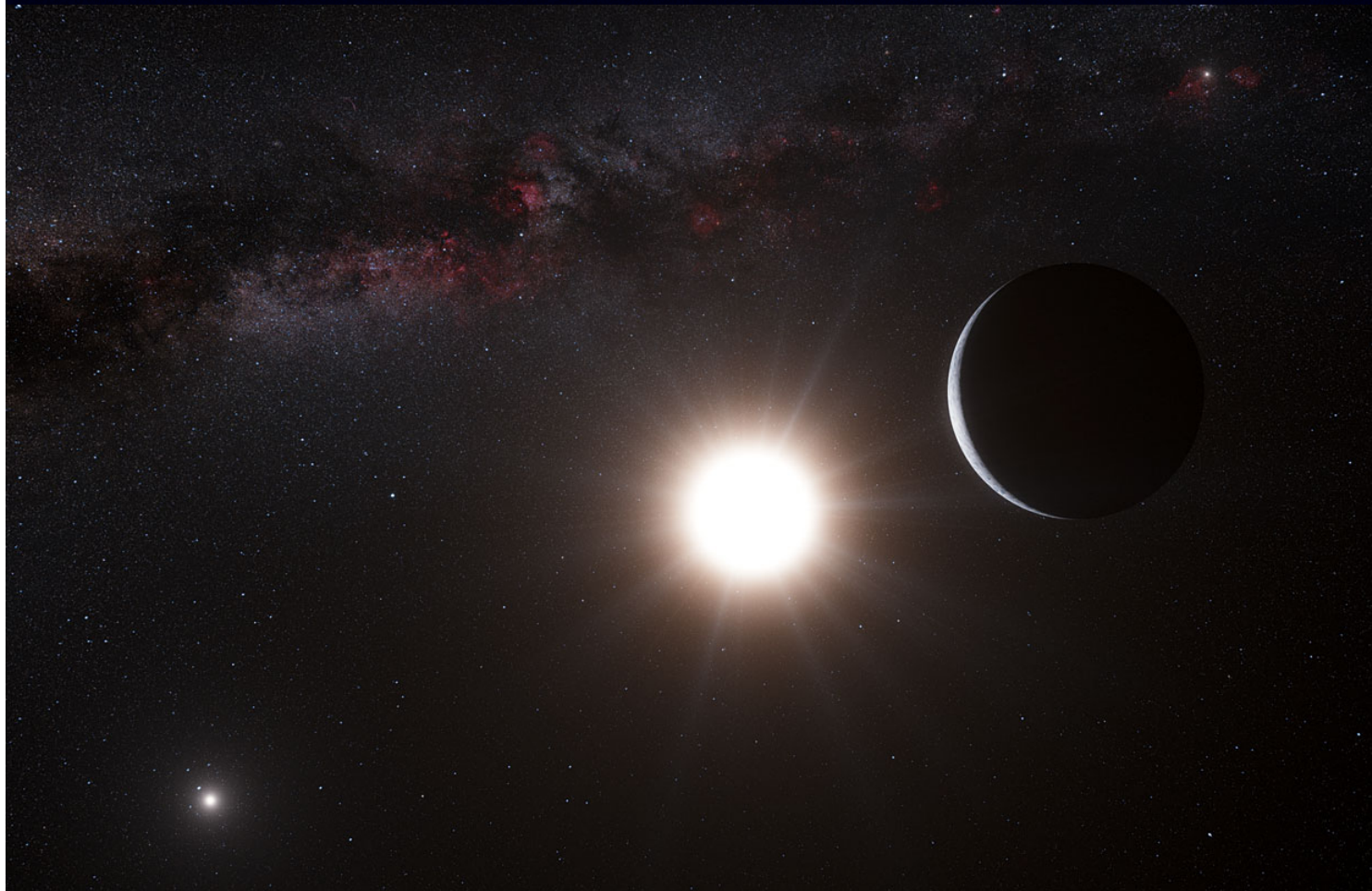
Brillanti per vedere atmosfera



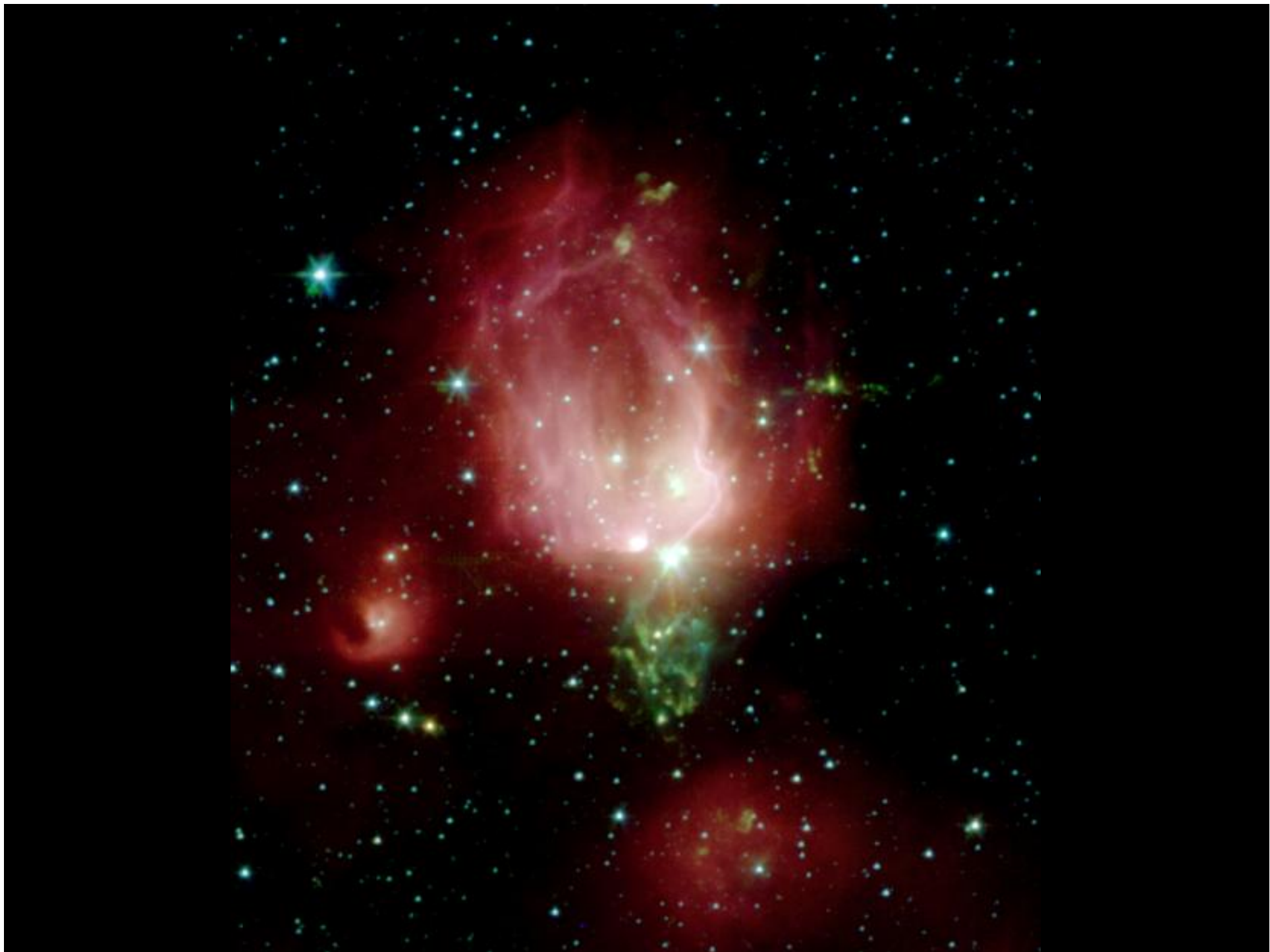
Ciclo del Carbonio



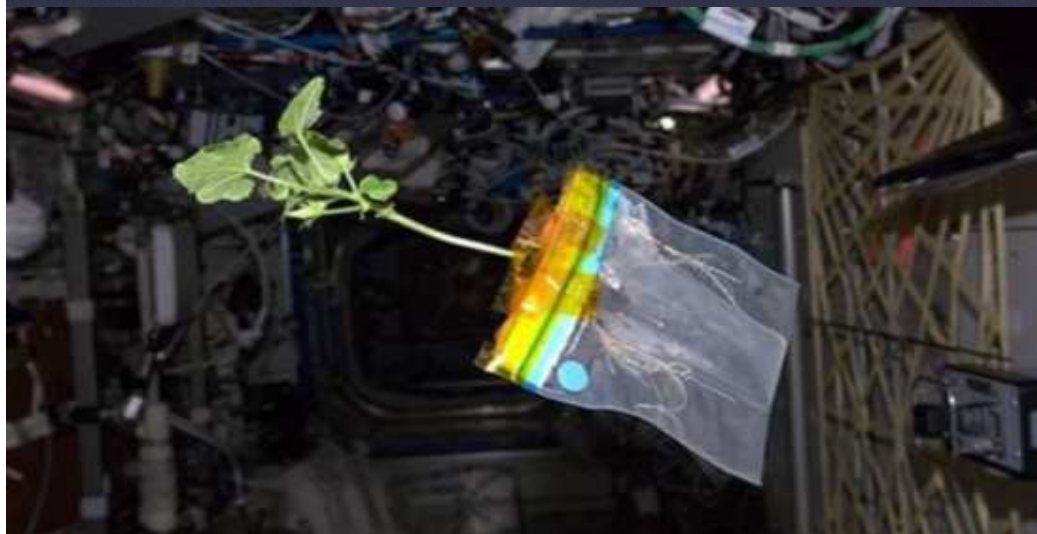
Un nuovo vino?



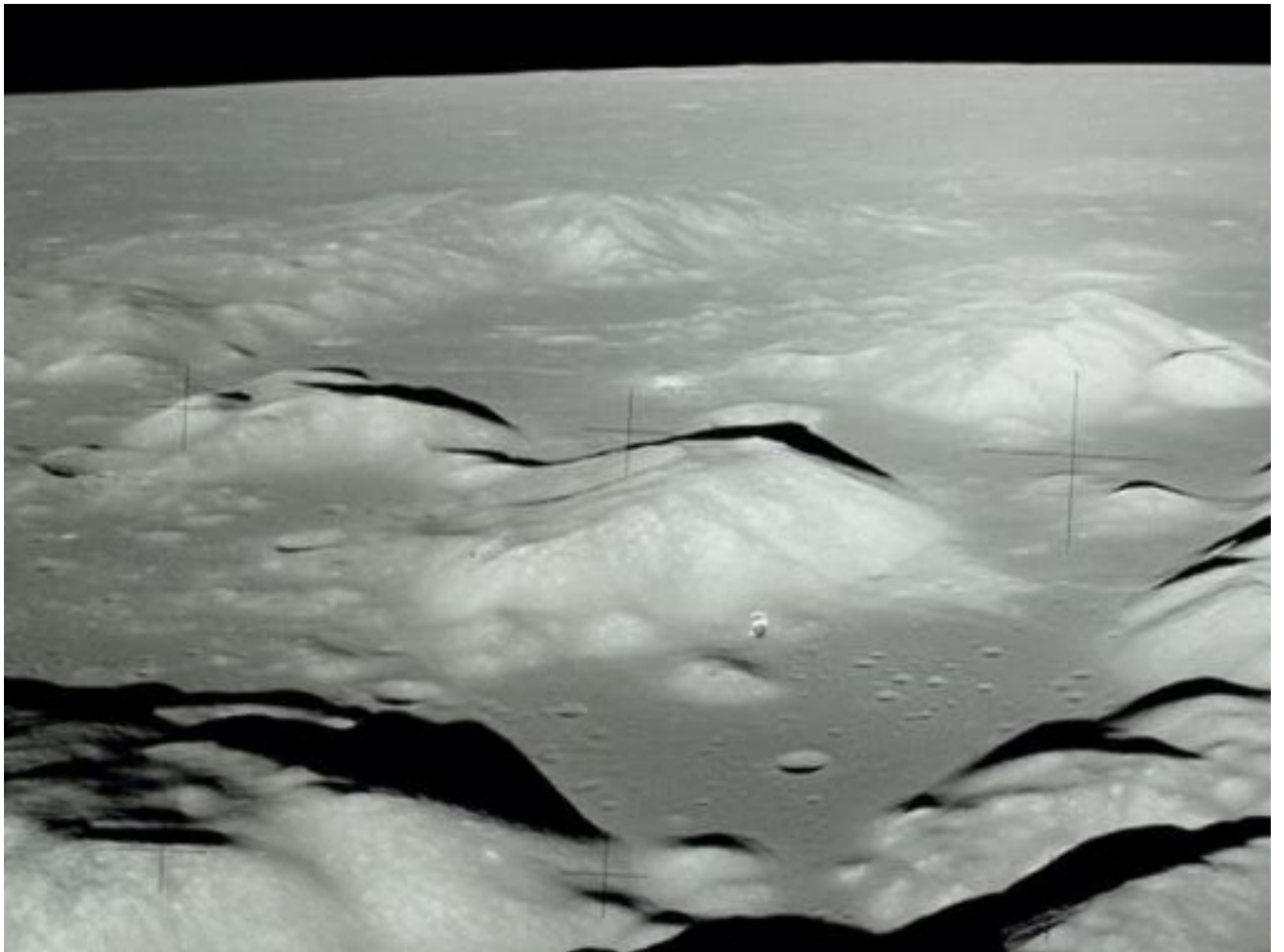




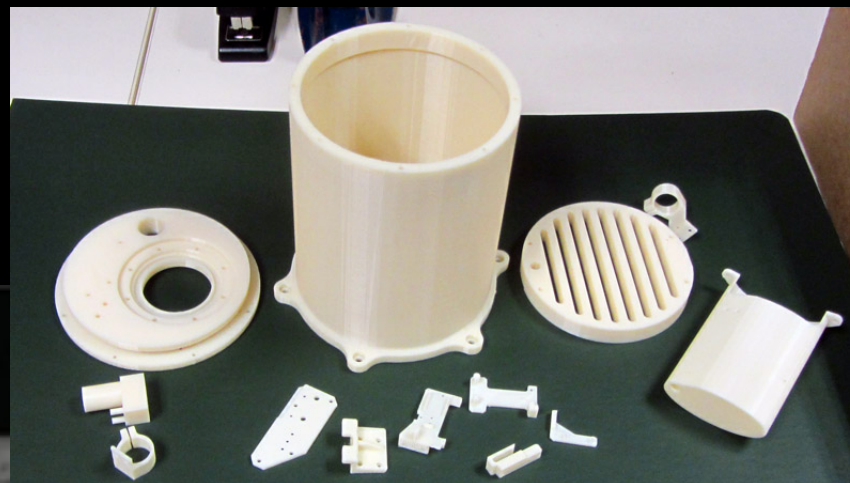




Piante sulla ISS
crescione basilico ..



Ascensio
Machinae
Ad
Lunam
Italica
Arte



Google Lunar X-Prize
Per il 2016



